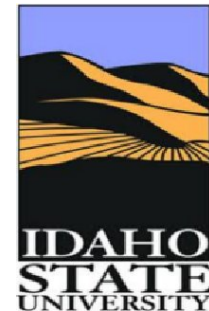
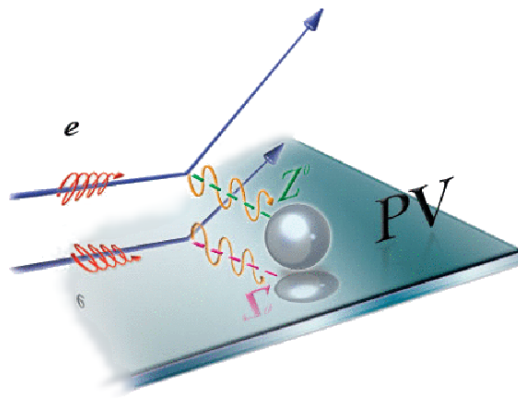


# Applications of Parity Violation

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Daniel Sluder, Brady Lowe, Joey McCullough

January 29, 2018





# Applications of Parity Violation in the Weak Nuclear Force

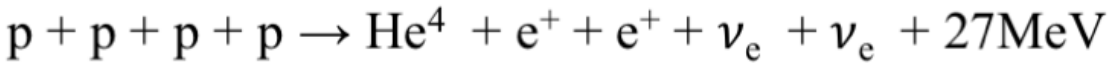
## Outline

- Beta Decay and Parity Violation
- Standard Model and the Weak Force
- Experiments: PREX/CREX and MOLLER
- Quartz Cerenkov Detector R&D at ISU
- Summary and Outlook



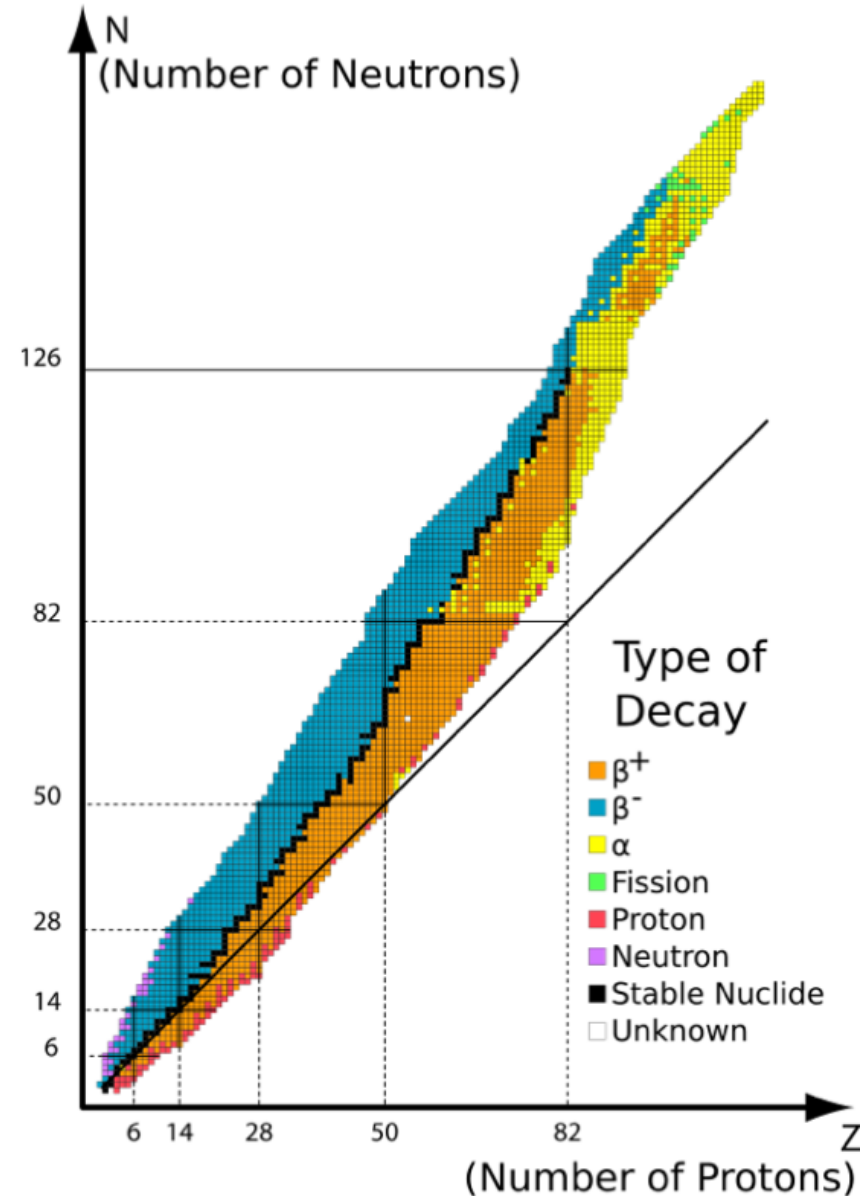
## The Weak Force: Oh, I didn't know that?

- Through a series of nuclear reactions, four protons (hydrogen nuclei) in the core of our Sun combine to form a helium nucleus emitting two positrons and two neutrinos and releasing 27 MeV of energy:



- Thermonuclear fusion--Perhaps the most important reaction for all life on planet Earth is caused by a fundamental force of nature that is rarely discussed in the classroom: Weak Interactions or the weak nuclear force

- Responsible for nearly all radioactive decay processes
- Beta decay is most common
- Theoretical understanding is at same level as Quantum Electro Dynamics

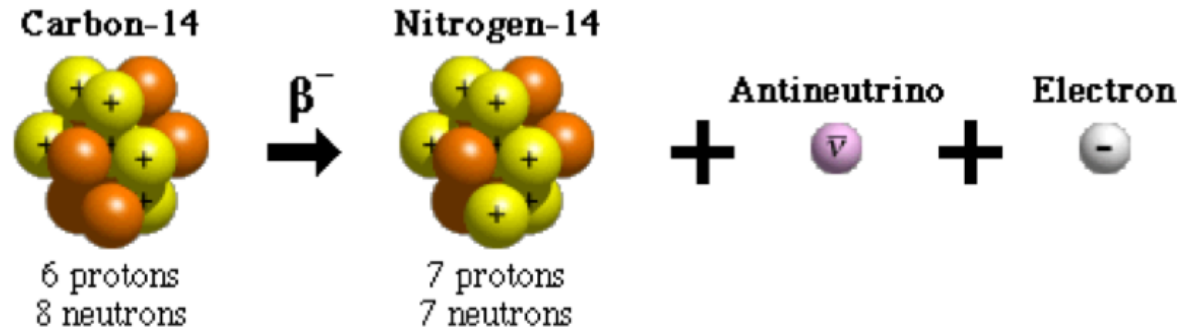




# Beta Decay Examples

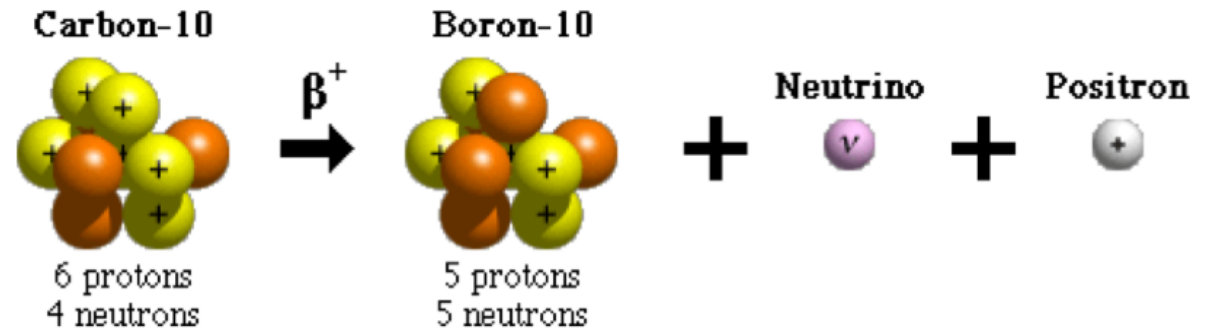
- $\beta^-$  decay:  $n \rightarrow p + \bar{\nu}_e + e^-$
- Moves nuclei up the periodic table ( $Z \rightarrow Z + 1$ )

## Beta-minus Decay



- $\beta^+$  decay:  $p \rightarrow n + \nu_e + e^+$
- Moves nuclei down in the periodic table ( $Z \rightarrow Z - 1$ )

## Beta-plus Decay

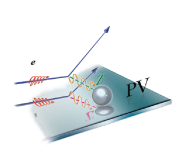




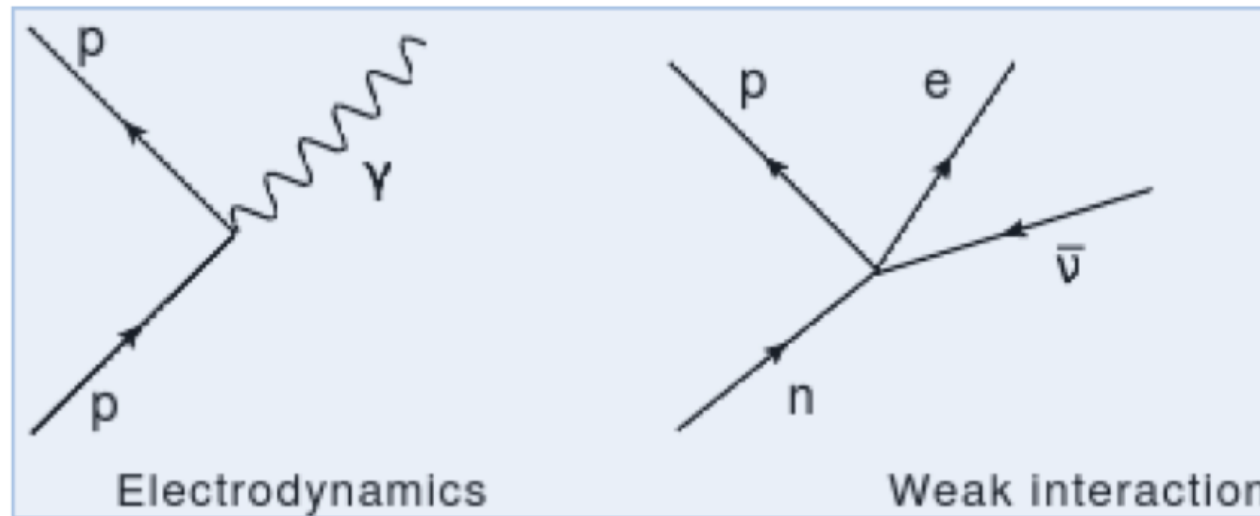
# Beta Decay – Nature’s Window into the Weak-nuclear Force

## A Quick History:

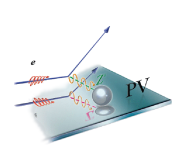
- 1899 Rutherford Rutherford classifies three types of radioactive emissions: alpha, beta, and gamma
- 1931 Pauli postulates existence of neutrino to explain non-discrete energy spectra of  $\beta$ -decay electrons
- 1933 Fermi develops theory to explain  $\beta$  decay -- precursor to theory for weak interaction
- 1956 Neutrino discovered by experiment.  $\bar{\nu}_e + p \rightarrow n + e^+$
- 1957 Parity Violation discovered in  $\beta$  decay of  $^{60}\text{Co}$



## Fermi's Interaction – Precursor to Weak Theory



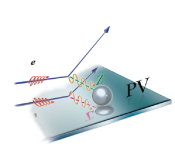
- Fermi's theory invented a physical mechanism for  $\beta$  decay
- 4-fermion contact interaction at single space-time point
- Modeled after electrodynamic field interactions -- where  $\vec{J}_E$  of a charged particle interacts with  $\vec{A}$  to create a photon
- For Fermi's theory, the ``weak'' current of pn-pair interacts with the ``weak'' current of  $e\bar{\nu}$ -pair
- Fermi's ``weak'' currents/potentials had vector form just as EM.



## First Neutrino Observations 1956

- Clyde L. Cowan, Frederick Reines (Awarded 1995 Nobel Prize)
- Experiment conducted near nuclear reactor ( $\sim 10^{13}$   $\nu$ 's /s/cm<sup>2</sup>)
- Two water tanks 12m underground and 11m from reactor
- Used inverse beta decay reaction:  
 $\bar{\nu}_e + p \rightarrow n + e^+$
- The  $e^+$  annihilated with an  $e^-$  producing two  $\gamma$  rays (detected)



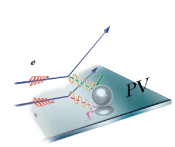


## Parity Symmetry

$$\mathbf{P}: \begin{bmatrix} x \\ y \\ z \end{bmatrix} \longrightarrow \begin{bmatrix} -x \\ -y \\ -z \end{bmatrix}$$

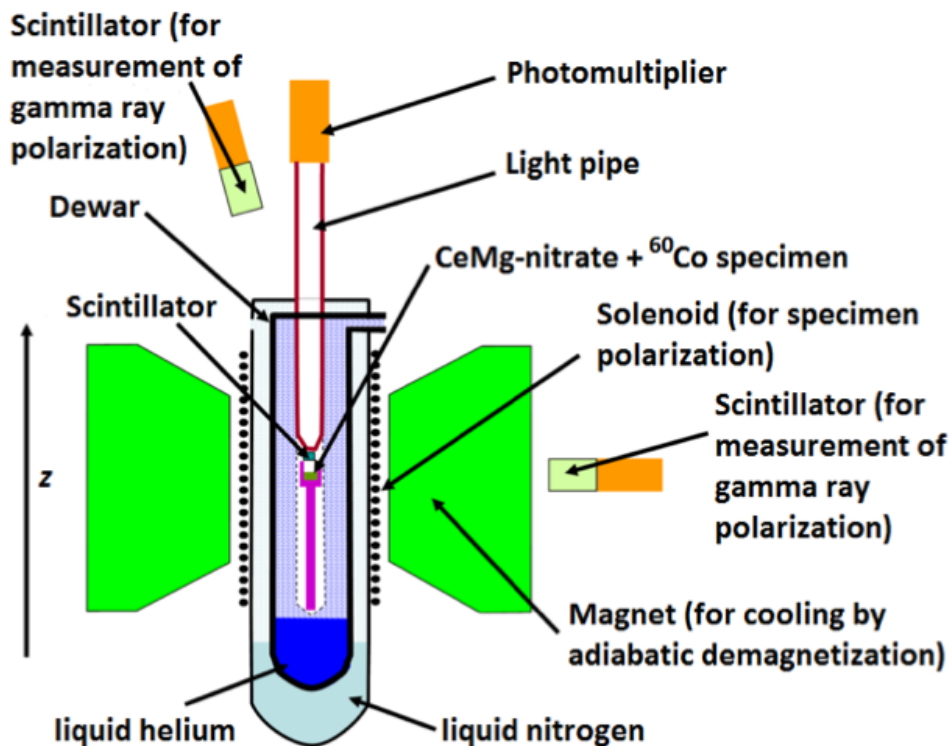
- Parity operation: Spatial reflection through the origin
- “Even” functions:  $\mathbf{P} f(x, y, z) \Rightarrow +f(x, y, z)$
- “Odd” functions:  $\mathbf{P} f(x, y, z) \Rightarrow -f(x, y, z)$
- *Classically*, scalar quantities ( $m, E, \rho, V, M, \dots$ ) are mainly “even” while vector quantities ( $\vec{x}, \vec{a}, \vec{F}, \vec{E}, \vec{A}, \dots$ ) are mainly “odd”
- *Quantum Mechanically*, if  $\mathbf{P}$  commutes with the Hamiltonian, then Parity is conserved (invariant or symmetric)
- Fundamental symmetry of nature known to be conserved in electromagnetism, strong interactions, and gravity





## Parity Violation Discovered in $\beta$ -decay: 1957

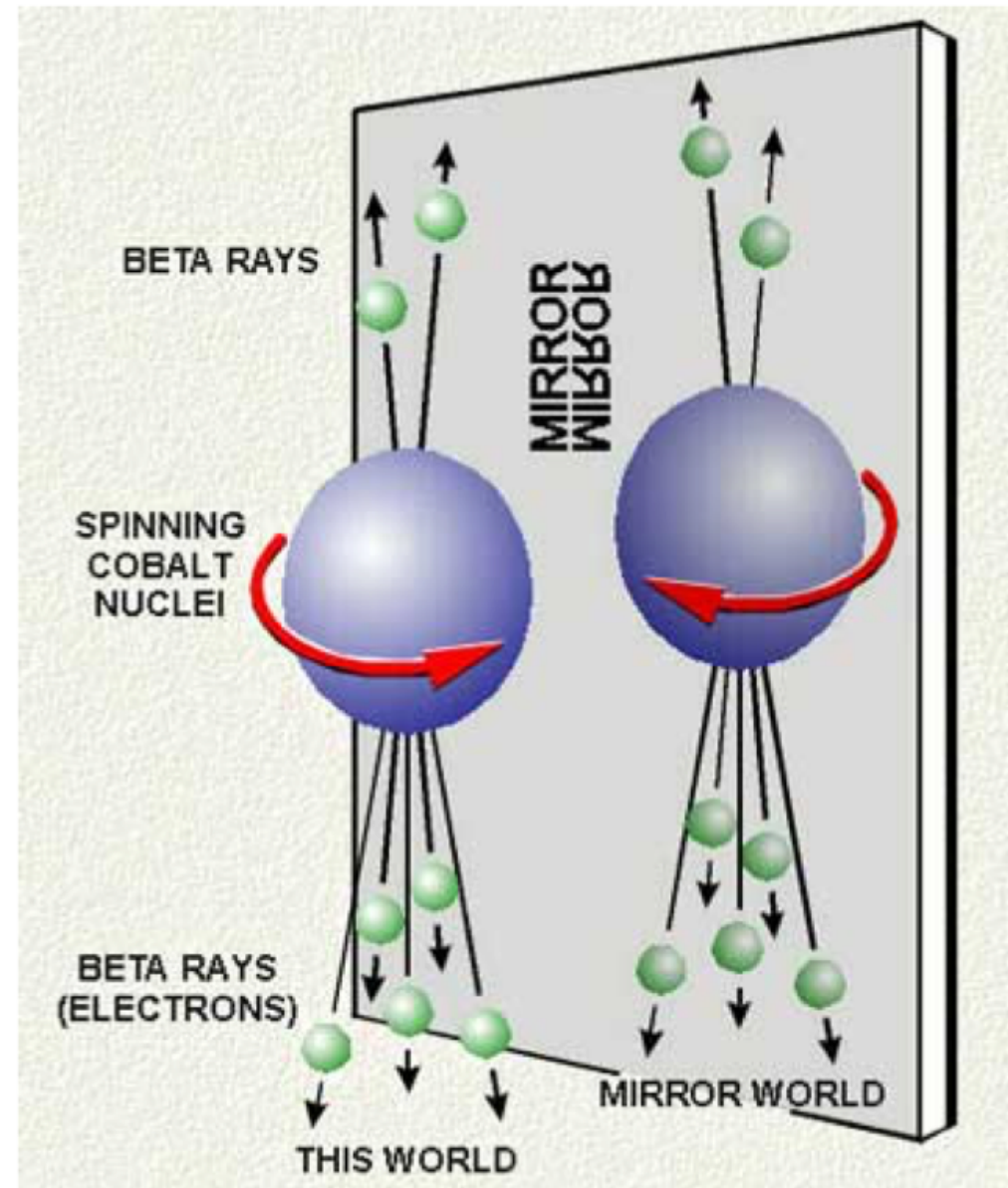
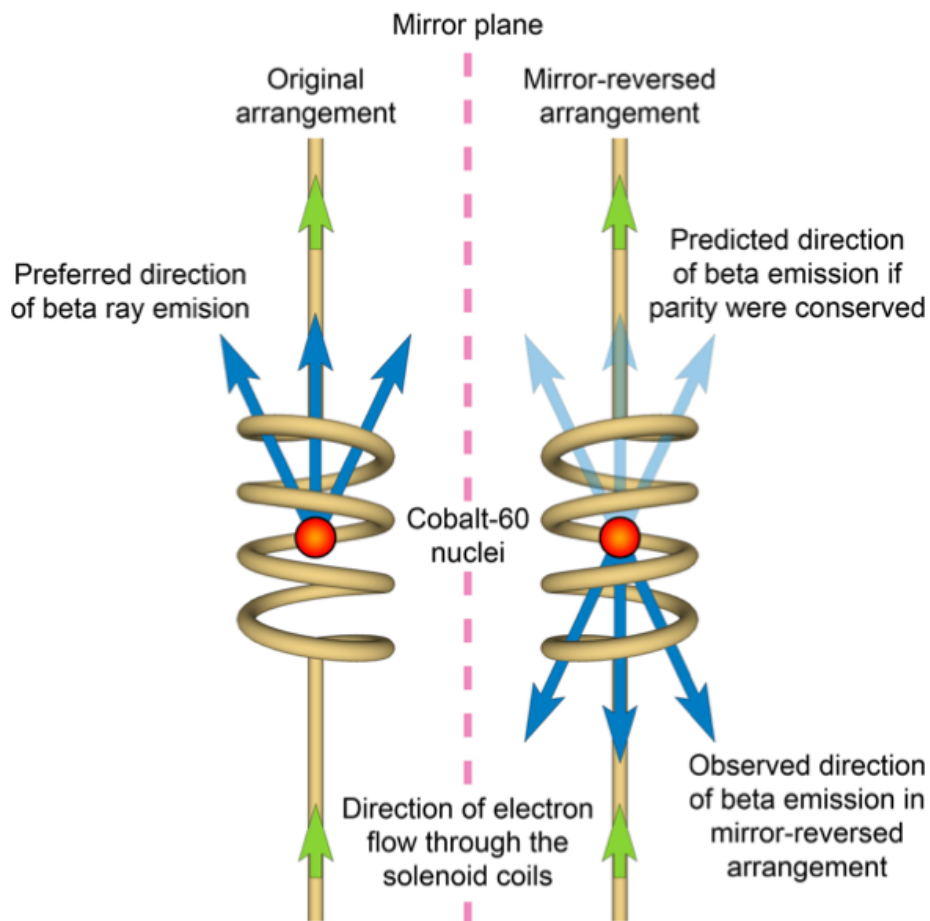
- Chien-Shiung (Madame) Wu Experiment
- Took place at NBS (now NIST)
- Studied  $\beta^-$  decay of super-cooled, spin-aligned  $^{60}\text{Co}$  nuclei
- $^{60}_{27}\text{Co} \rightarrow ^{60}_{28}\text{Ni} + e^- + \bar{\nu}_e + 2\gamma$
- Achieved 3 mK and 60% polarization





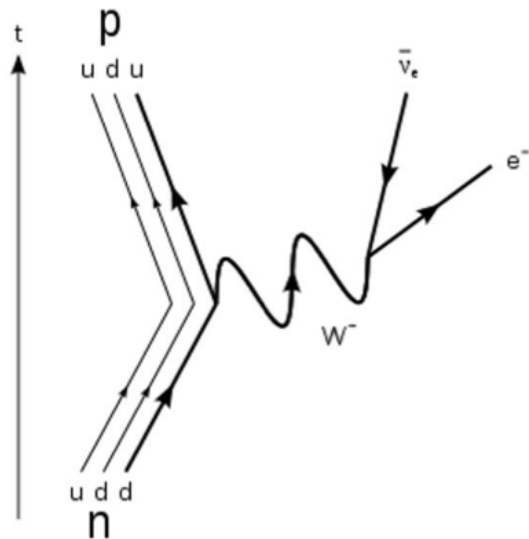
## Parity Violation Discovered in $\beta$ -decay: 1957

- Parity found to be maximally violated
- T.D. Lee and C.N. Yang awarded 1957 Nobel Prize





## $\beta^-$ Decay and Standard Model

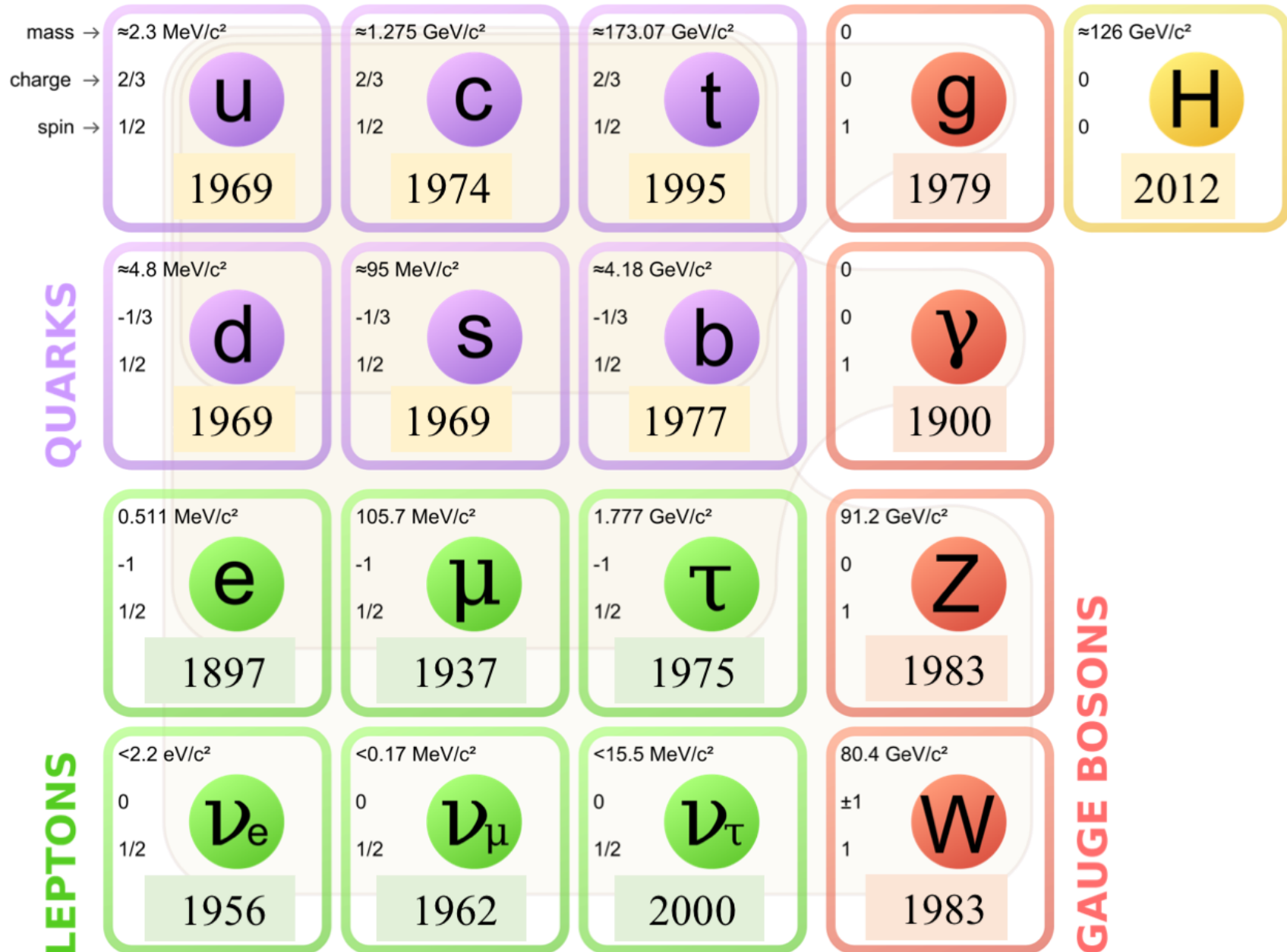


	mass → $\approx 2.3 \text{ MeV}/c^2$ charge → $2/3$ spin → $1/2$ <b>u</b> up	mass → $\approx 1.275 \text{ GeV}/c^2$ charge → $2/3$ spin → $1/2$ <b>c</b> charm	mass → $\approx 173.07 \text{ GeV}/c^2$ charge → $2/3$ spin → $1/2$ <b>t</b> top	mass → $0$ charge → $0$ spin → $1$ <b>g</b> gluon	mass → $\approx 126 \text{ GeV}/c^2$ charge → $0$ spin → $0$ <b>H</b> Higgs boson
<b>QUARKS</b>	mass → $\approx 4.8 \text{ MeV}/c^2$ charge → $-1/3$ spin → $1/2$ <b>d</b> down	mass → $\approx 95 \text{ MeV}/c^2$ charge → $-1/3$ spin → $1/2$ <b>s</b> strange	mass → $\approx 4.18 \text{ GeV}/c^2$ charge → $-1/3$ spin → $1/2$ <b>b</b> bottom	mass → $0$ charge → $0$ spin → $1$ <b><math>\gamma</math></b> photon	
	mass → $0.511 \text{ MeV}/c^2$ charge → $-1$ spin → $1/2$ <b>e</b> electron	mass → $105.7 \text{ MeV}/c^2$ charge → $-1$ spin → $1/2$ <b><math>\mu</math></b> muon	mass → $1.777 \text{ GeV}/c^2$ charge → $-1$ spin → $1/2$ <b><math>\tau</math></b> tau	mass → $91.2 \text{ GeV}/c^2$ charge → $0$ spin → $1$ <b>Z</b> Z boson	<b>GAUGE BOSONS</b>
<b>LEPTONS</b>	mass → $< 2.2 \text{ eV}/c^2$ charge → $0$ spin → $1/2$ <b><math>\nu_e</math></b> electron neutrino	mass → $< 0.17 \text{ MeV}/c^2$ charge → $0$ spin → $1/2$ <b><math>\nu_\mu</math></b> muon neutrino	mass → $< 15.5 \text{ MeV}/c^2$ charge → $0$ spin → $1/2$ <b><math>\nu_\tau</math></b> tau neutrino	mass → $80.4 \text{ GeV}/c^2$ charge → $\pm 1$ spin → $1$ <b>W</b> W boson	

- Julian Schwinger modifies Fermi's theory to incorporate parity violating potential term (V-A) and idea of intermediate vector bosons; Glashow, Weinberg, and Salam 1979 Nobel Prize
- $W^\pm$  only couples to left-handed particles and right-handed anti-particles
- $Z^0$  couples predominantly to left-handed particles



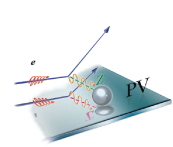
# Standard Model of Elementary Particles



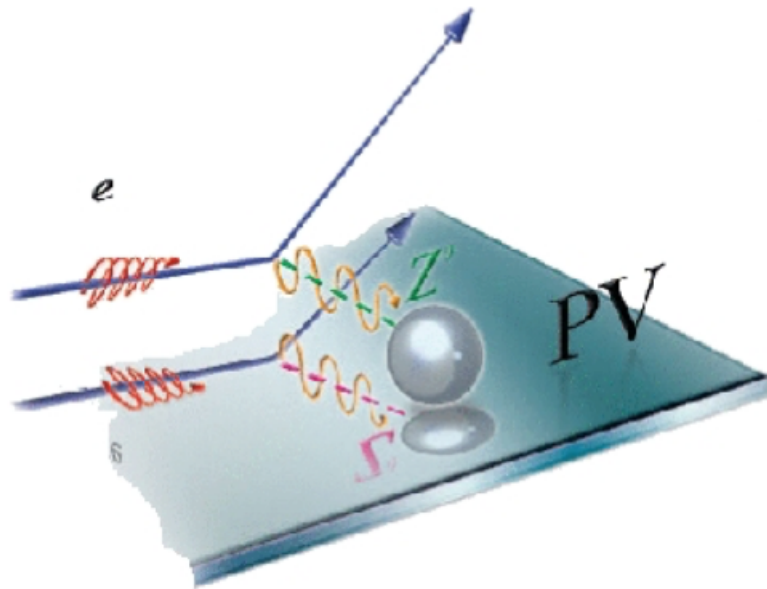


## Parity Violation and Electron Scattering

- Electron scattering experiments make first measurement of neutral ( $Z^0$ ) weak current in late 1970's (at SLAC)
- PVeS experiments scatter longitudinally spin-polarized electron beams (with relatively low energies) off unpolarized, fixed nuclear targets
- Since  $Z^0$  couples to opposite spin (helicity) particles with different strengths, one can measure cross section ( $\sigma$ ) differences for opposite helicity beams to access the neutral
- Following technological breakthroughs (at SLAC), high beam polarizations and fast helicity reversals become possible
- PVeS experiments measure an Asymmetry:  $A_{PV} = \frac{\sigma_R - \sigma_L}{\sigma_R + \sigma_L}$
- Since weak scattering process is only tiny fraction of total  $\sigma$ , PV asymmetries are tiny and difficult to measure accurately



$A_{PV}$ : Dominated by Electroweak Interference



$$A^{PV} = \frac{\sigma_R - \sigma_L}{\sigma_R + \sigma_L}$$

$$\propto \frac{\text{[diagram: } \gamma \text{ and } Z^0 \text{ exchange]} \cdot 10^{-4} Q^2}{\text{[diagram: } \gamma \text{ exchange}]^2} \sim \frac{10^{-4} Q^2}{\text{GeV}^2}$$

- Amplitude for Scattering Process:  $S_{\text{tot}} \rightarrow S_{\text{em}} + S_{\text{w}}$ , but cross section  $\sigma \rightarrow |S_{\text{tot}}|^2 = |S_{\text{em}}|^2 + |S_{\text{w}}|^2 + 2S_{\text{em}}S_{\text{w}}$  (quantum interference term)
- Since  $\sigma^{\text{em}}_R = \sigma^{\text{em}}_L$  and  $|S_{\text{w}}|^2$  is negligible,

$$A_{PV} \rightarrow \frac{2S_{\text{em}}S_{\text{w}}}{2|S_{\text{em}}|^2} = \frac{S_{\text{w}}}{S_{\text{em}}} \sim 10^{-4} \cdot Q^2$$

where  $Q^2$  is 4-momentum transferred during interaction (GeV)



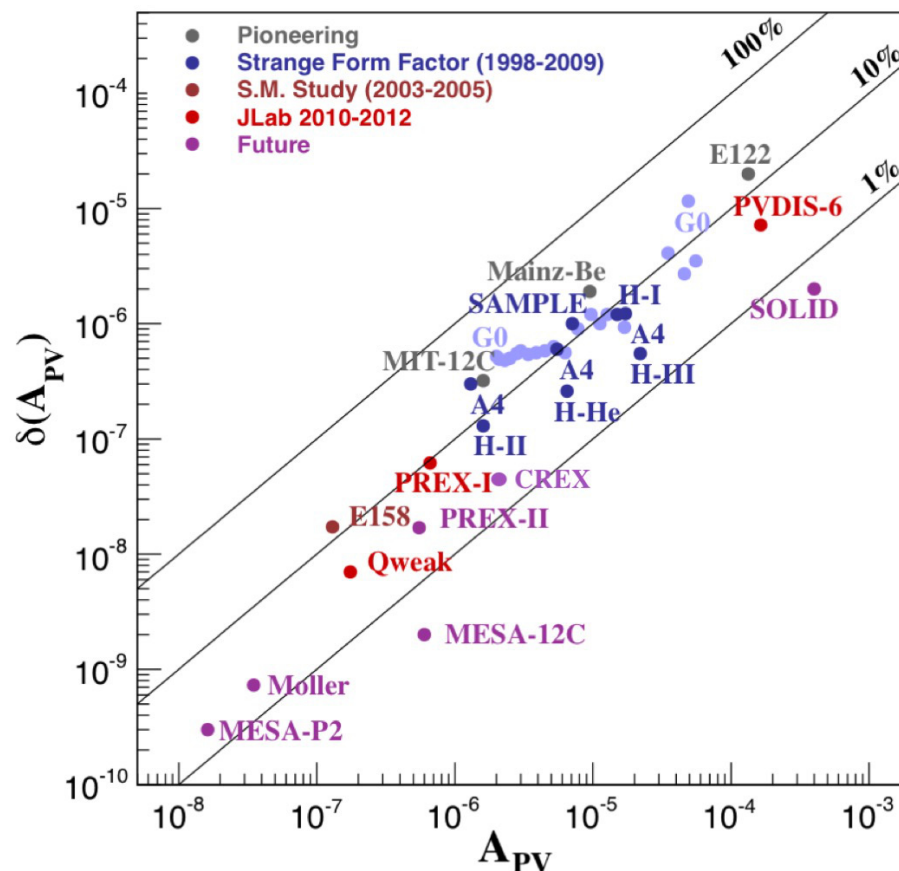
### 3 Decades of Technical Progress

photocathodes, polarimetry, high power cryotargets, nanometer beam stability, precision beam diagnostics, low noise electronics, rad-hard det

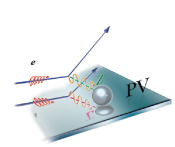
PVeS Experiment Summary

- 1st generation
- 2nd generation
- 3rd generation
- 4th generation

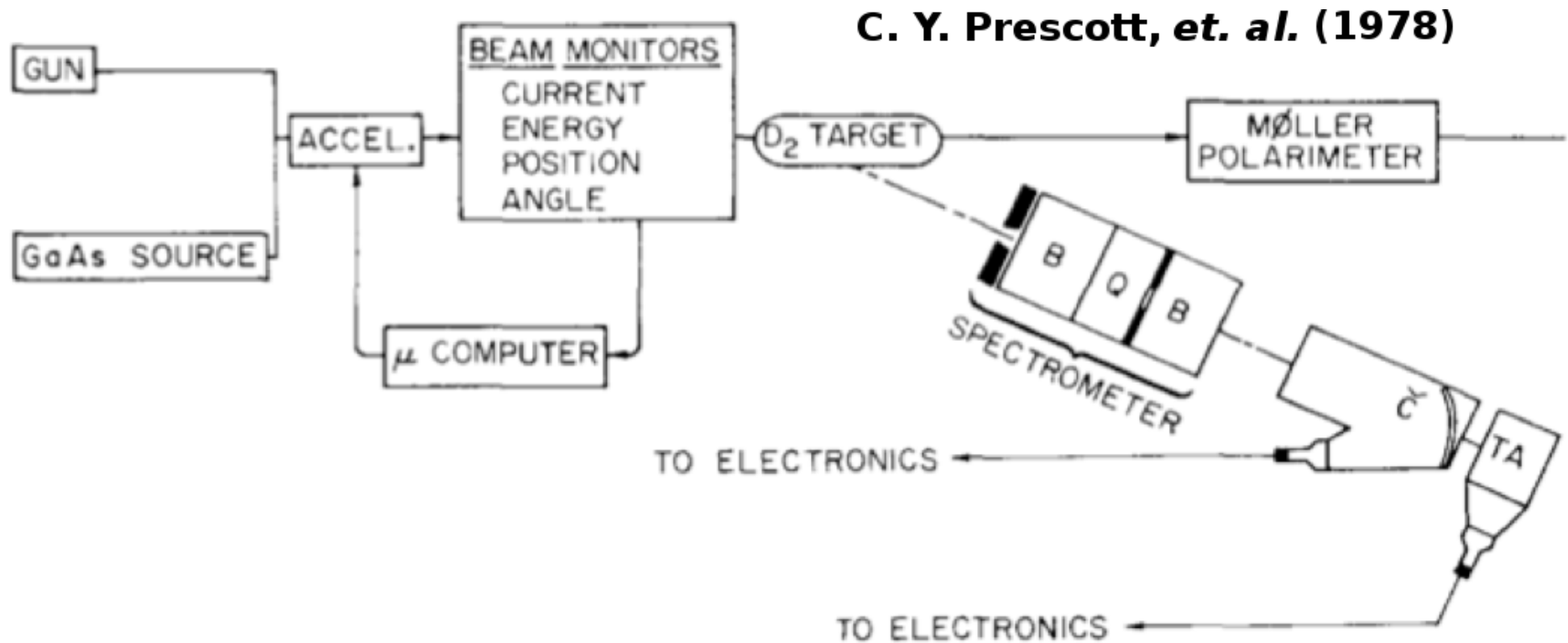
E122 – 1<sup>st</sup> PVES Expt (late 70’s at SLAC)  
 Mainz & MIT-Bates in mid 80’s  
 JLab program launched in mid 90’s  
 E158 at SLAC meas PV Møller scattering



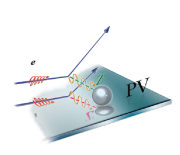
- Parity-violating electron scattering has become a precision tool



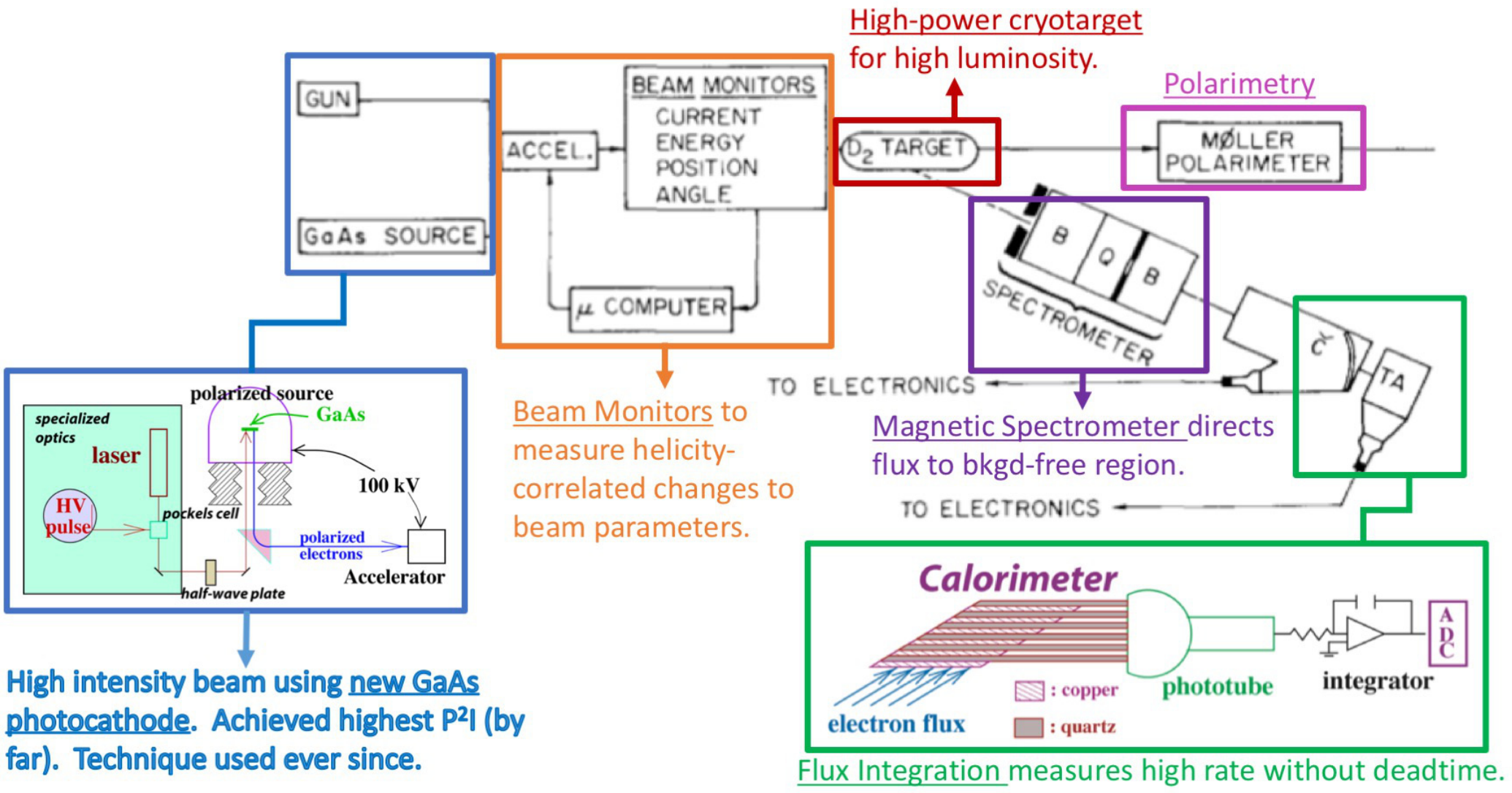
# Blueprint of a PVES Experiment (E122 at SLAC)







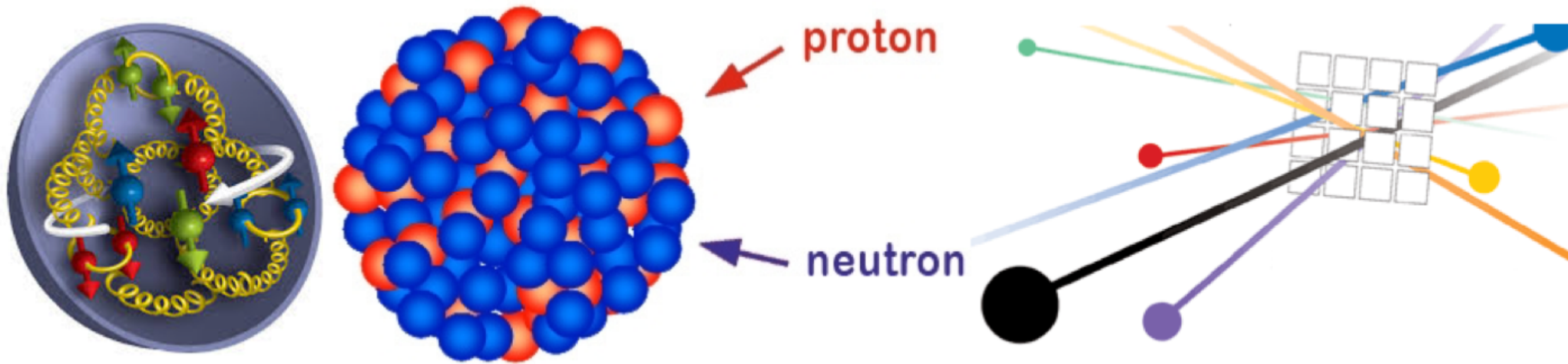
# Anatomy of a PVES Experiment (E122 at SLAC)

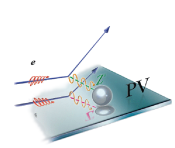




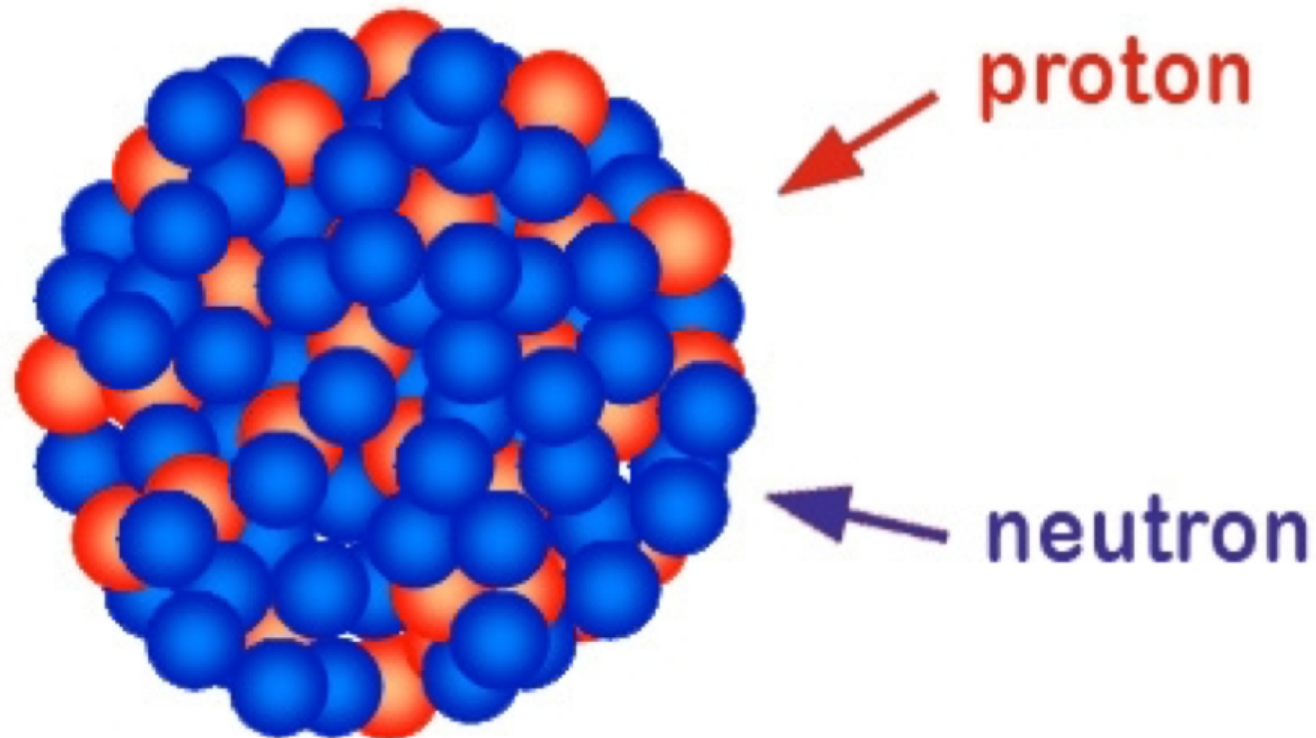
## Selected Applications of PVeS

- **Strange Quarks:** What is the role of strange quarks in the electromagnetic structure of the proton or nucleon?
- **Size of Nucleus:** What is the size of a neutron-rich, complex nucleus? What is  $R_n$ ,  $n_{\text{skin}}$ ? Implications for Neutron Stars? LIGO overlap
- **BSM Searches:** Searching for physics Beyond the Standard Model. Obvious Motivations here: Dark sector matter,...





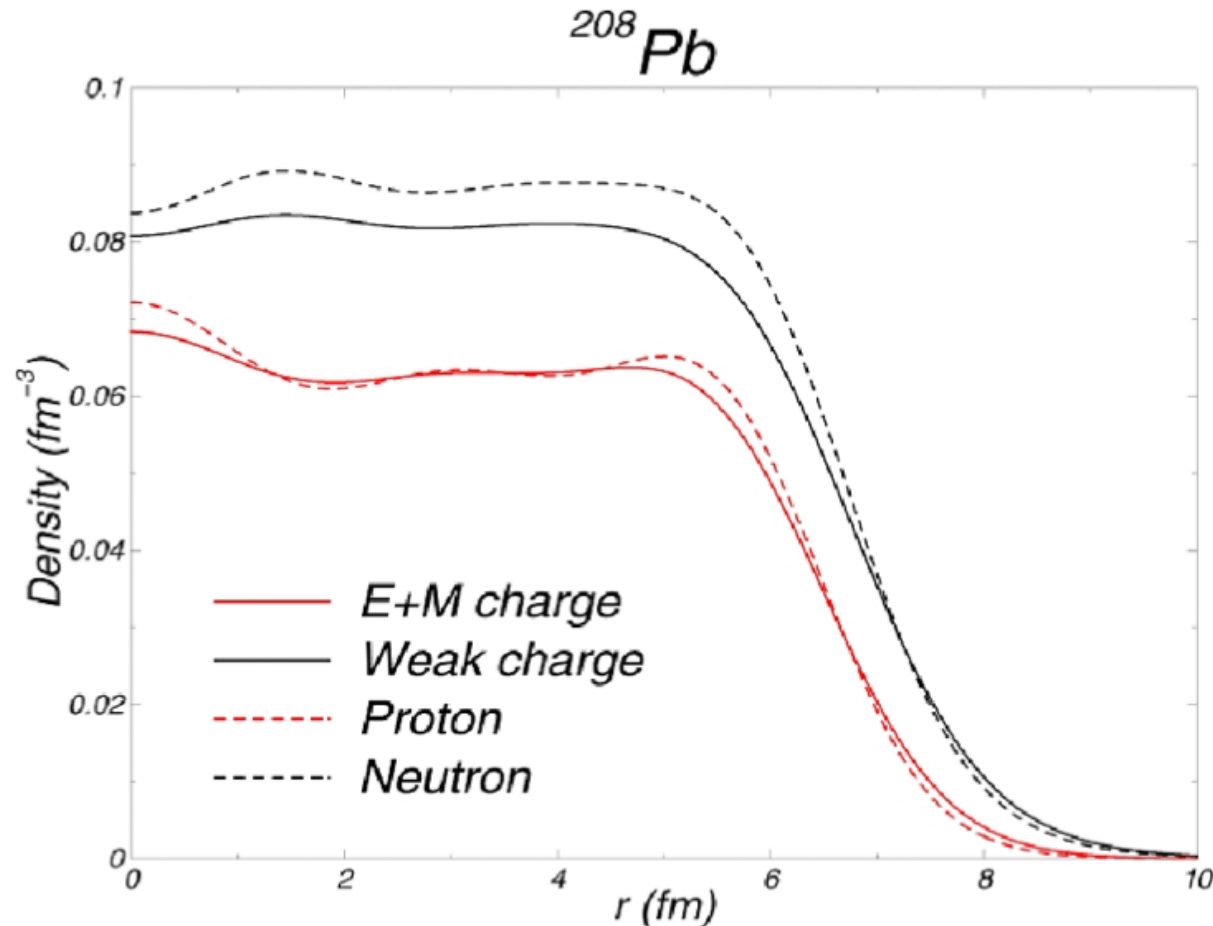
What is the size of a  $^{208}\text{Pb}$  nucleus (82p+126n)?



- What do we mean by size? The mass radius, the charge radius?
- PREX (Pb Radius EXperiment) addresses this question in a unique way: Uses a “Weak” nuclear force probe to measure how much neutrons stick out past protons in  $^{208}\text{Pb}$  nucleus (called the Neutron “skin”)
- CREX (Calcium Radius EXperiment) performs same measurement but on  $^{48}\text{Ca}$



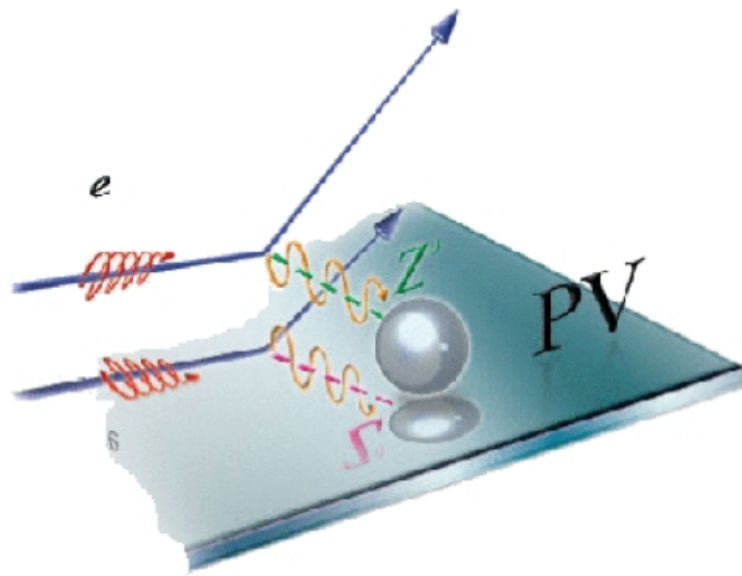
# Mass versus EM Charge Radii of $^{208}\text{Pb}$



- Electromagnetism: Force mediated by  $\gamma$  exchange; protons have EM charge “+e” while neutrons have 0...
- Weak Nuclear: Force mediated by  $Z^0$  and  $W^\pm$ ;  
**neutrons have ~12 times more Weak charge than protons**



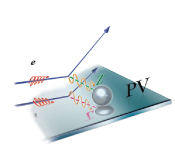
## PREX Measurement (Pb Radius EXp)



$$A^{PV} = \frac{\sigma_R - \sigma_L}{\sigma_R + \sigma_L}$$

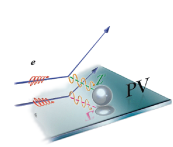
$$\propto \frac{\langle \gamma \rangle \langle Z^0 \rangle}{\langle \gamma \rangle^2} \sim \frac{10^{-4} Q^2}{\text{GeV}^2}$$

- Uses  $\sim 1$  GeV elastically scattered electrons (at  $\pm 5$  deg) off 0.5 mm thick isotopically pure  $^{208}\text{Pb}$  target
- $e^-$  beam is longitudinally spin-polarized, target is unpolarized
- Measurement relies on the maximal parity-symmetry violating nature of the Weak force
- $e^-$ 's dominant interaction is EM, but it can also interact via the Weak force; but it does so predominately for only one of the polarization states and not the other -- thus the Asymmetry ( $A_{PV}$ ) measurement



## Motivation: Nuclear Radii in Heavy Nuclei

- Measurements are important for understanding the strong nuclear force
- Calculations are difficult due to non-pQCD regime and complicated due to many-body physics
- Interesting for:
  - Fundamental nuclear structure
  - Isospin dependence and nuclear symmetry
  - Dense nuclear matter and neutron stars
- Proton radius is relatively easy - electromagnetic probes
- Neutron radius is difficult:
  - Weakly couples to electroweak probes
  - Hadronic probes have considerable uncertainty
  - Theory has range of  $R_n - R_p$  for Pb of 0 - 0.4 fm

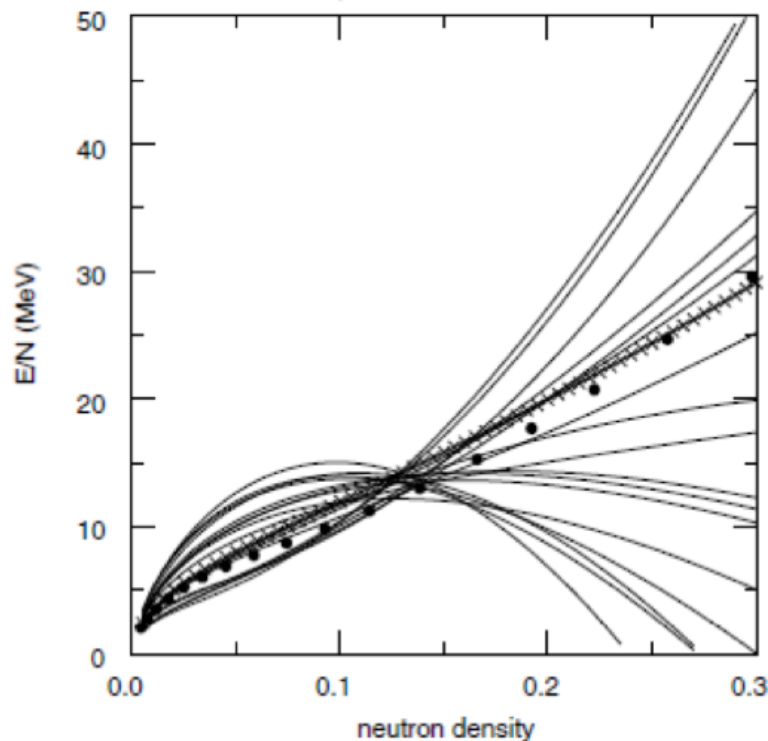


## Motivation: What do we learn from $R_n$ ?

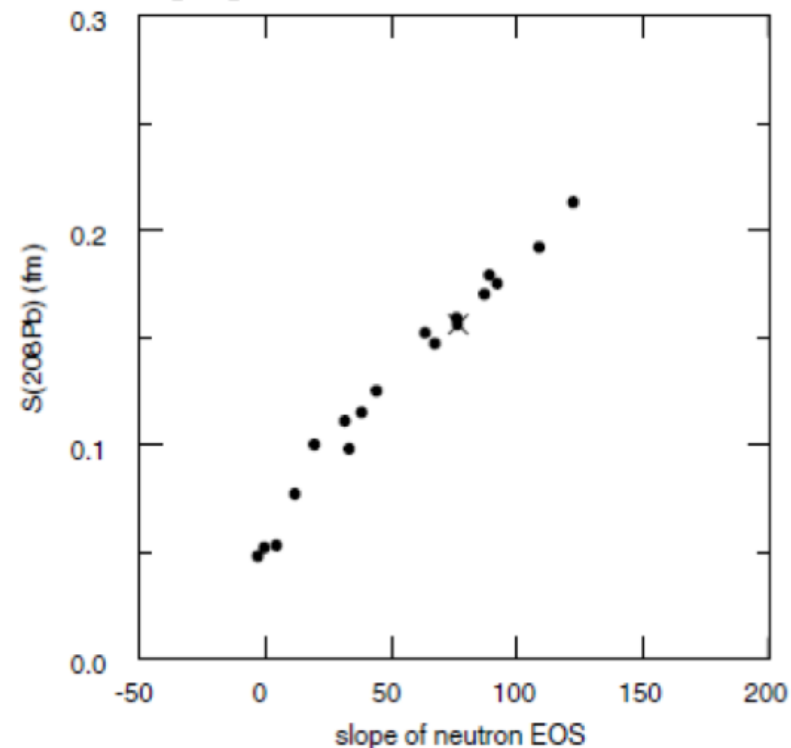
- Constraints on Equation of State (EOS) and symmetry energy of neutron rich matter -- the energy cost for asymmetric matter ( $N \neq Z$ )
- Slope of EOS can be used to constrain potential models

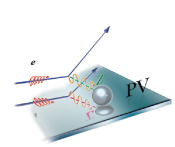
B.A. Brown, PRL 85, 5296 (2000)

Neutron EOS for 18 Skyrme model sets;  
dots are FP calcs; crosses are SkX;  
neutron density in neutron/ $\text{fm}^3$



Lead neutron skin vs. corresponding slope parameter of the 18 EOS sets



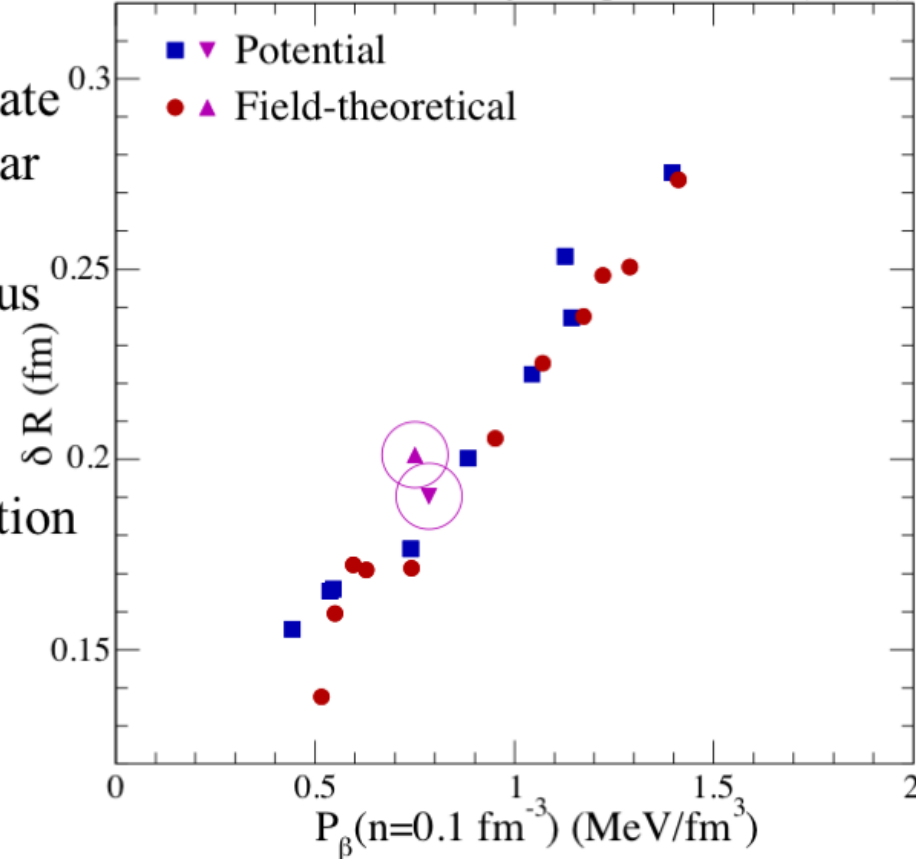


## Motivation: Neutron Stars

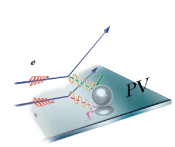
- Neutron star structure is better understood with measurements of  $R_n$
- Larger  $R_n$  (and thus  $\Delta R$ --the skin) correlates with larger pressure  $P$
- Larger  $P$  pushes neutrons out against surface tension increasing  $R_n$ :
  - Thus measurement of  $R_n$  (and  $\Delta R$ ) could calibrate the pressure of neutron star matter at sub-nuclear densities
  - Combining  $\Delta R$  with observed neutron star radius could allow access to pres.-dens. rel't inside neutron stars
- Additionally, symmetry energy governs proton fraction
  - Direct URCA cooling depends on processes:
 
$$n \rightarrow p + e^- + \bar{\nu}$$

$$e^- + p \rightarrow n + \nu$$
  - Larger symmetry energy gives larger proton fraction (need 11\%)

A.W. Steiner *et al.*, Phys Rep 411, 325 (2005)







## Methods used to Measure $R_n$

- Hadronic Probes

- Elastic  $pN$ ,  $\vec{p}N$ ,  $nN$ ,  $\pi^\pm N$
- $\pi^0$  photo-production (Kruche, et al.)
- GDR
- Antiproton scattering
- Have theoretical uncertainty

- Electroweak Probes

- Parity violating electron scattering
- Atomic parity violation
- “Clean” measurements, fewer systematics
- Technically challenging



# Non-Parity Violating Electron Scattering

- Electron scattering  $\gamma$  exchange provides  $R_p$  through nucleus form factors ( $F(Q^2)$ )

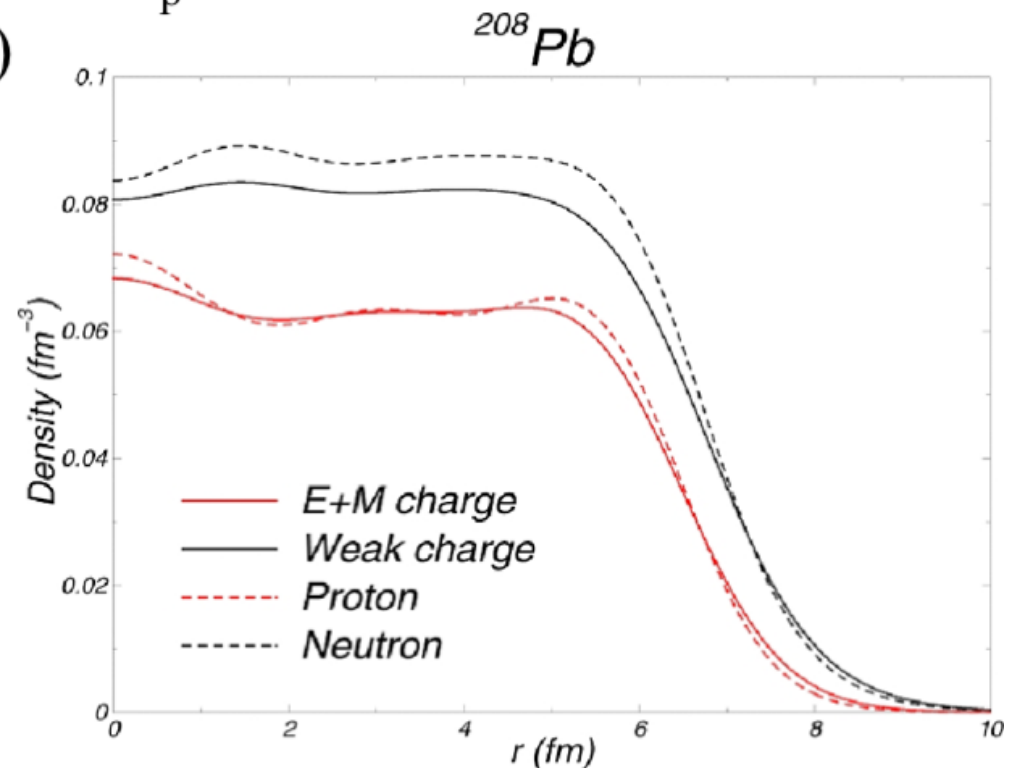
➤ For spin 0 nucleus:

$$\frac{d\Omega}{d\sigma} = \frac{\alpha^2 \cos^2 \frac{\theta}{2}}{4E^2 \sin^4 \frac{\theta}{2}} |F(Q^2)|^2$$

➤ In limit of small  $Q^2$ :

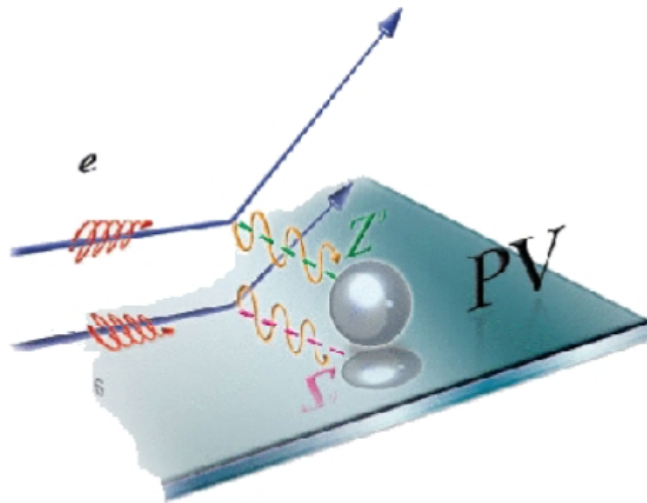
$$F(Q^2) \approx F(0) + \left. \frac{dF}{dQ^2} \right|_{Q^2=0} + \dots = \int \rho(\vec{x}) d^3x - \frac{1}{6} Q^2 \langle r_{\text{charge}}^2 \rangle$$

➤ So small  $Q^2$  measurements give density and RMS electromagnetic radius (dominated by  $R_p$ )





# Parity Violating Electron Scattering



$$A^{PV} = \frac{\sigma_R - \sigma_L}{\sigma_R + \sigma_L}$$

$$\propto \frac{\text{[Diagram: } \gamma \text{ and } Z^0 \text{ exchange]} \cdot 10^{-4} Q^2}{\text{[Diagram: } \gamma \text{ exchange}]^2 \cdot \text{GeV}^2}$$

- The  $e^-$  can also exchange a  $Z^0$  which is parity violating (PV)
- $Z^0$  primarily couples to the neutron, since:

$$Q_{\text{weak}}^{\text{proton}} \propto 1 - 4\sin^2\theta_W \approx 0.076, \quad Q_{\text{weak}}^{\text{neutron}} = -1$$

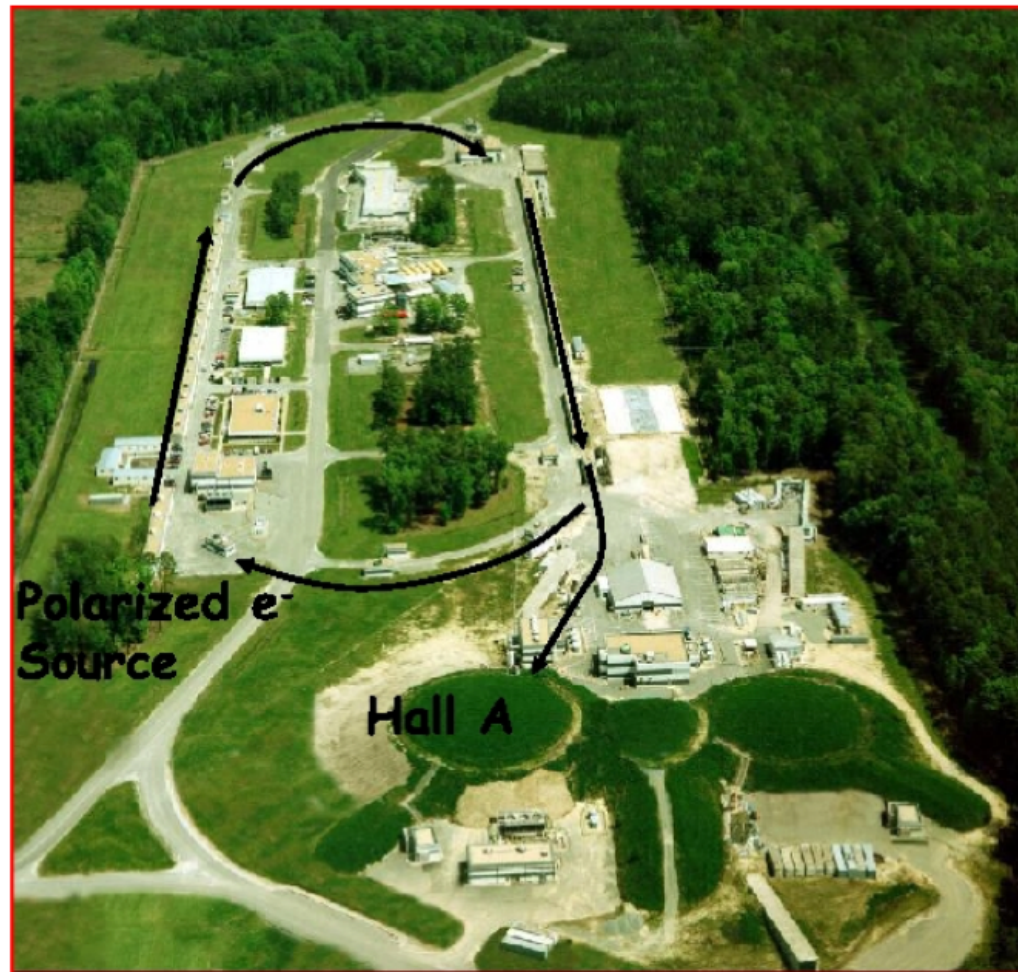
- Detectable in PV asymmetries of  $e^-$  with opposite helicities
- In Born approximation,  $Q^2 \ll M_Z$ , from  $\gamma - Z$  interference:

$$A_{PV} = \frac{G_F Q^2}{4\pi\alpha\sqrt{2}} \left[ 4\sin^2\theta_W - 1 + \frac{F_n(Q^2)}{F_p(Q^2)} \right], \quad F_n(Q^2) = \frac{1}{4} \int d^3r' j_0(qr) \rho_n(r)$$

- For fixed target experiment, typical  $A_{PV} \sim 10^{-8} - 10^{-4}$



# JLab's CEBAF is Excellent Facility for PV Measurements



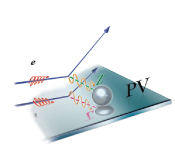
- High quality polarized beam,  $P_e \sim 85 - 90\%$
- PV expt's need quiet beam parms over helicity windows:  
 $\Delta x < 10 \mu\text{m}$                        $\Delta x' < 2 \mu\text{rad}$                        $\Delta E/E < 10^{-3}$



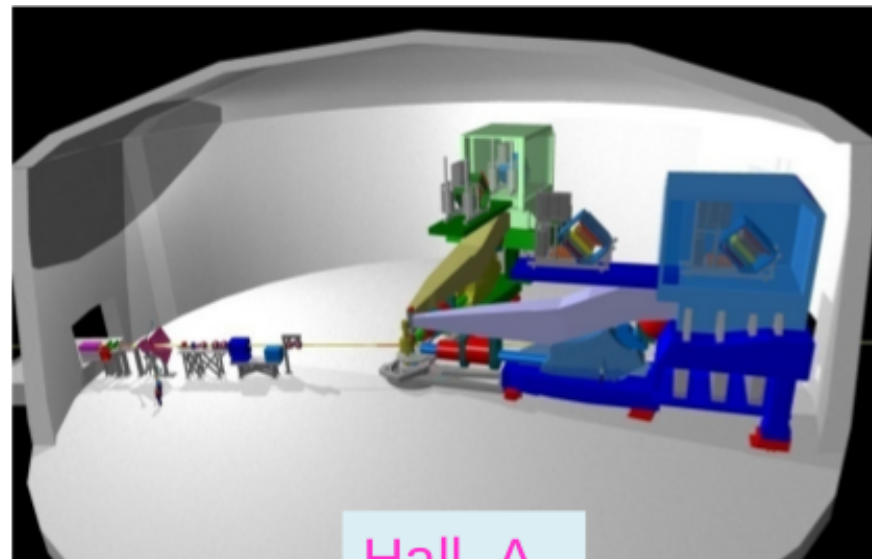
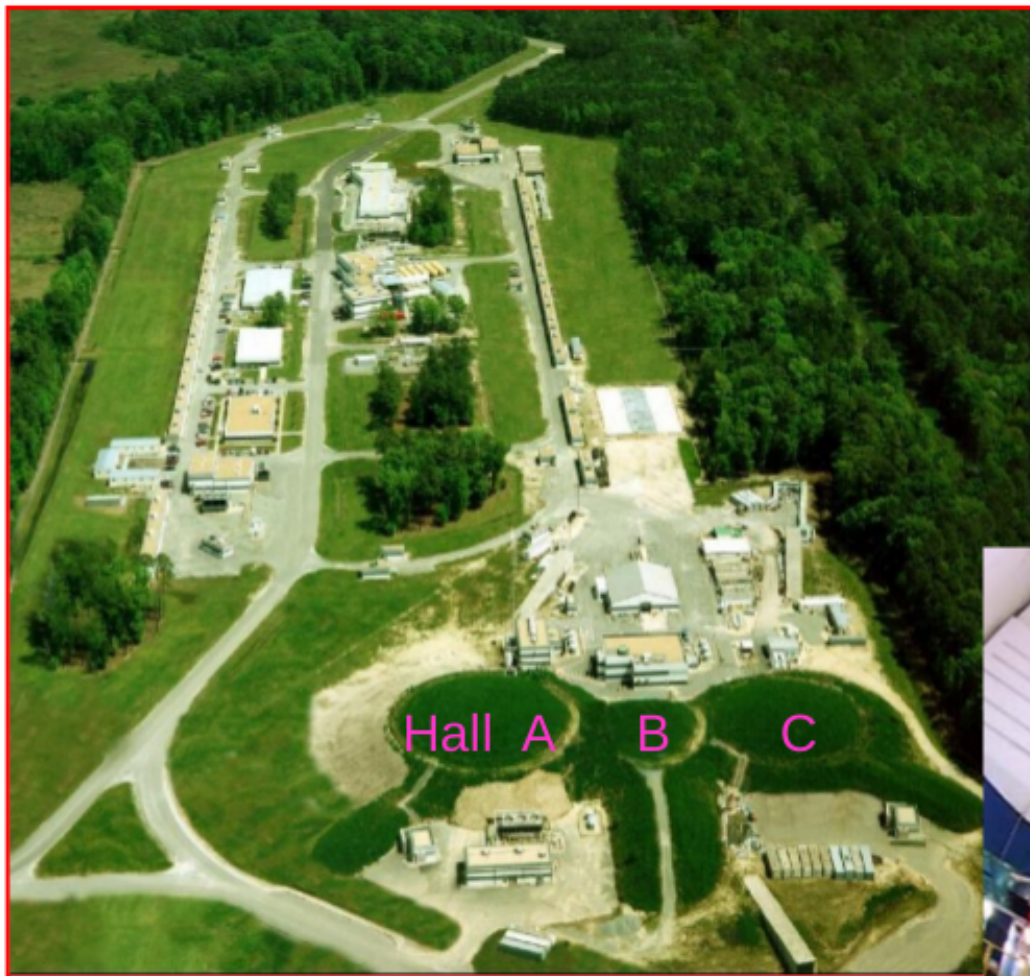
## PREx Measurement

### PREx measures $R_n$ of $^{208}\text{Pb}$

- Lead is nice because:
  - Excess of neutrons (44 more--with some expected to form a neutron-rich skin)
  - Doubly magic nucleus (82 protons, 126 neutrons)
  - Nearest excited state is 2.6 MeV from elastic peak (possible to exclude inelastics using HRS)
- Ran in Spring 2010 (approved 30 PAC days)
- $E_e = 1.063 \text{ GeV}$ ,  $\theta_e \approx 5^\circ$ ,  $Q^2 \approx 0.009 \text{ GeV}^2$
- $I_e \sim 50 - 75 \mu\text{A}$
- Proposed uncertainty on  $A_{\text{PV}}$  of 3%,  $R_n \sim 1\%$
- Uncertainty dominated by statistical error



# Jefferson Lab Hall A (Newport News, Virginia)



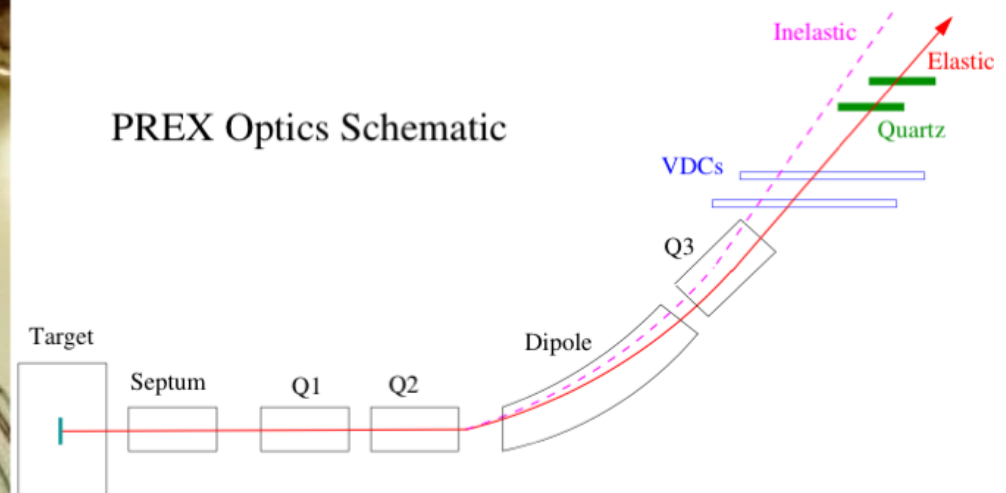
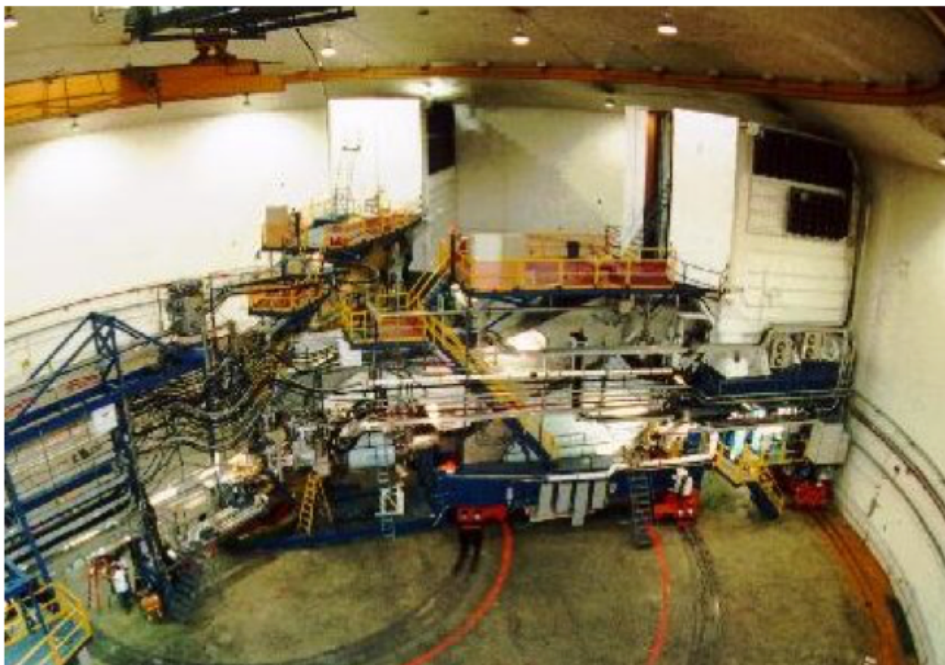
Hall A

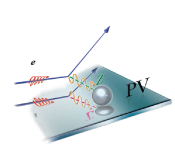




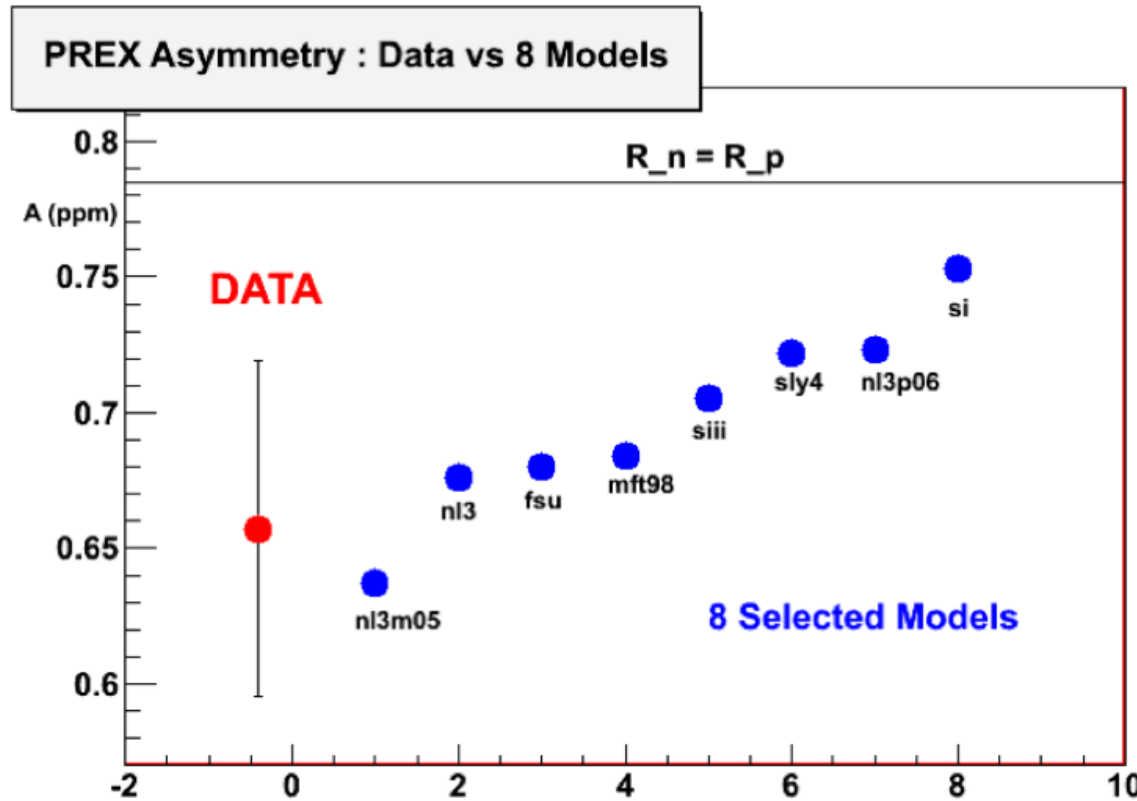
## Experimental Setup

- Std. Hall A HRS's with detector huts well shielded against bkgds
- Run dual, symmetric arms-- cancels out  $A_{\text{trans}}$  and other systematics
- Use septum magnet to bend  $5^\circ$  to  $12.5^\circ$
- Upgraded polarimetry (non-inv. Compton  $\sim 1\%$ , Inv. Moller  $\sim 1\%$ )
- 0.5mm thick Pb foil in between two 0.15mm Diamond targets ( $\sim 1 \text{ in}^2$ ) with cryogenically cooled frame; used fast rastered beam
- Quartz Cerenkov detectors with 18-bit integrating ADCs



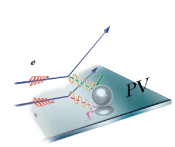


# PREX Result



- Set 95% CL on existence of neutron skin
- $R_n = 5.78^{+0.15}_{-0.17}$  fm,  $\Delta R = R_n - R_p = 0.34^{+0.15}_{-0.17}$  fm
  - Each model of neutron density is folded into numerical solution of Dirac eqn with Coulomb and weak axial potential
  - Full acceptance (apertures, septum optics, detectors) applied to  $A_{PV}$





## Result and Error Budget

$$A_{PV} = 0.658 \pm 0.0604 \pm 0.0130 \text{ ppm}$$
$$\pm 9.2\%(\text{stat}) \pm 2.0\%(\text{syst})$$

Contributions	abs (ppm)	rel (%)
Polarization	0.0071	1.1%
Detector Lin.	0.0071	1.1%
Beam Corrections	0.0072	1.1%
$Q^2$	0.0028	0.4%
$^{12}\text{C}$ Asymmetry	0.0025	0.4%
Transverse Pol.	0.0012	0.2%
BCM Lin.	0.0010	0.1%
Target Thick	0.0006	0.1%
Rescattering	0.0001	0.0%
Inelastic Cont.	0.0000	0.0%



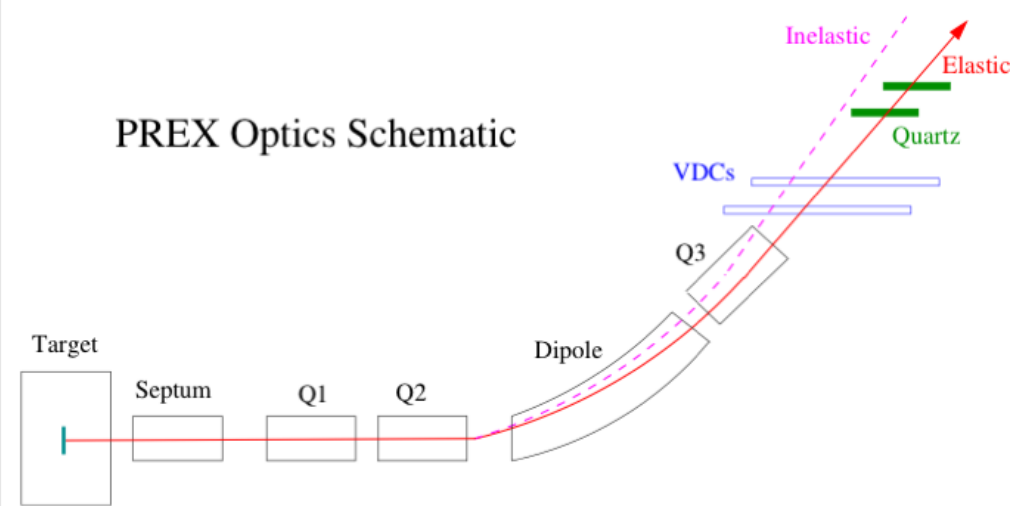
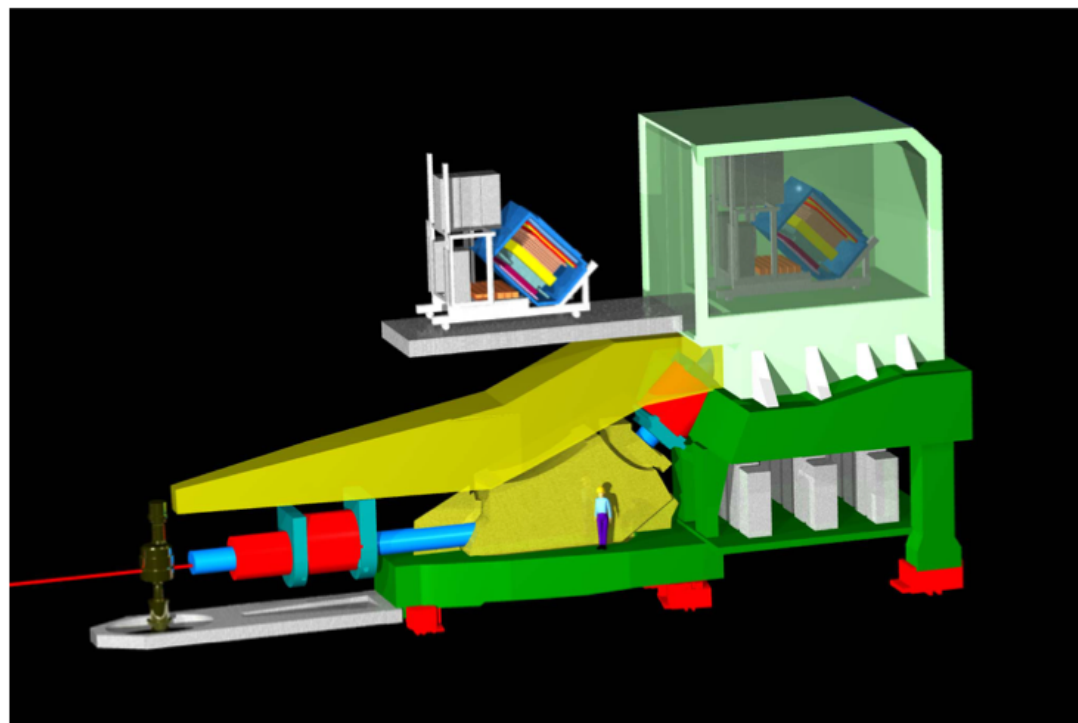
## Summary (PREX)

- PREx exp. ran March - June 2010 to measure  $R_n$  on  $^{208}\text{Pb}$ ;  
Published in Phys. Rev. Lett. 108, 112502 (2012)
- After all corrections:  $A_{\text{PV}}^{\text{Pb}} = 0.658 \pm 0.0604$  (9.2%)  $\pm 0.0130$   
(2.0%) ppm (statistics dominated uncertainty)
- From simple fit over calcs:  $R_n = 5.78_{-0.17}^{+0.15}$  fm
- Neutron skin:  $R_n - R_p = 0.34_{-0.17}^{+0.15}$  fm
- Established existence of neutron skin with 95% CL
- PREx-II experiment set to run in late 2016 – will improve stat. err of PREx-I by factor of 3
- PREx-II precision will better discriminate between models allowing predictions relevant for the description of neutron stars



## Experimental Setup (Spectrometer & Detectors)

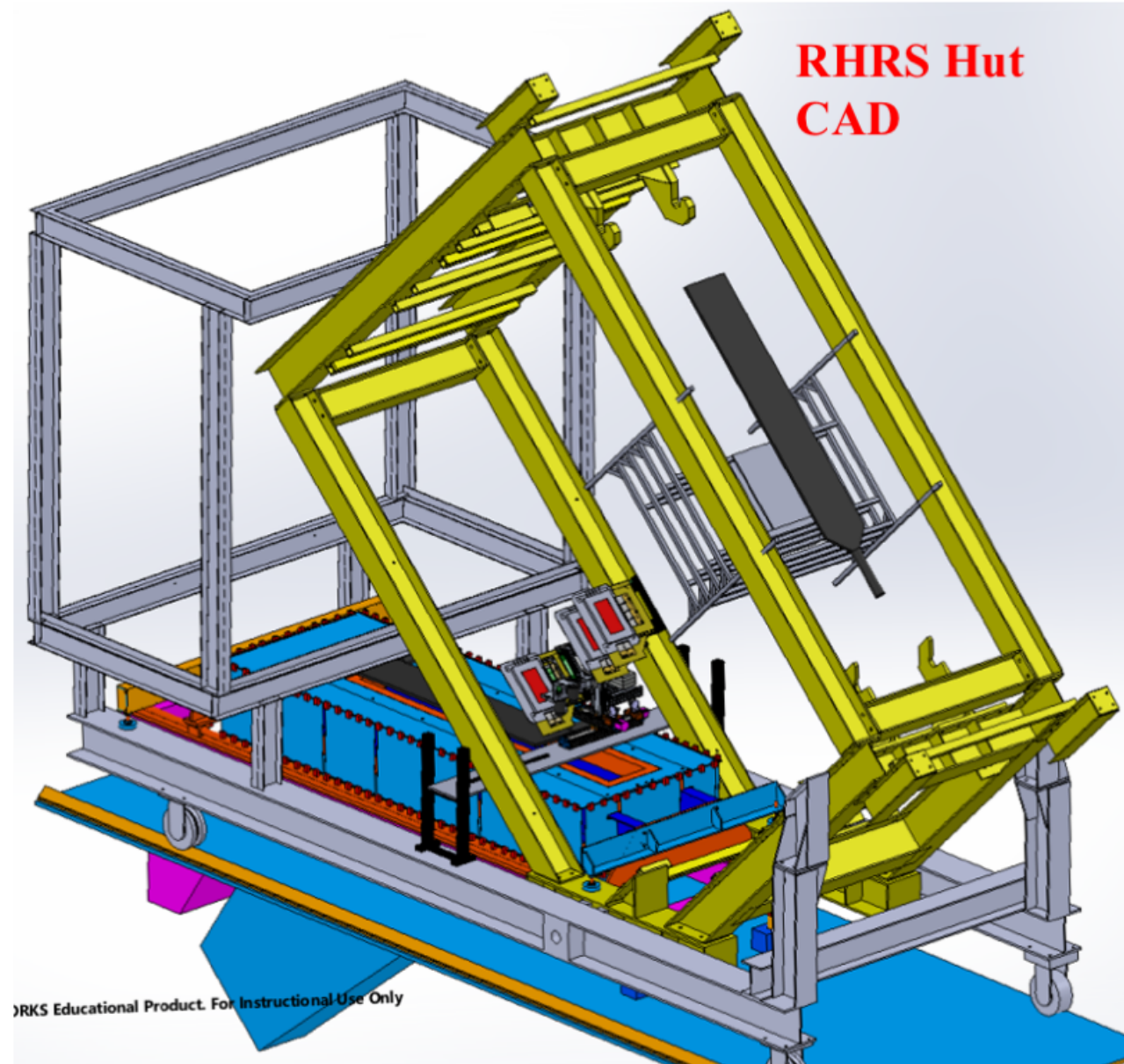
- Thin quartz Cerenkov detectors with PMTs used to measure scattered electron flux
- Highly relativistic electrons travel faster than light travels through the quartz, thus creating Cerenkov radiation (UV light)
- High purity quartz necessary due to its extreme radiation hardness (maintains transparency during high doses (Grad) of radiation)

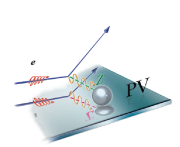




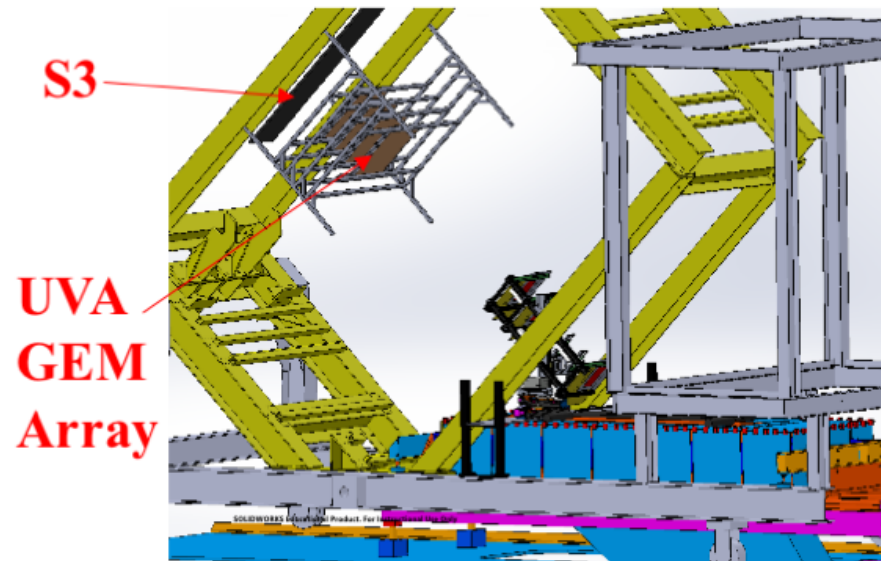
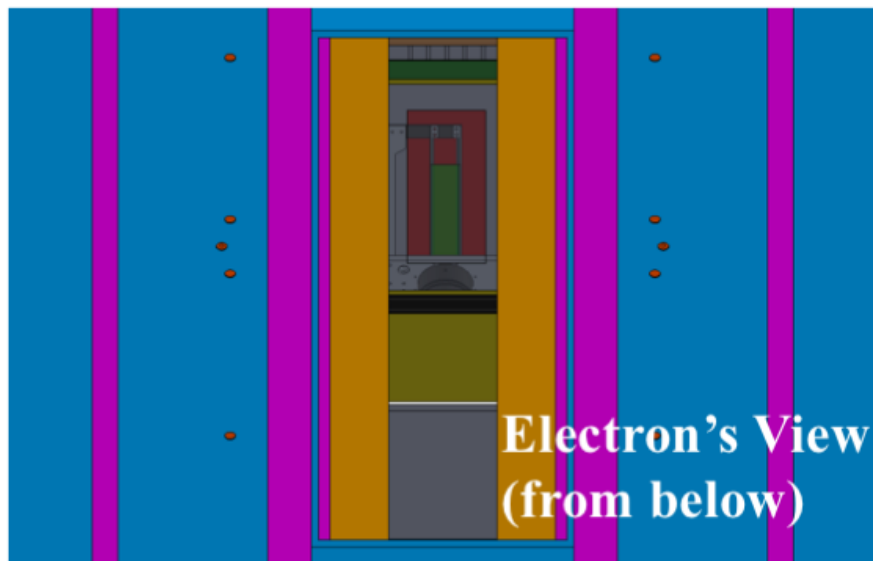
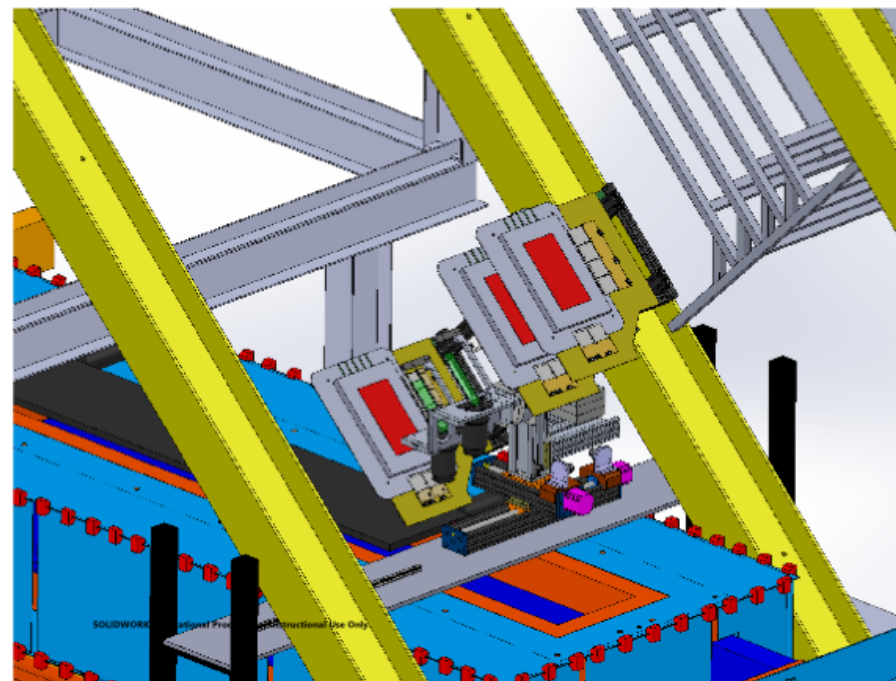
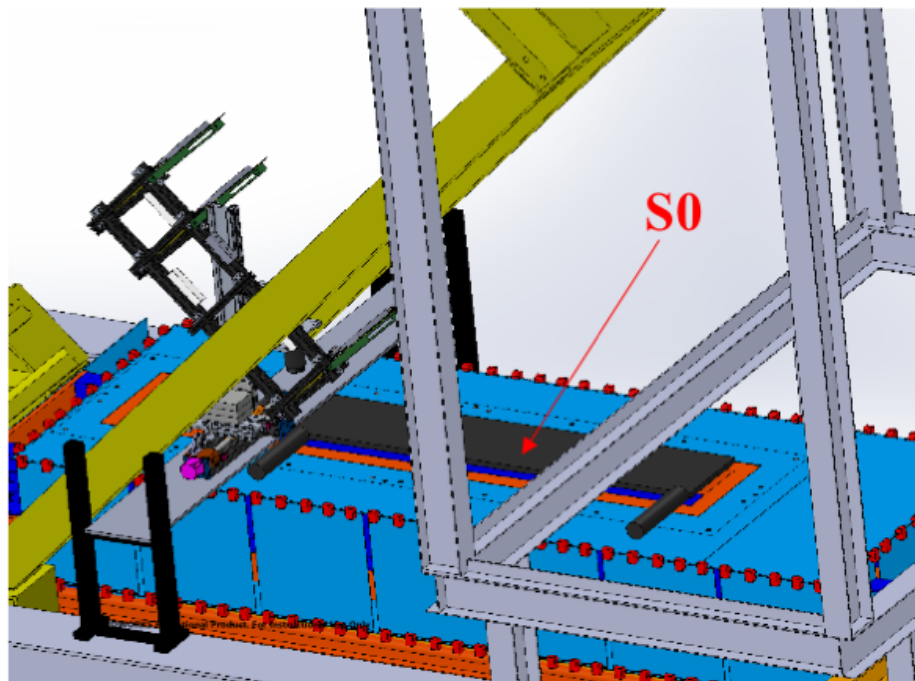
# HRS Detector Package for PREX-II/CREX

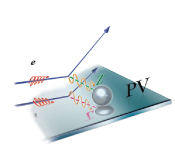
- All HRS standard detector packages removed except for VDCs: No S1, S2, Cerenkov, or Calorimeter
- For counting-mode operation: Use S0 + S3 for triggering
- Additional array of large GEMs from UVA group installed above PREX detector package
- A\_T detector not shown: will mount just above small GEMs
- Plan to reuse same hardware and mounting/installation concept developed for PREX-I



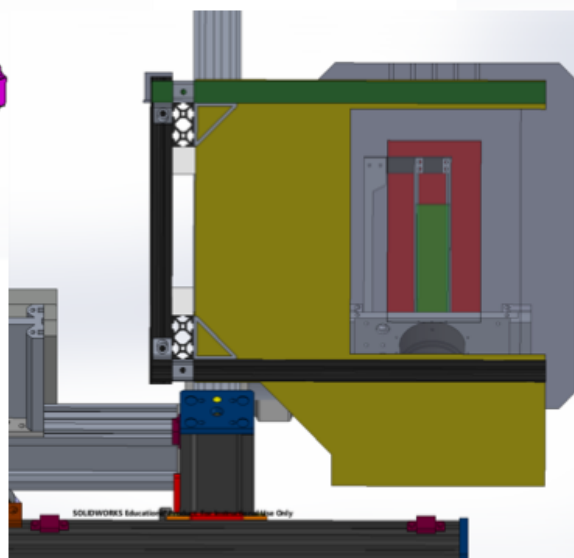
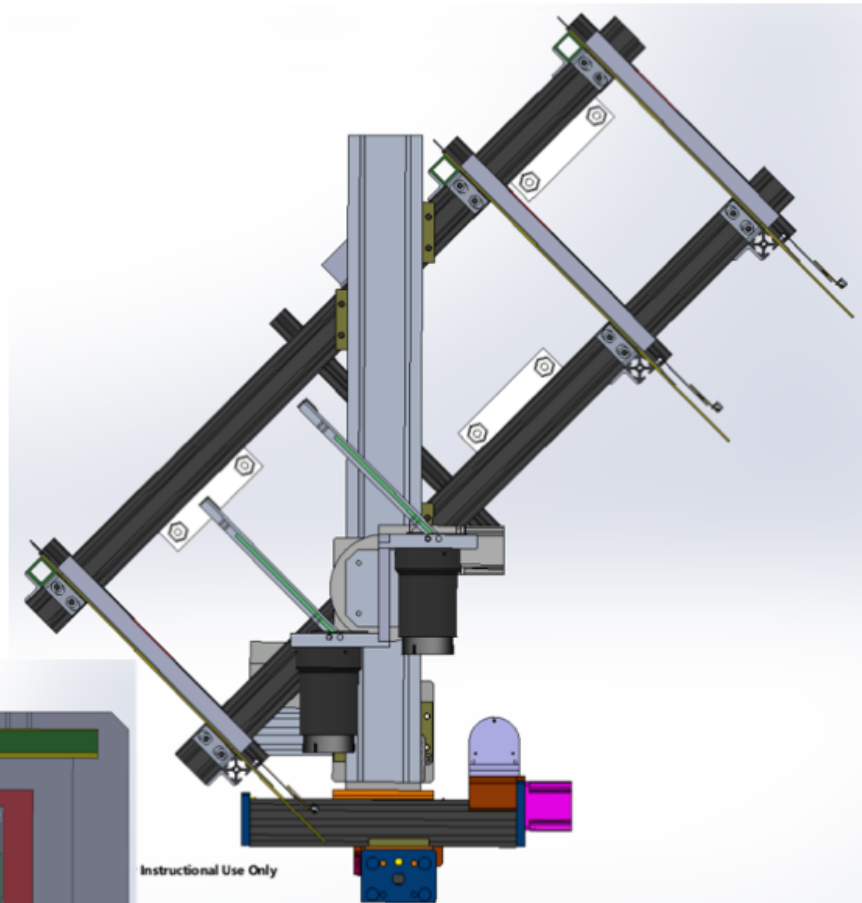
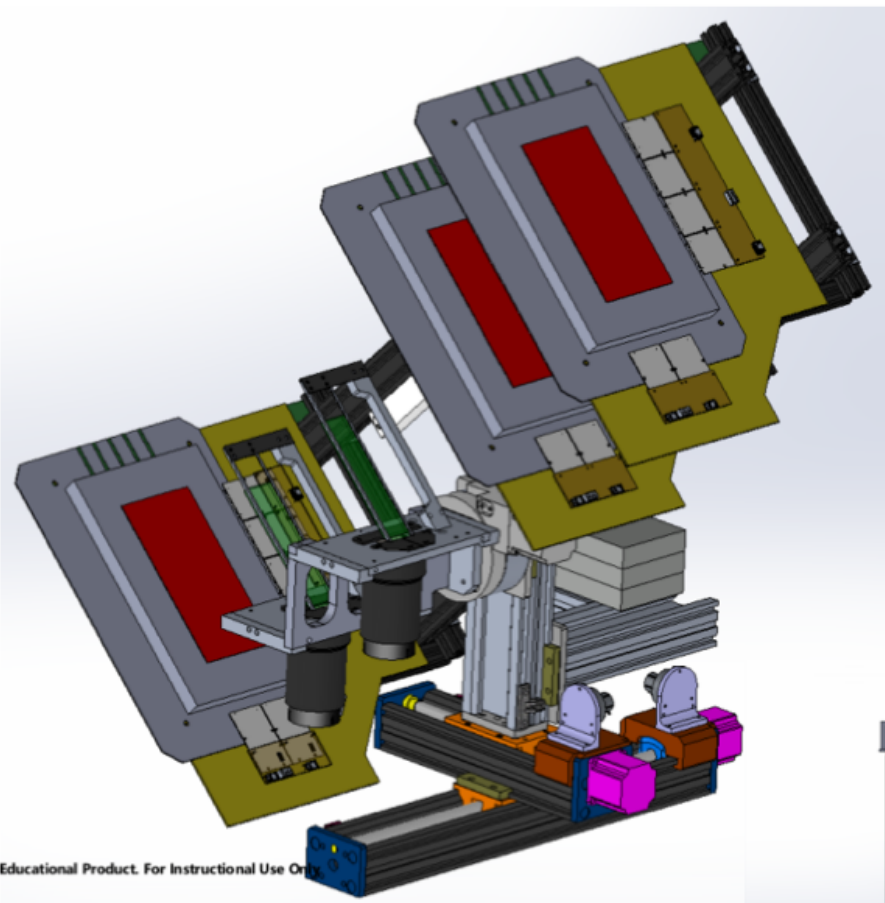


# HRS Detector Package for PREX-II/CREX





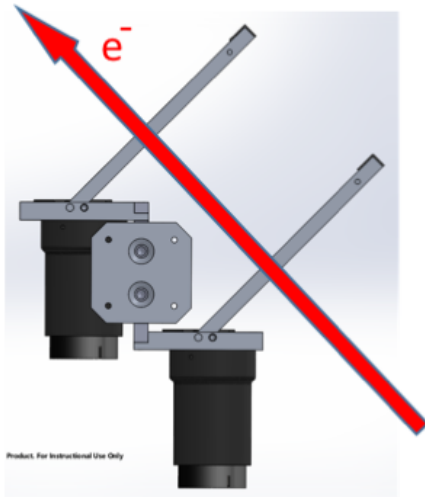
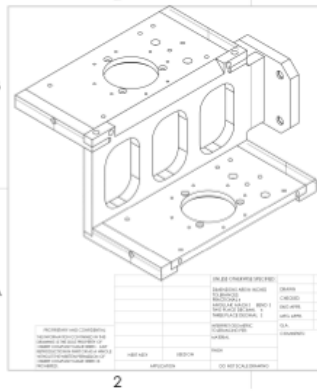
# RHRS Tandem Quartz Mount with GEMs



Electron's view (from below)

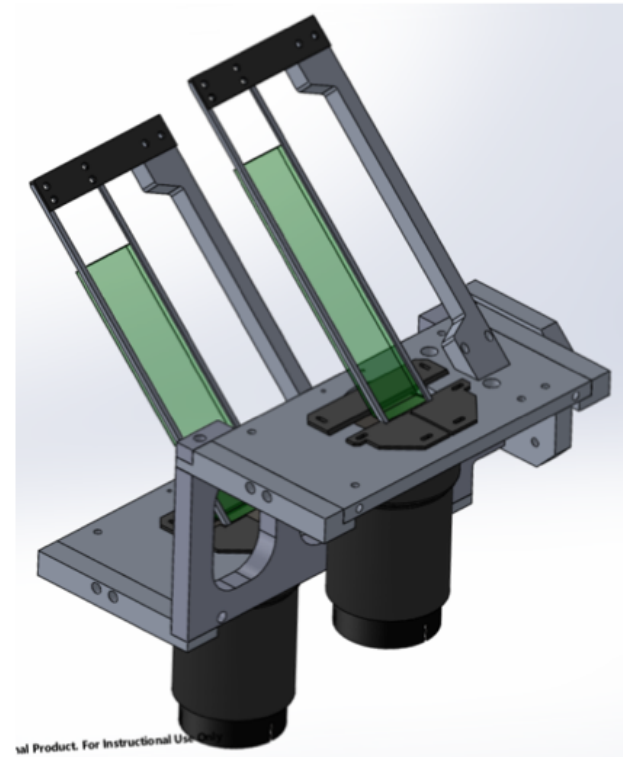
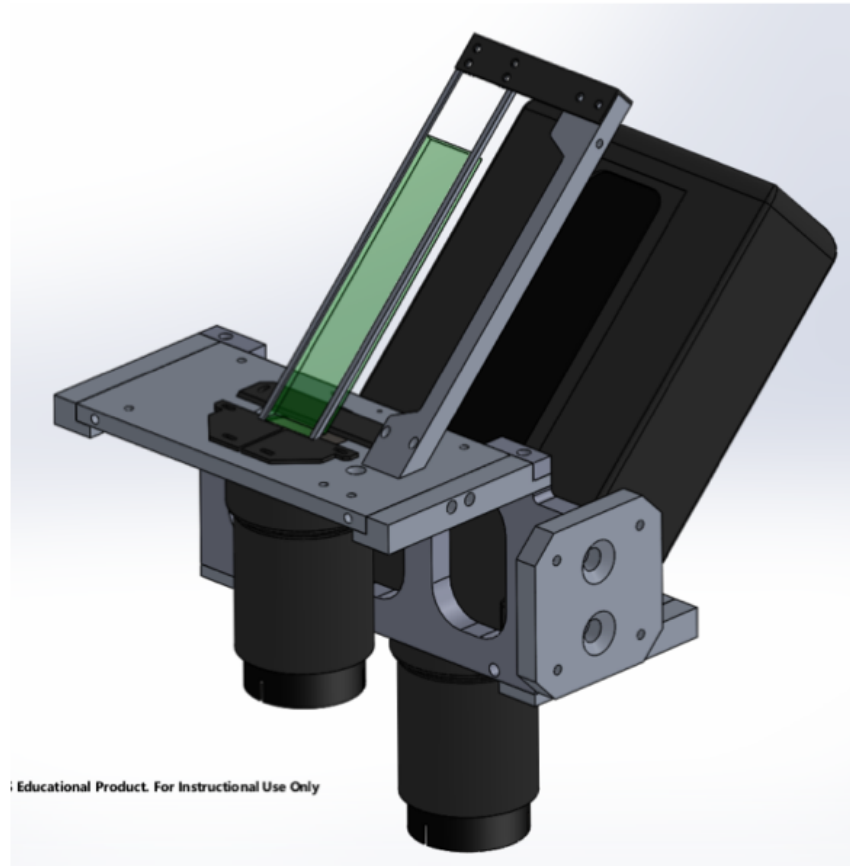


# Main Integrating Tandem Detector Design

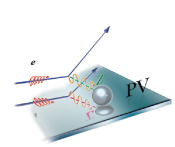



Part No.	Part	Description	Quantity
1	Front Plate	2.000 inch Thick 4013 Aluminum Plate	2
2	Main Extension	5.125 inch Thick 4013 Aluminum Plate	1
3	Lower Left Extension	5/8 x 1.4061 Aluminum Flat	1
4	Lower Right Extension	5/8 x 1.4061 Aluminum Flat	1
5	Upper Left Extension	5/8 x 1.4061 Aluminum Flat	1
6	Upper Right Extension	5/8 x 1.4061 Aluminum Flat	1
7	Base Mount	5.125 inch Thick 4013 Aluminum Plate	1

SEE DWG NO. **Tandem\_FrontPlates**  
SCALE: 1:2 INCHES SHEET 1 OF 12

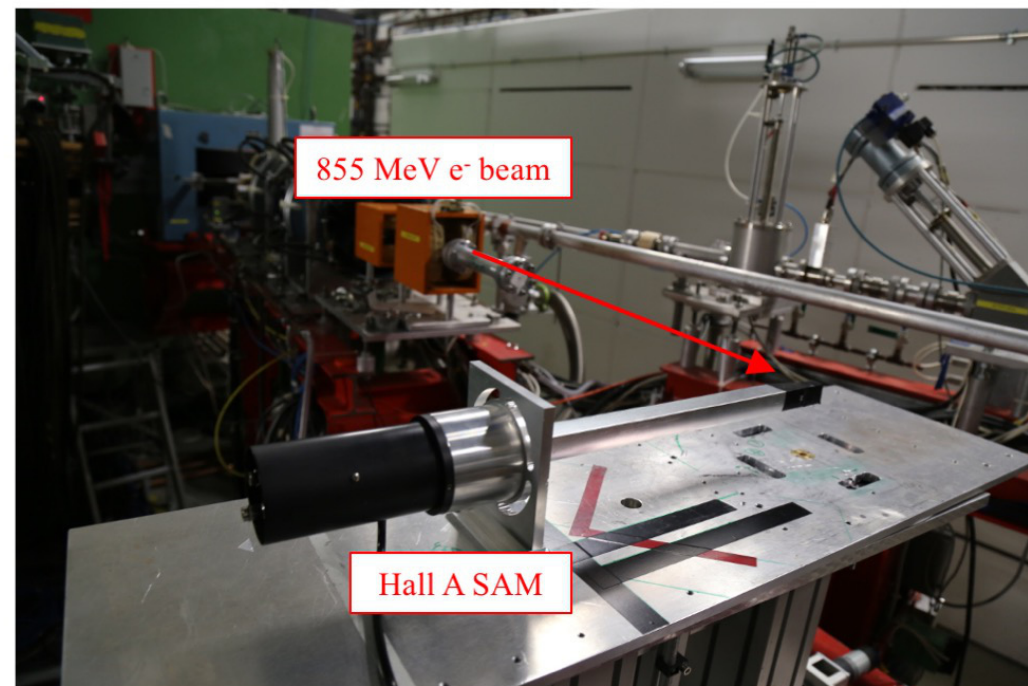
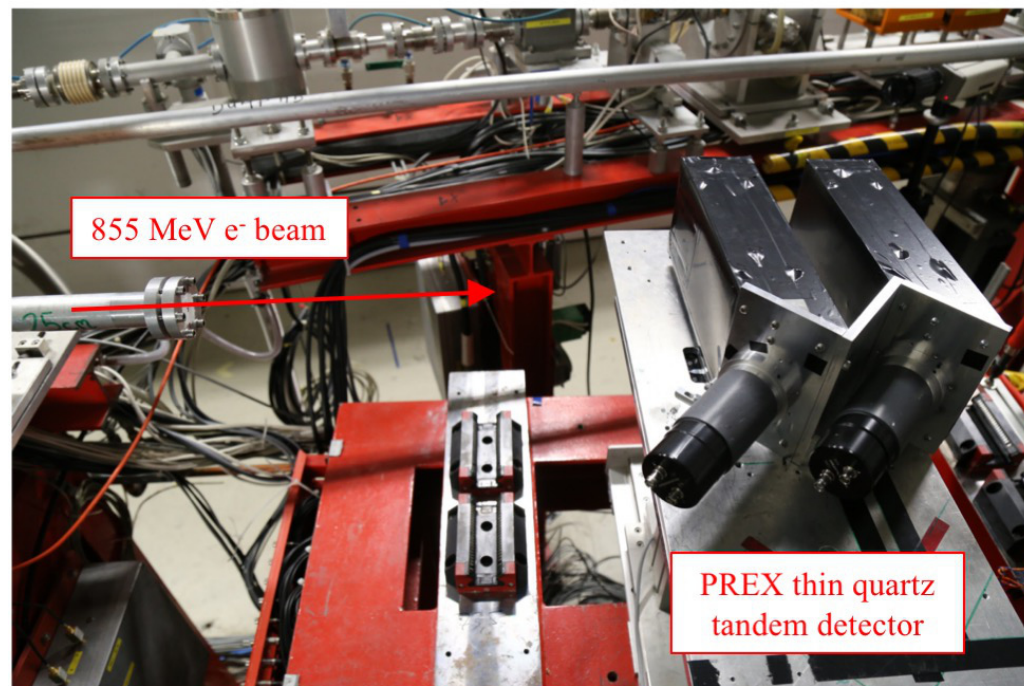


- PREX-II/CREX main detector design based on UMass Design-3.
- Rotatable tandem mount designed and prototype constructed
- New design has shorter quartz rails and incorporates mu-metal shields and 3D printed Nylon enclosure with Kapton windows



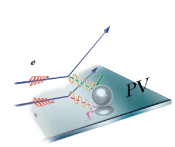
## MAMI testbeam May 24-27, 2016

- $\frac{3}{4}$  shift total for PREX/CREX and SAM tests

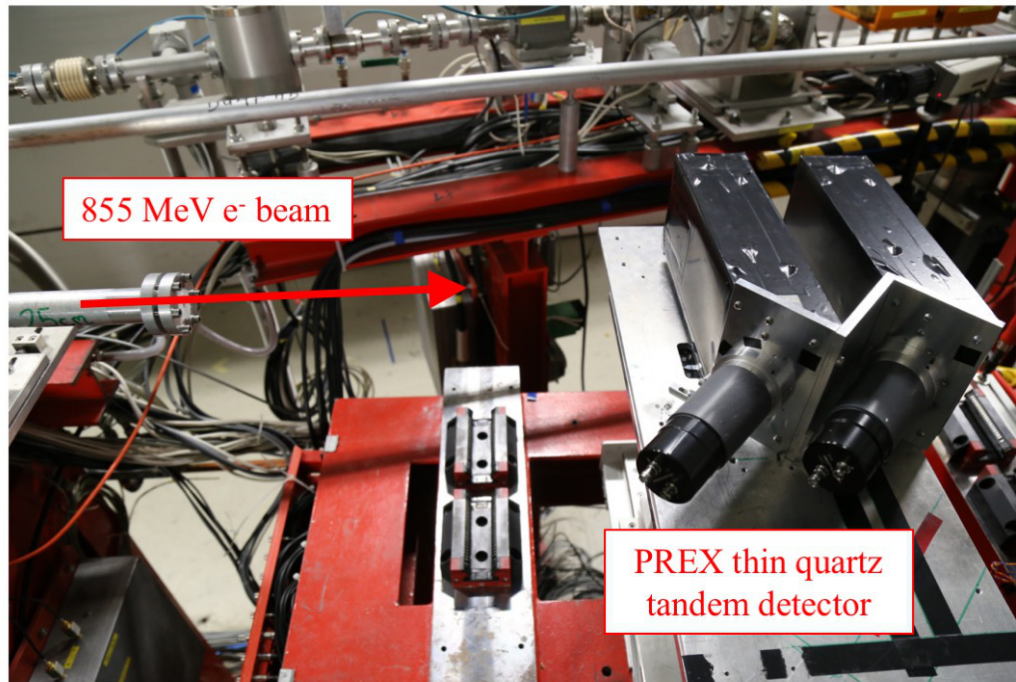


- 6mm and 10mm Tandem mount
- Near normal e<sup>-</sup> incidence
- Final SAM detector PE yield studies:
  - Miro27 and UVS light-guides
  - With and without 1cm tungsten pre-radiator





## PREX/CREX Tandem mount Tests

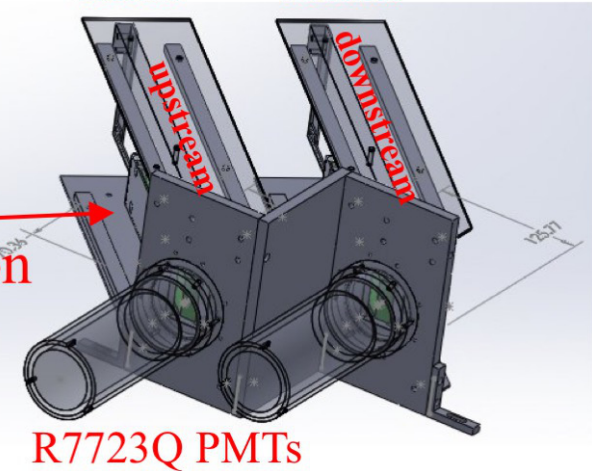


855 MeV e<sup>-</sup> beam

PREX thin quartz tandem detector

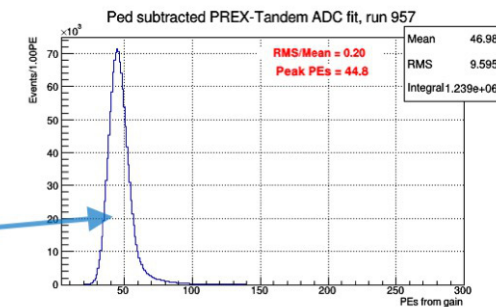
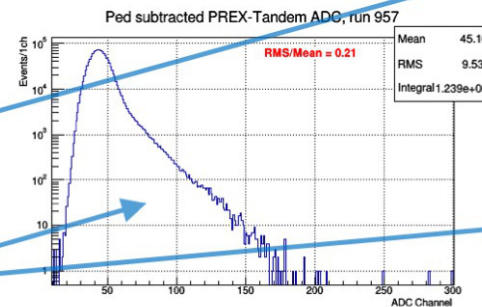
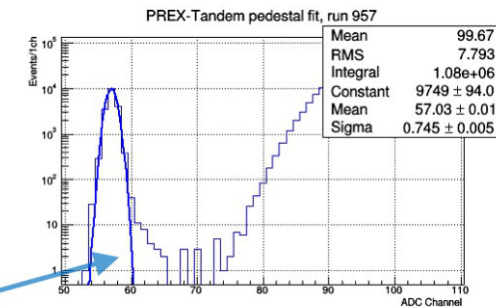
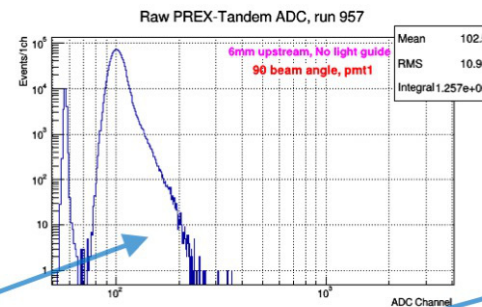
Spectrosil2000 thicknesses: 10mm and 6mm  
6mm and 10mm

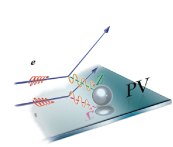
e<sup>-</sup> beam  
Centered on quartz at ~90°



R7723Q PMTs

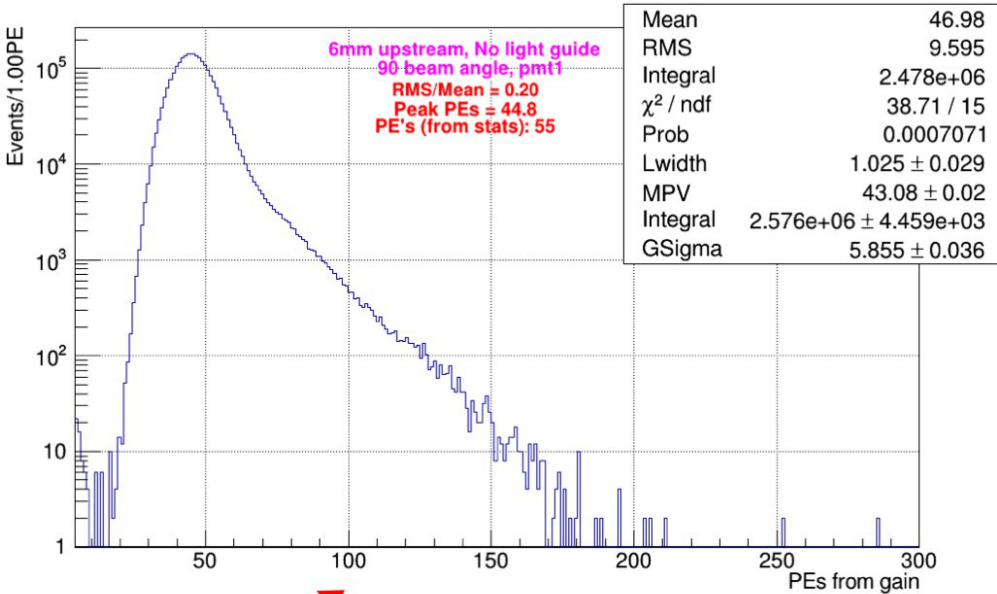
- Quartz spacing same as for rotary tandem mount (~16 cm)
- Used two Hamamatsu R7723Q pmts
- Quartz is wrapped with 1 mil Al. Mylar
- Took runs for each quartz thickness upstream and downstream
- Example raw data, pedestal fit, and ped-corrected ADC and PE dists



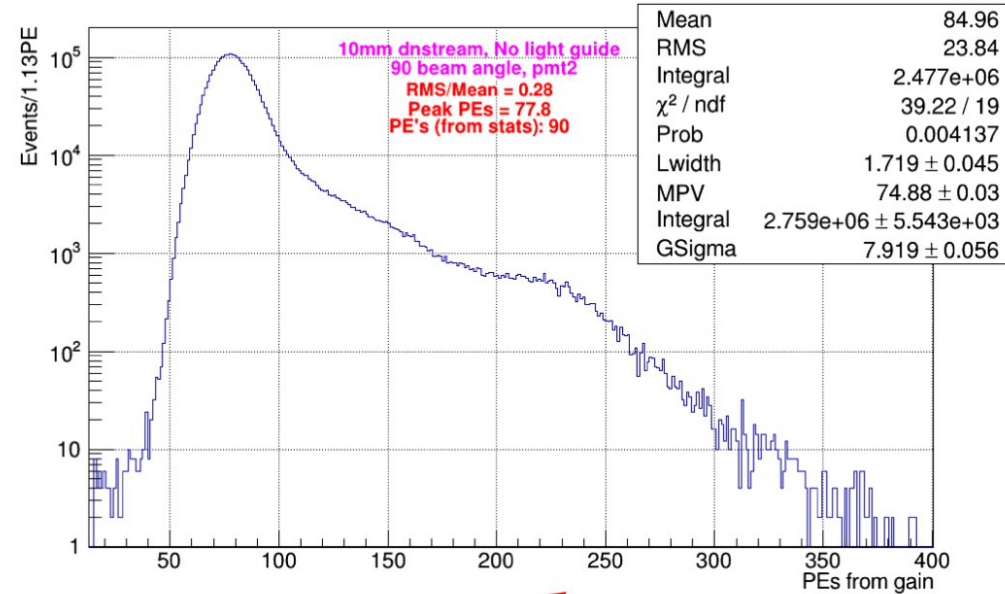


## 6 mm/10 mm Tandem Testbeam Results

Ped subtracted PREX-Tandem ADC fit, run 957



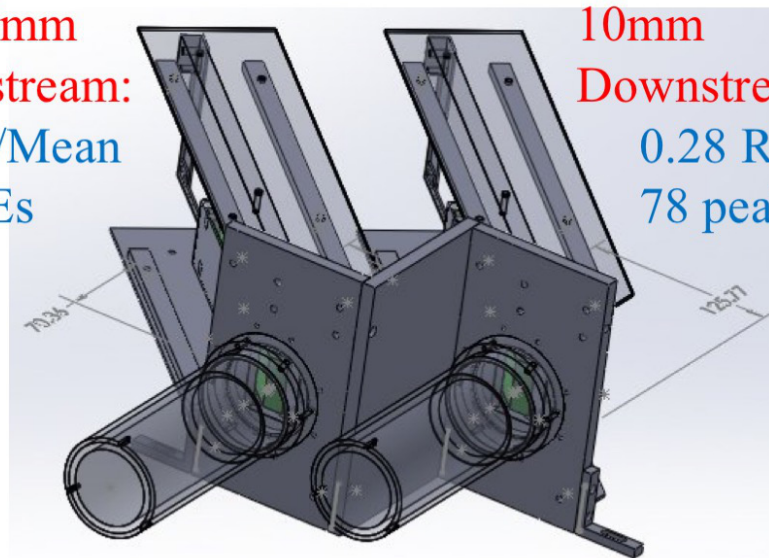
Ped subtracted PREX-Tandem ADC fit, run 957



- PE's converted from ADC units using PMT gains
- Peak PE's from Langau fit parameters did not agree with PE's from gain; fits were poor and very sensitive to fit domain around peak

6mm  
Upstream:  
0.20 RMS/Mean  
45 peak PE's

10mm  
Downstream:  
0.28 RMS/Mean  
78 peak PE's

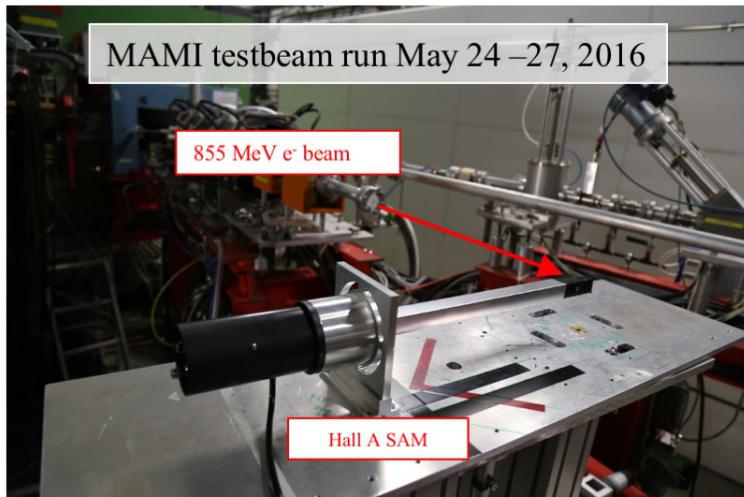


- Uncertainty in PMT gains between 5 – 10%

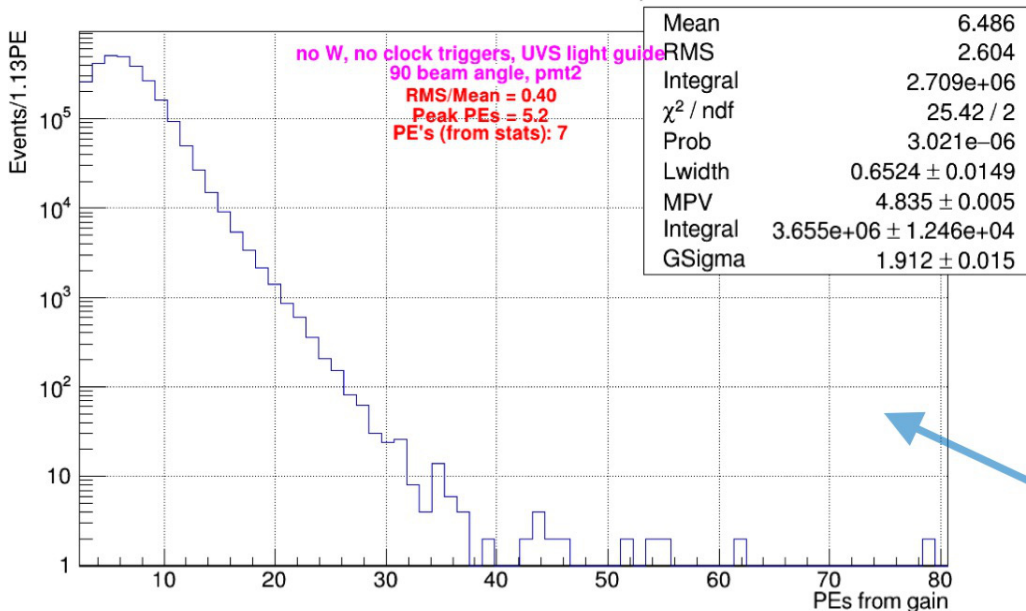


## SAM PE Yield & LG Testbeam Study: Miro-silver27 vs Anolux UVS (no tungsten)

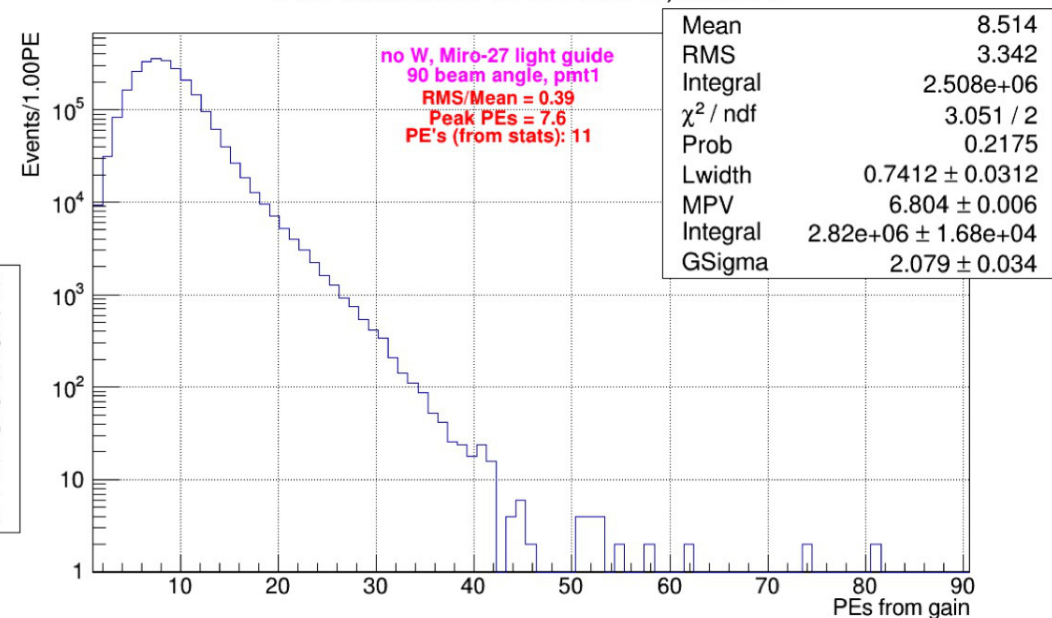
- Miro-silver27, no tungsten, N<sub>2</sub> gas flowing:  
~7 - 8 peak PEs (using PMT gain) with 39% relative width



Ped subtracted SAM ADC fit, run 1103



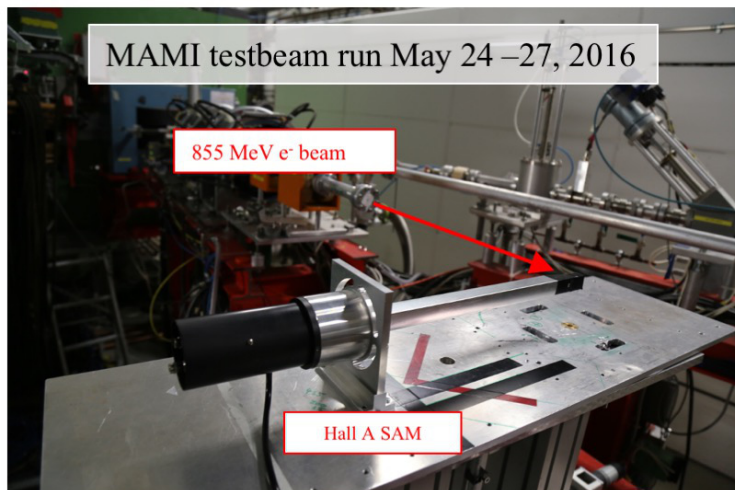
Ped subtracted SAM ADC fit, run 976



- Anolux UVS, no tungsten, N<sub>2</sub> gas flowing, and no clock-triggers: ~5 peak PEs (using PMT gain) with 40% relative width



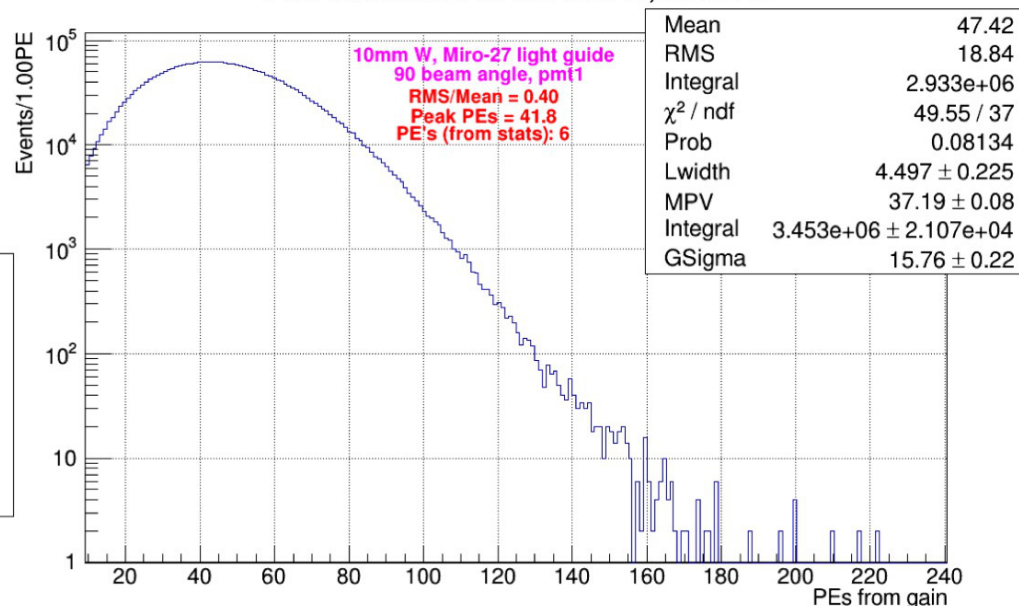
## SAM PE Yield & LG Testbeam Study: Miro-silver27 vs Anolux UVS (w/ tungsten)



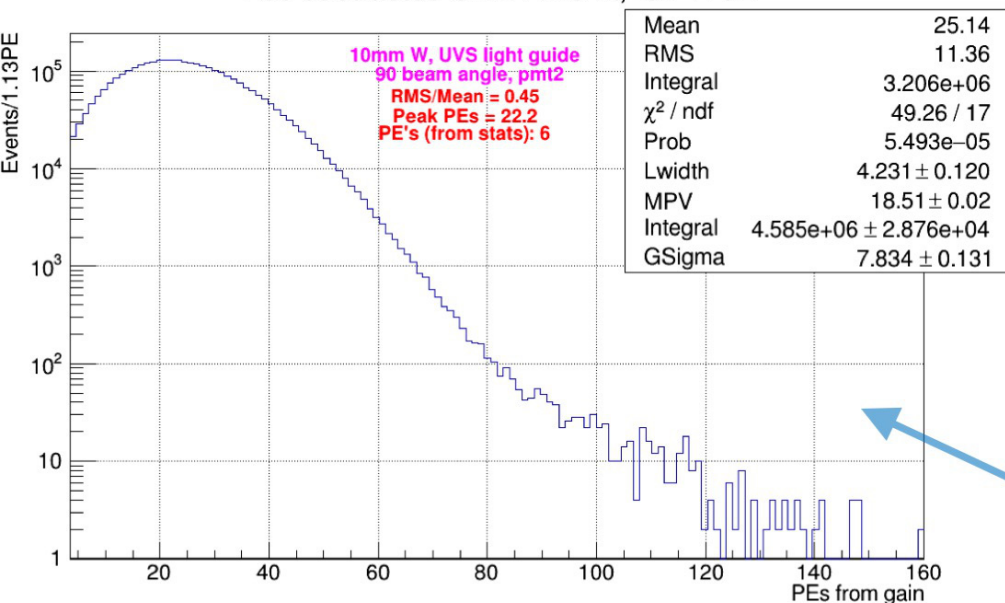
MAMI testbeam run May 24 -27, 2016

- Miro-silver27, w/ 10mm tungsten, N<sub>2</sub> gas flowing: **~42 peak PEs** (using PMT gain) **with 40% relative width**

Ped subtracted SAM ADC fit, run 978



Ped subtracted SAM ADC fit, run 1102



- Anolux UVS, w/ tungsten, N<sub>2</sub> gas flowing: **~22 peak PEs** (using PMT gain) **with 45% relative width**



## Final SAM Design and 2016 Testbeam

MAMI testbeam run May 24 –27, 2016

855 MeV e<sup>-</sup> beam

Hall A SAM

- Final SAM detector PE yield studies:
  - MiroSilver27 and UVS light-guides
  - With and without 1cm tungsten pre-radiator



Small Ange Monitors:  
Detect ~0.5° target scattering

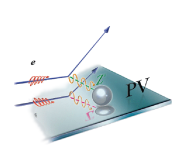


Assembled & Installed in Hall A Fall 2015

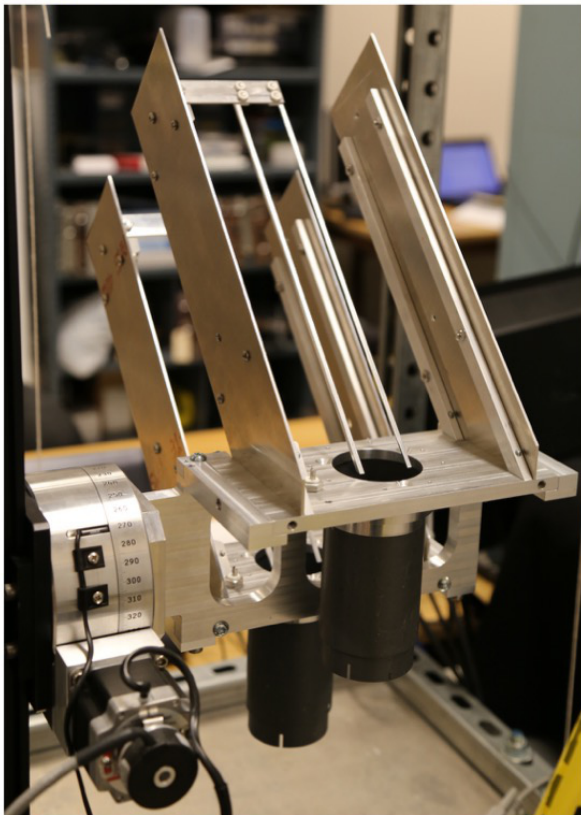
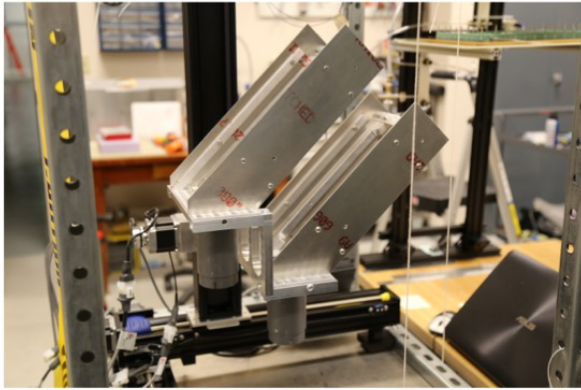


### Final SAM detector

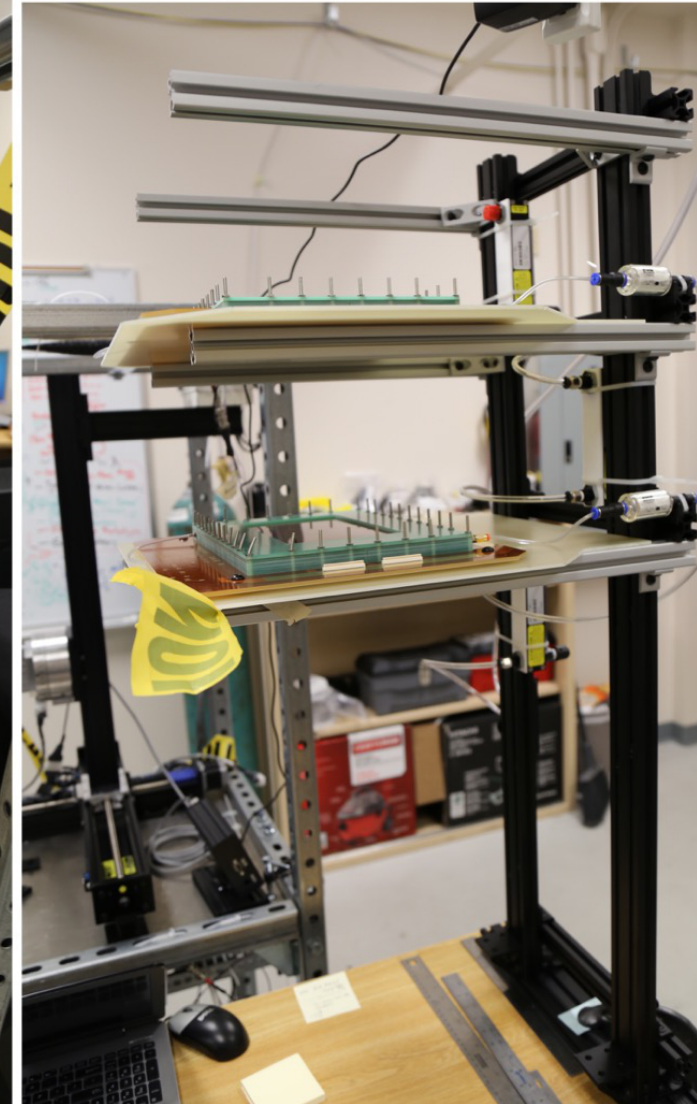
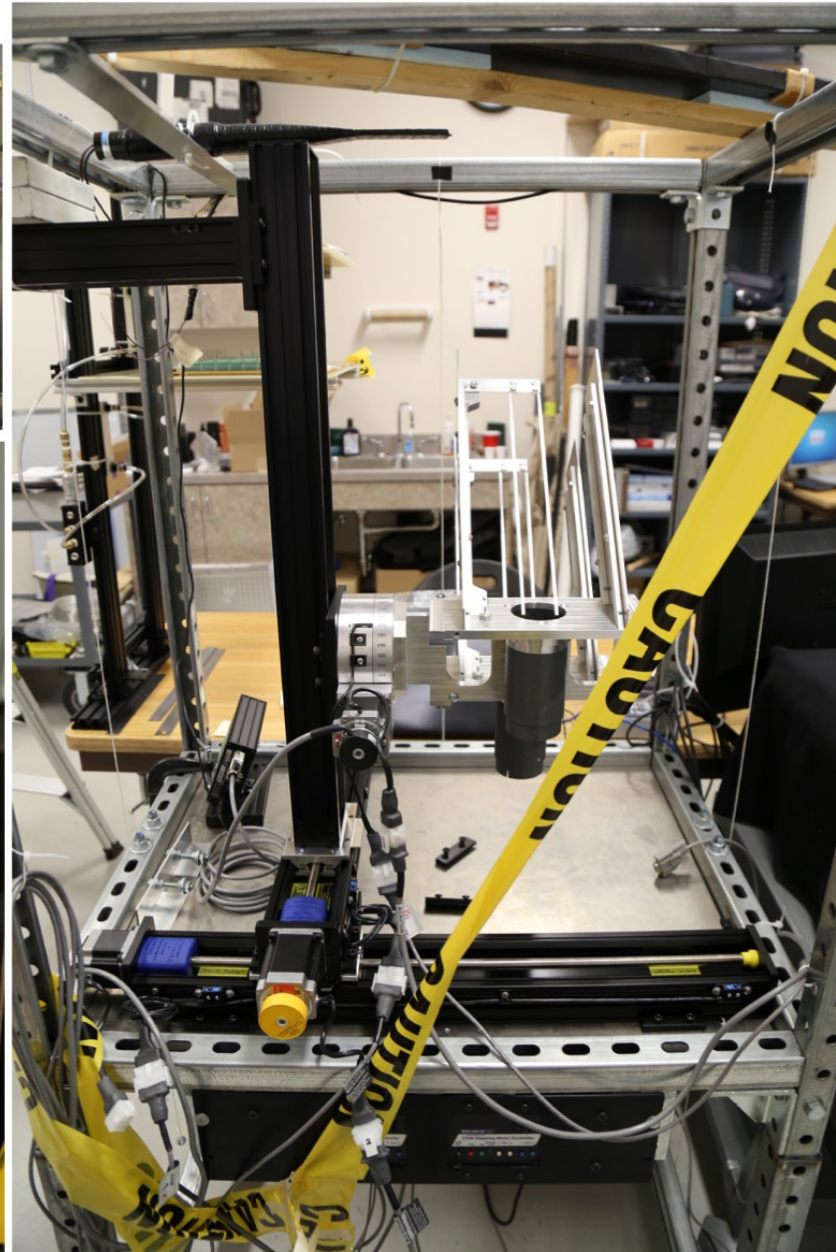
- Quartz: 33 x 20 x 13 mm<sup>3</sup>
- Miro27 LG: 36 x 2.6 x 2.1 cm<sup>3</sup>
- Optimized 1-bounce funnel mirror
- Unity or high-gain R375 2" PMTs
- Use of pre-radiator not decided
- Dry-air inlet and outlet ports
- Custom flange adapter for easy de-install/re-install (radcon permitting)



# Prototype Development at ISU



Prototype LHRs Tandem mount

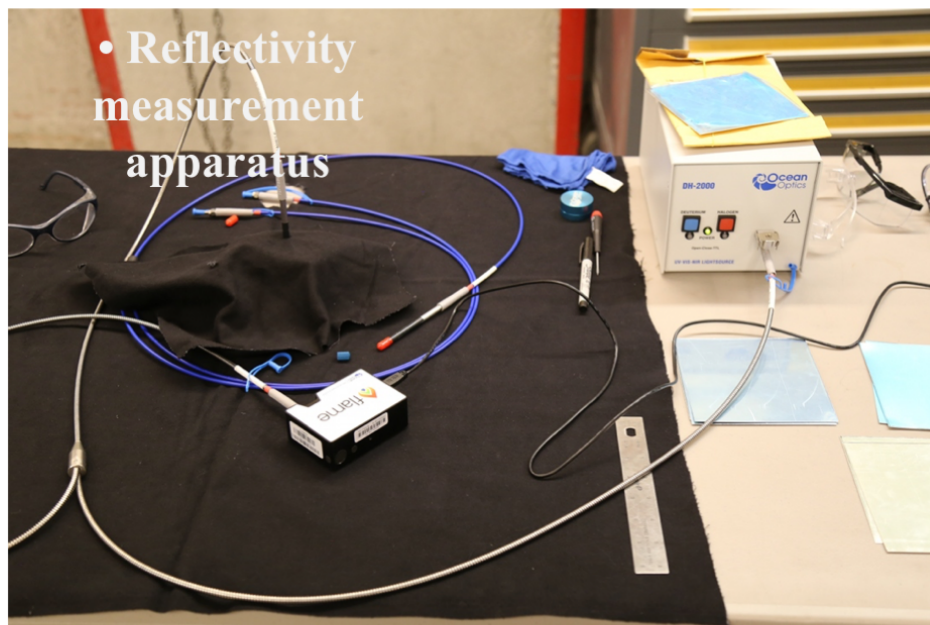


Prototype GEM stand  
(under construction)



## Light guide reflectivity measurements

- Measuring light guide (LG) reflectivity as function of angle (10 – 90°) and  $\lambda$  (200 – 800nm); ongoing



- Light source: Ocean Optics DH2000: 200 - 800nm, 25W Deuterium bulb
- Spectrometer: Ocean Optics USB Flame, enhanced sensitivity, UV-VIS grating
- NIST specular calibration standard

Light guide materials tested:

Miro-silver 4270

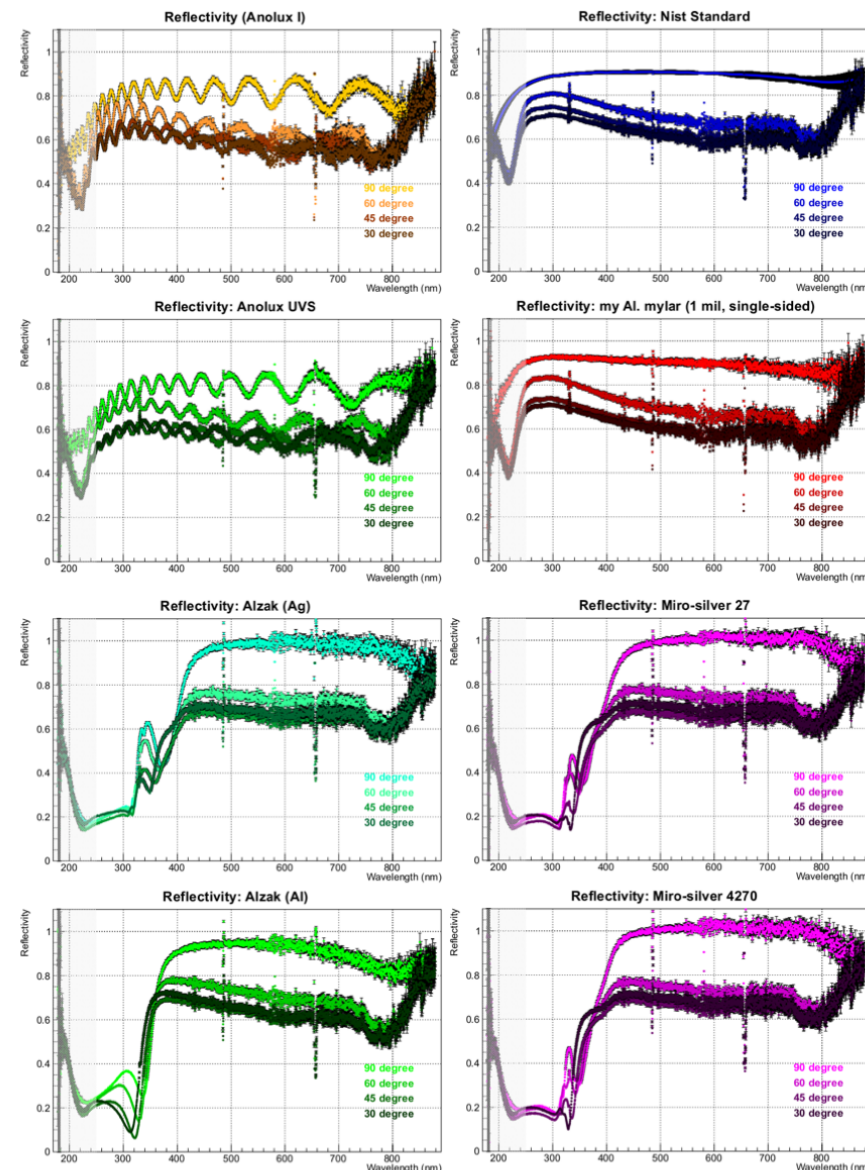
Miro-silver 27

Anolux I and UVS

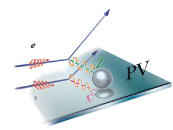
Alzak-Al and Alzak-Ag

Miro 2000Ag (diffuse)

1 mil, single-sided aluminized mylar



Reflectivity vs.  $\lambda$  for various materials at diff. angles



## LG reflectivity radiation hardness study

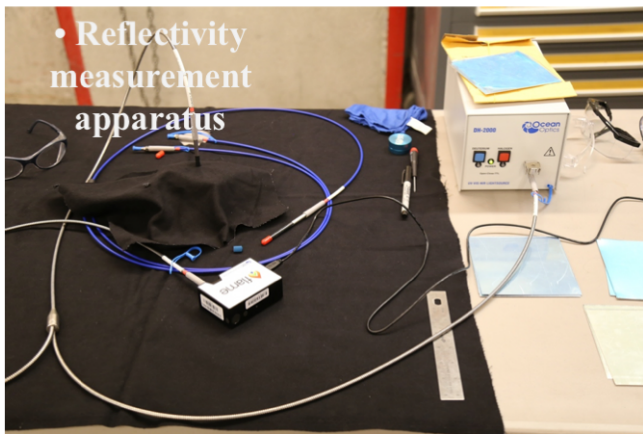
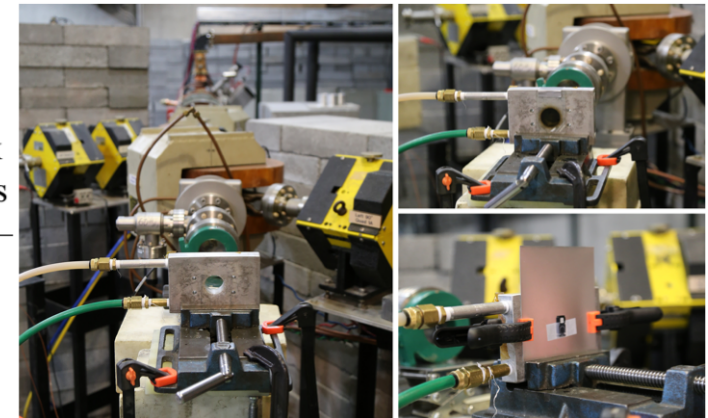
**25 MeV LINAC (Main Hall and Airport)**

RF Frequency: 2856 MHz (S-Band)  
 Energy Range: ~4~25 MeV (current varies)  
 Pulse Width: ~50ns to 4 micro seconds  
 Repetition Rate: single pulse to 360 Hz  
 Ports: 0 degree, 45 degree and 90 degree (Beam energy resolution ~ 1+/- 15%)

25B Energy vs Current			
Energy (MeV)	0 port (mA)	45 port (mA)	90 port (mA)
23	55	55 @ 3.8uS	46 @ 3.6 uS
20	100	70 @ 4. uS	65 @ 4. uS
16	100	48 @ 3.6 uS	48 @ 3.6 uS
13	80	30 @ 3.3 uS	15 @ 3.3 uS
10	60	18 @ 3.1 uS	7.5 @ 3.1 uS
9	110	30 @ 4. uS	15 @ 4. uS
6	100	60 @ 4. uS	60 @ 4. uS
4	50	20 @ 4. uS	20 @ 4. uS

- Used 8 MeV e<sup>-</sup> beam, 65 - 110mA I<sub>peak</sub>, 4μs pulse width at 250 Hz, 310 – 880 W

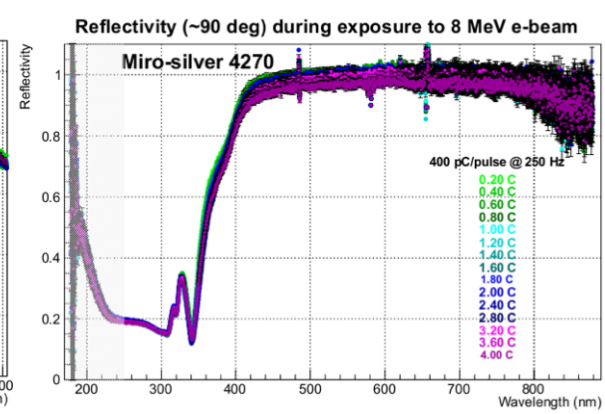
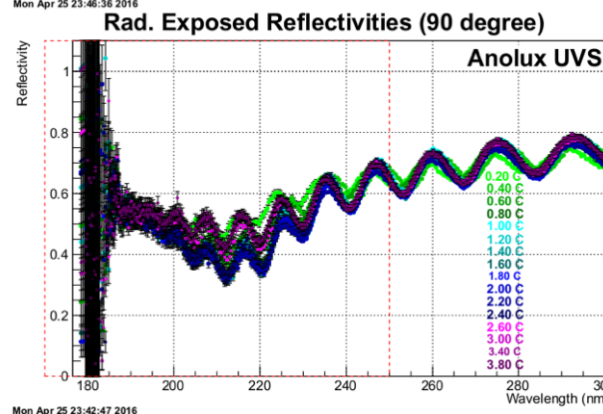
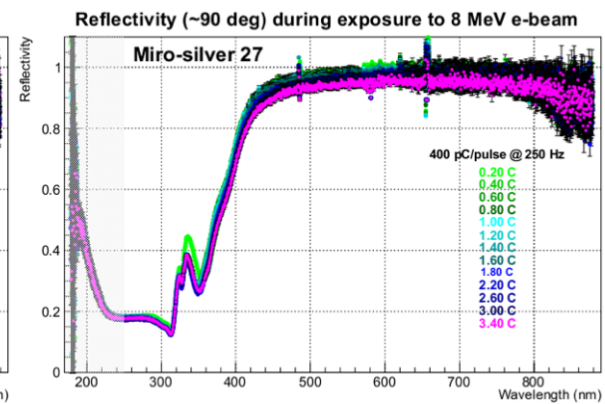
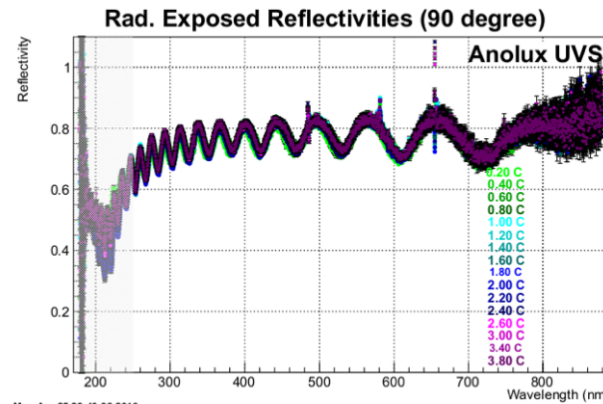
- Water-cooled (15° C) aluminum brick w/ 1.5 cm radius hole (for beam) – more than adequate cooling.



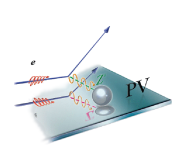
• Reflectivity measurement apparatus

Irradiated several light guide material samples over a 3 day period from Mar 22 - 24, 2016:

- Miro-silver 4270
- Anolux UVS
- Miro 2000Ag (diffuse)
- Miro-silver 27 (from Michael)
- Alzak-Al and Alzak-Ag (from KK)
- 1 mil, single-sided aluminized mylar





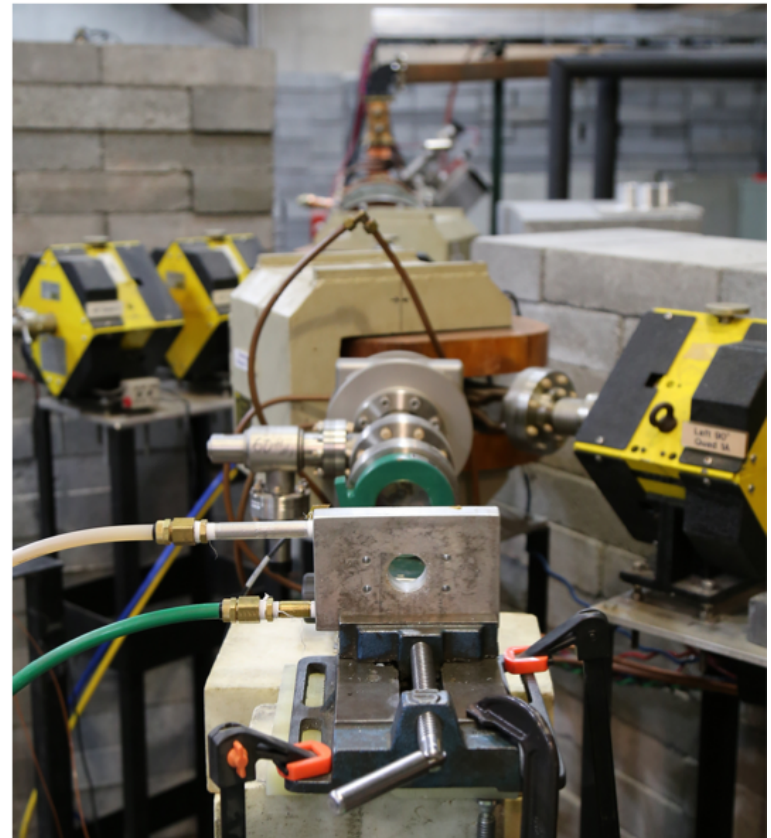
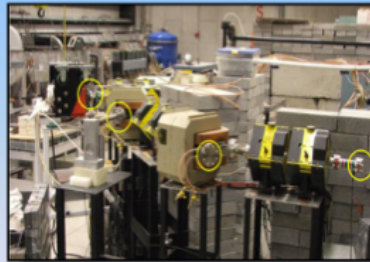


## Radiation Hardness QC for quartz and other components

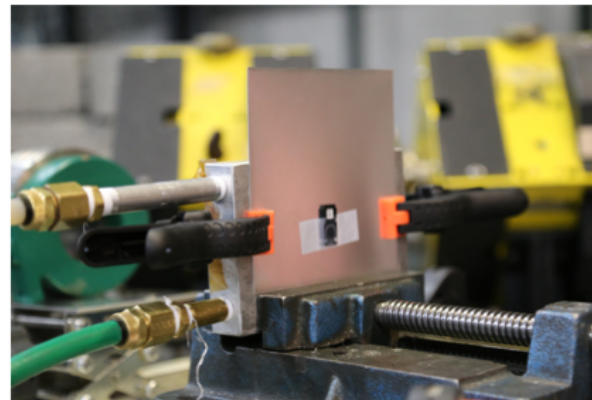
### 25 MeV LINAC (Main Hall and Airport)

**RF Frequency:** 2856 MHz (S-Band)  
**Energy Range:** ~4~25 MeV (current varies)  
**Pulse Width:** ~50ns to 4 micro seconds  
**Repetition Rate:** single pulse to 360 Hz  
**Ports:** 0 degree, 45 degree and 90 degree (Beam energy resolution ~ +/- 15%)

25B Energy vs Current			
Energy (MeV)	0 port (mA)	45 port (mA)	90 port (mA)
23	55	55 @ 3.8uS	46 @ 3.6 uS
20	100	70 @ 4 uS	65 @ 4 uS
16	100	48 @ 3.6 uS	48 @ 3.6 uS
13	80	30 @ 3.3 uS	15 @ 3.3uS
10	60	18 @ 3 uS	7.5 @ 3 uS
9	110	30 @ 4uS	15 @ 4 uS
6	100	60 @ 4 uS	60 @ 4 uS
4	50	20 @ 4 uS	20 @ 4 uS



- A key issue is how well can we calibrate dose exposure?
- Another issue is how low can we go in beam current (while still monitoring it)

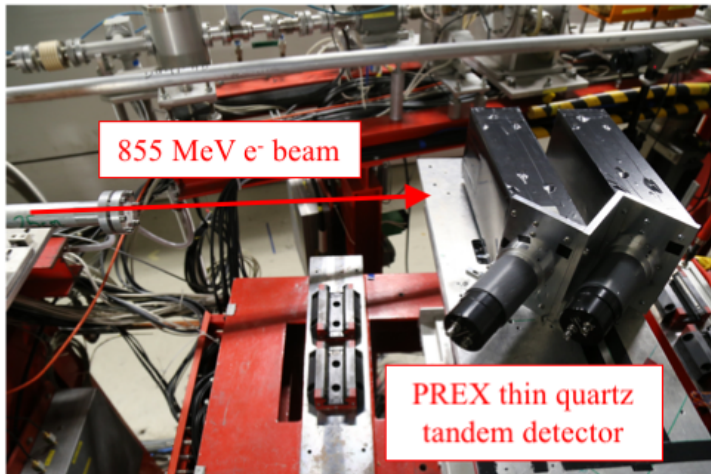


- Planning for a 1 - 2 day engineering run late spring or early summer to address these questions.

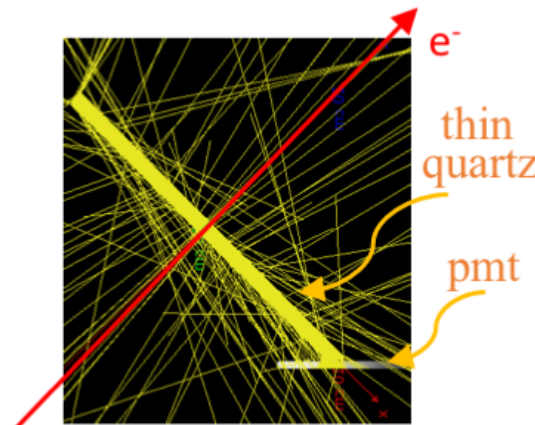


## Monte Carlo tuning and Shower-max Simulations

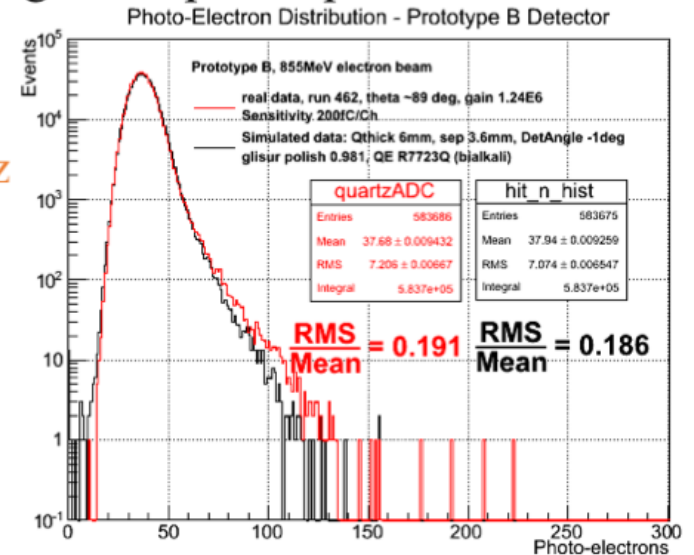
- Quartz optical G4 properties benchmarked at MAMI: Glisur ground polish parameter  $\sim 0.981$



MAMI testbeam with PREX detector



G4 event visualization for PREX detector



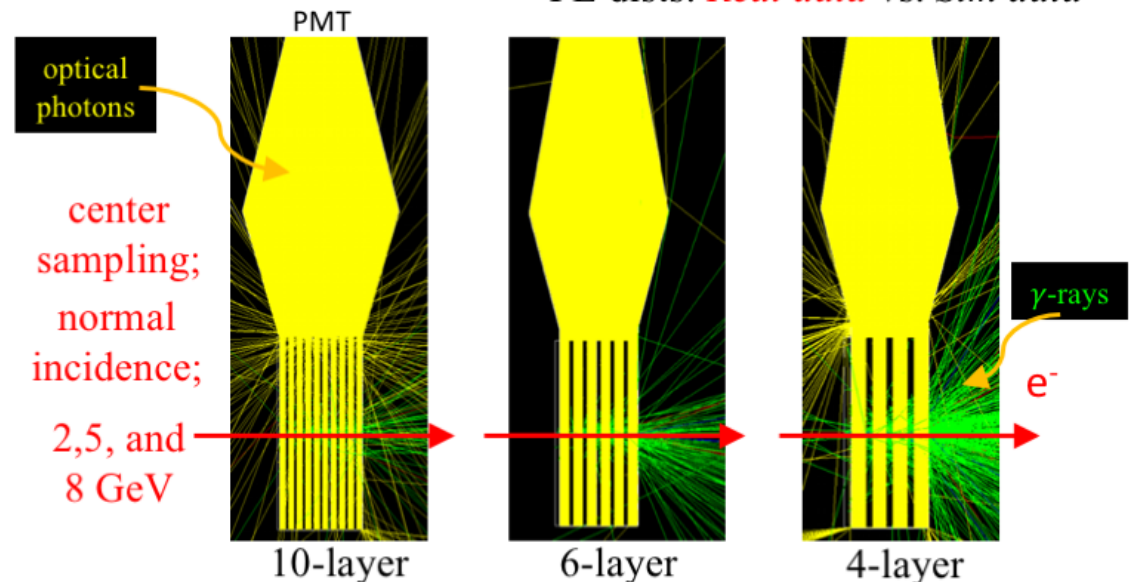
PE dists: *Real data* vs. *Sim data*

- Stack configuration MC study:

- ❖ Stack thicknesses all same ( $7.2 X_0$ )
- ❖ 2, 5, and 8 GeV incident electrons
- ❖ PE dists generated using tuned polish parameter and 60% LG reflectivity

### Conclusion:

4-layer gives comparable performance to 10-layer (and is easier and cheaper to build)



Shower-max event visualizations



# The MOLLER Project at Jefferson Lab:

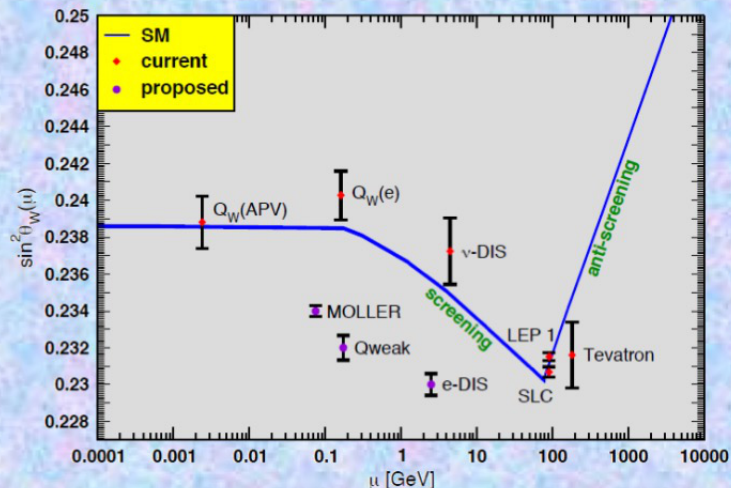
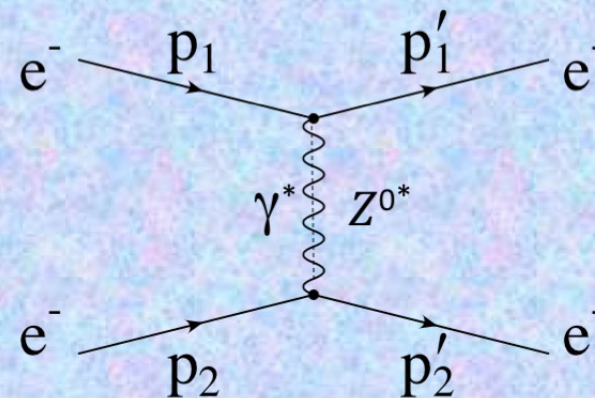
## Measurement of

## Lepton

## Lepton

## Electroweak

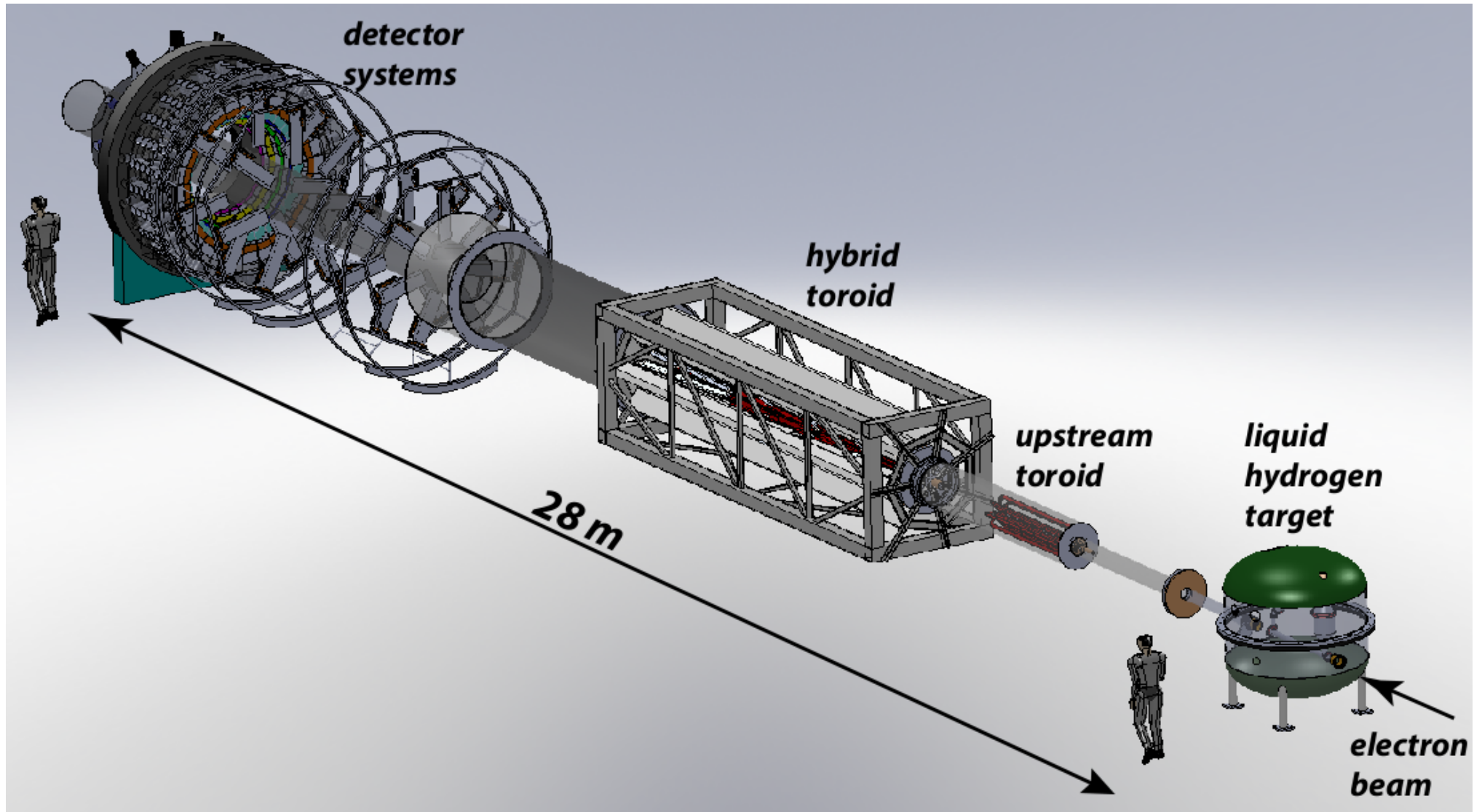
## Reaction





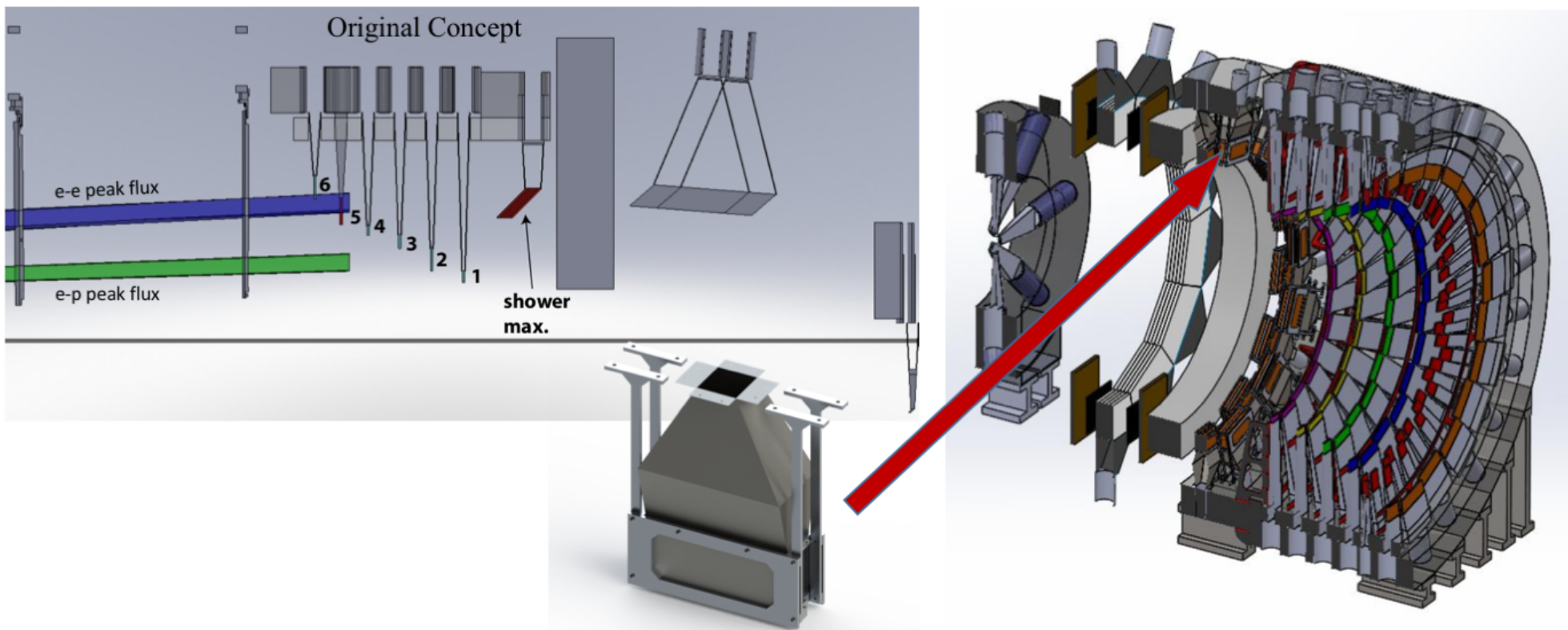
# MOLLER Apparatus

(major new installation experiment for Hall A)

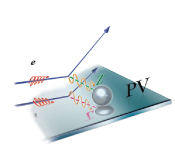




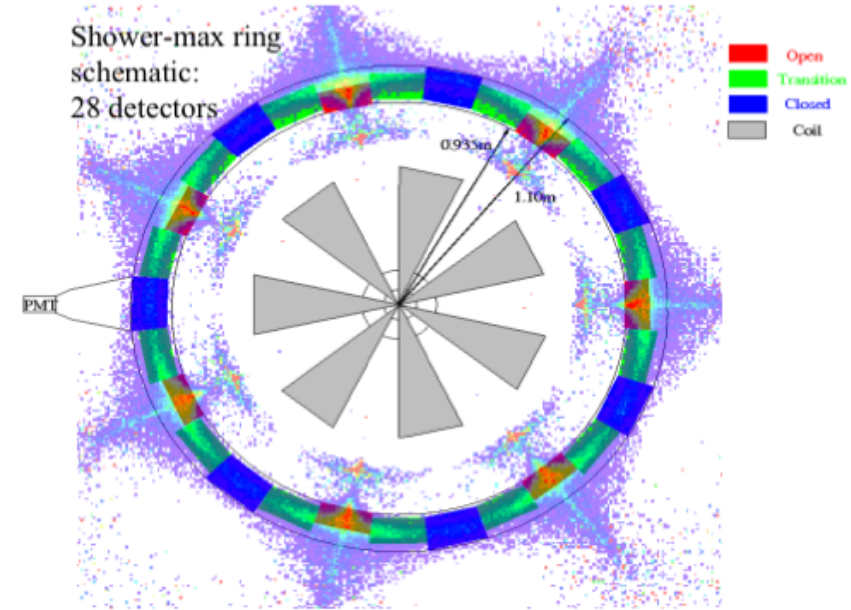
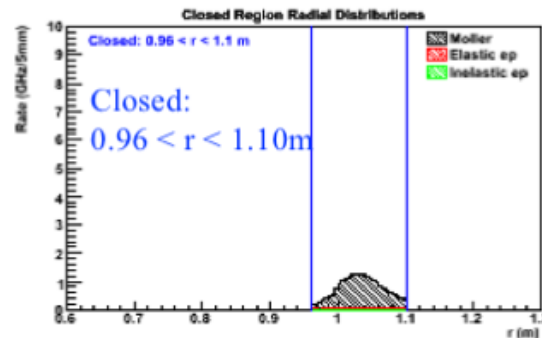
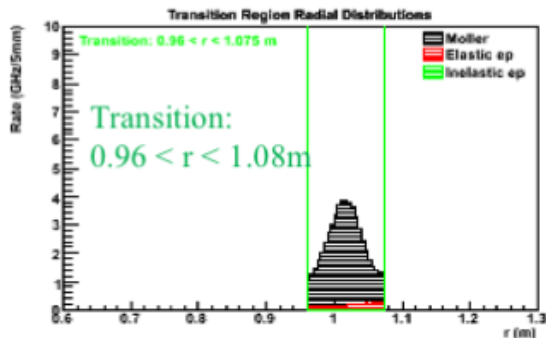
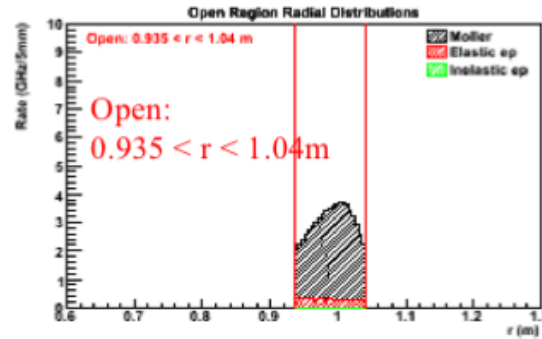
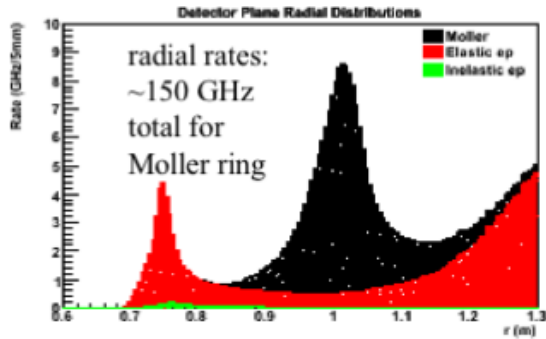
## ShowerMax Motivation & Requirements



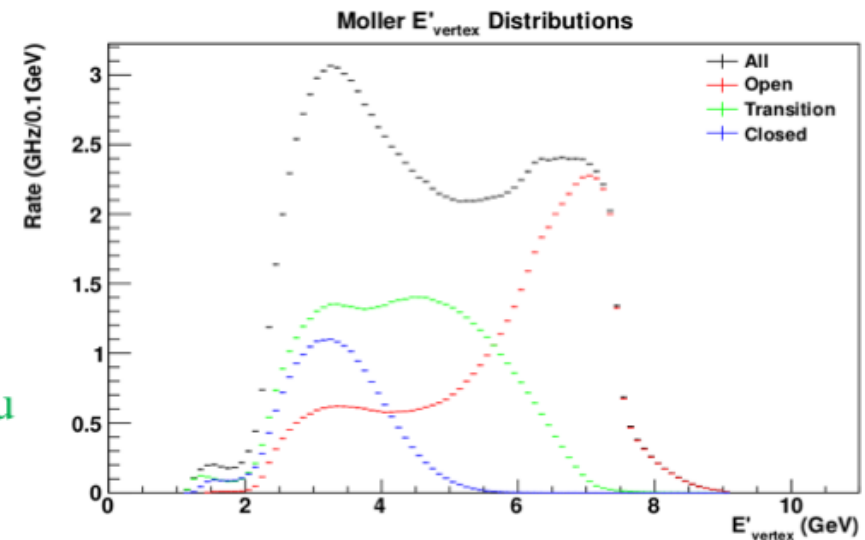
- Provides additional measurement of e-e ring integrated flux
- Weights flux by energy  $\Rightarrow$  less sensitive to low energy and hadronic backgrounds
- Will also operate in tracking mode to give additional handle on background (pion) identification – gives MIP-like signal
- Should have good resolution over full energy range ( $\frac{\sigma}{\langle n \rangle} \lesssim 25\%$ ), long term stability and be radiation hard



## Shower-max phi-segmentation, rates and energies



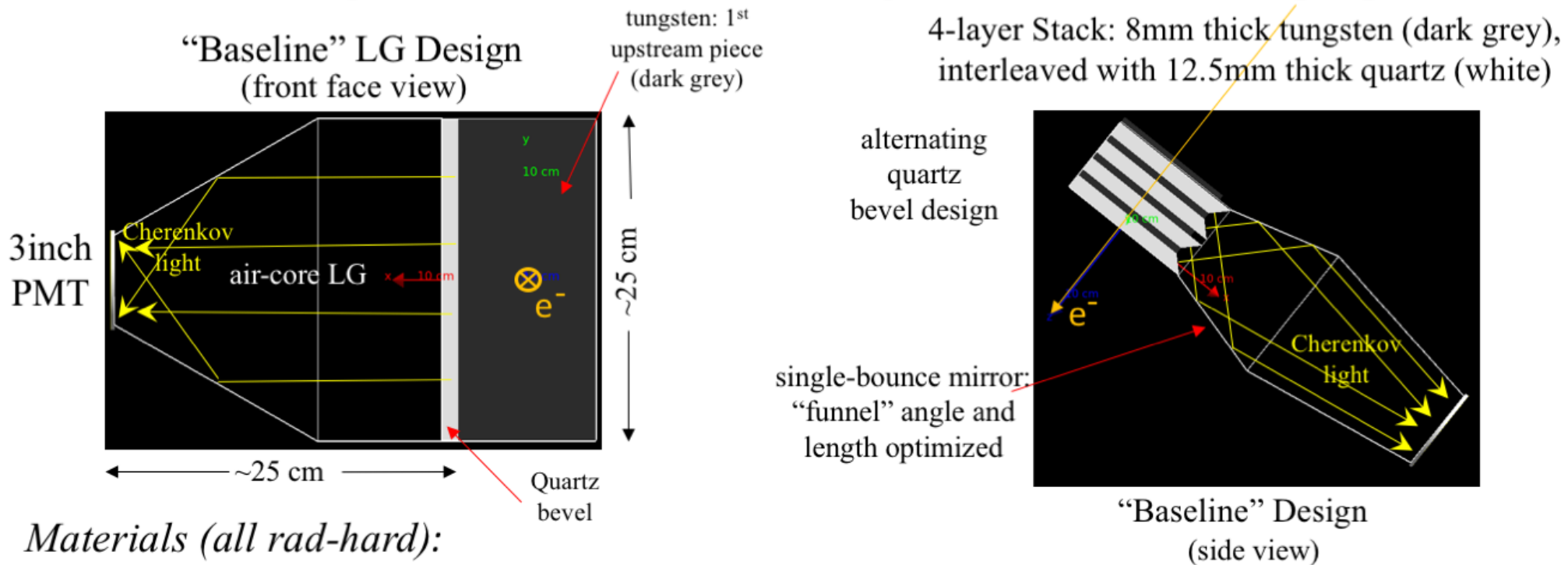
- Large range of rates and energies for different phi-region detectors:
  - Open ~9 GHz/det; 2 - 9 GeV, peak at 7 GeV...
  - Closed ~3.5 GHz/det; 2 - 5 GeV, peak at ~3 GeV
  - Transition ~4.5 GHz/det; 2 - 7 GeV, 3 - 5 GeV plateau





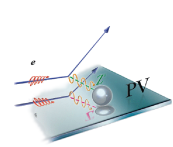
## Baseline Design Stack and Light Guide Concepts

- Detector concept uses a layered “stack” of tungsten and fused silica (quartz) to induce EM showering and produce Cherenkov light
- “Baseline” design developed using GEANT4 optical MC simulation:
  - Current design uses a **4-layer stack** with **8mm tungsten** and **12.5mm quartz** pieces
  - Cherenkov light directed to **3inch PMT** using **air-core, aluminum light guide**



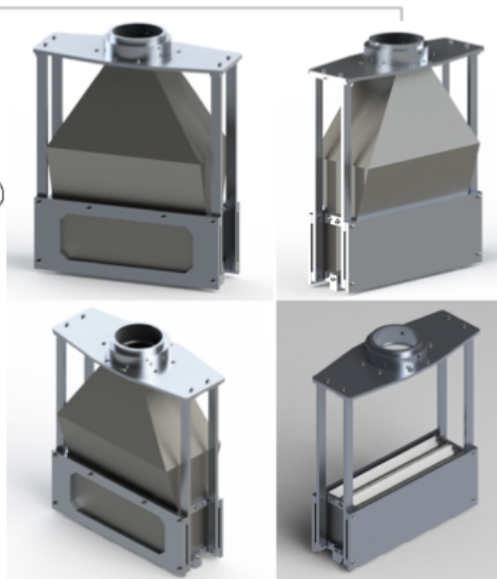
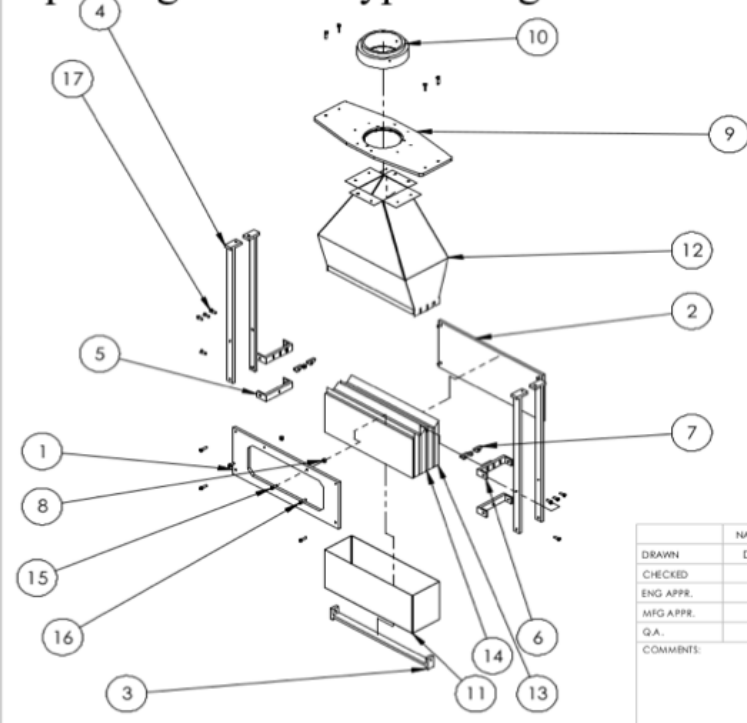
### Materials (all rad-hard):

- Tungsten is high purity (99.95%) and quartz is optically polished Spectrosil 2000
- Light guides are aluminum specular reflectors (Miro-silver 27, Anolux, or al. mylar, ...)
- Total radiation length:  $9.1 X_0$  tungsten +  $0.4 X_0$  quartz =  $9.5 X_0$

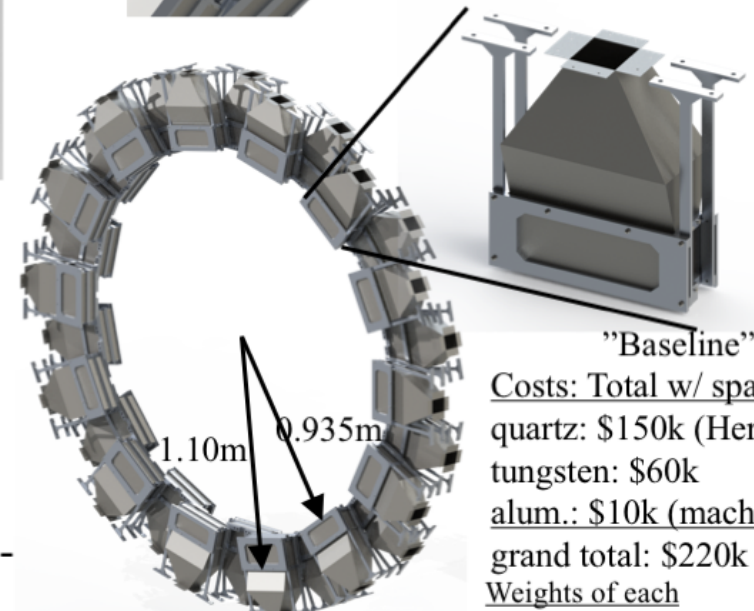
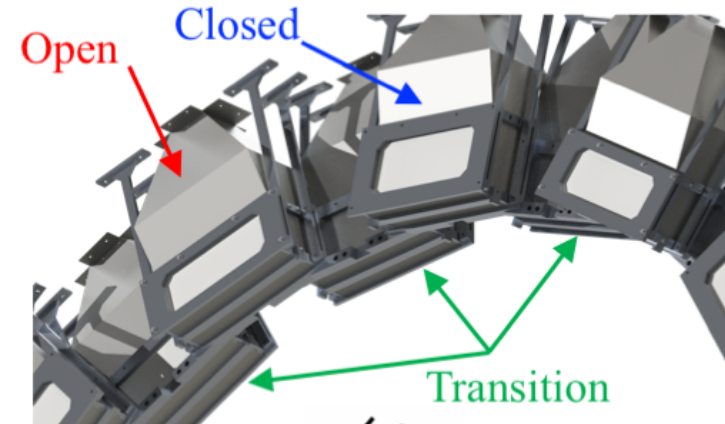


## Prototype Baseline Design and ring concept

### Open region Prototype Design



Idaho State University		
DRAWN	NAME	DATE
CHECKED	DS	05/16
ENG APPR.		
MFG APPR.		
Q.A.		
COMMENTS:		
TITLE: EXPLODED VIEW		
SIZE	DWG. NO.	REV
A	II	A
SCALE: 1:5		SHEET 2 OF 17

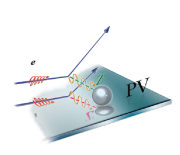


"Baseline"

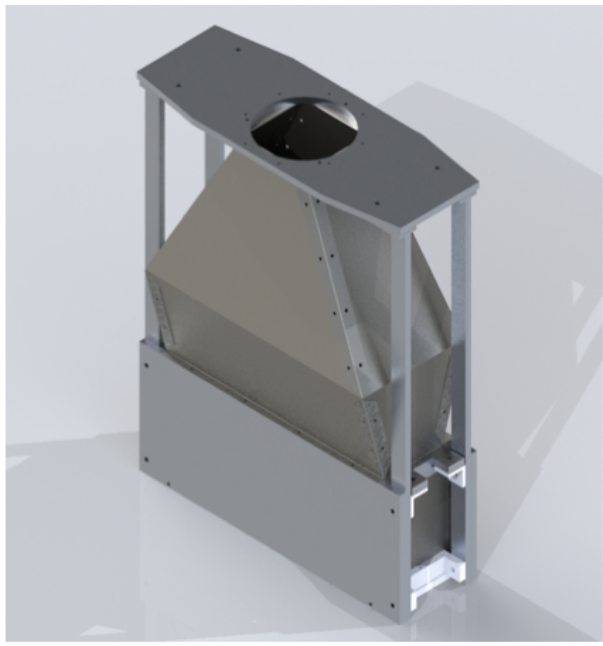
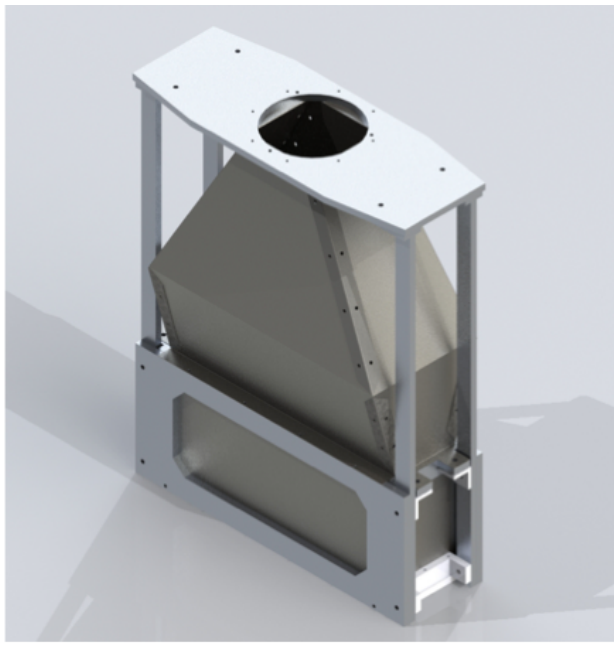
Costs: Total w/ spares  
 quartz: \$150k (Heraeus)  
 tungsten: \$60k  
 alum.: \$10k (machined)  
 grand total: \$220k  
Weights of each assembly:  
 Open: 39.7 lbs.  
 Transition: 42.5 lbs.  
 Closed: 50.8 lbs.  
 ring weight: 1230 lbs.

- Engineered shop drawings and Prototype CADs in hand
- **PLANS:** Finalize prototype Stack designs this summer, order quartz by fall, construct in winter 2018 and test in spring using 5 - 8 GeV electron testbeam at SLAC
- Shower-max ring design concept: staggered in  $\hat{z}$  with reinforced struts and brackets. 28 detectors in ring: 7 Open, 7 Closed, and 14 Transition



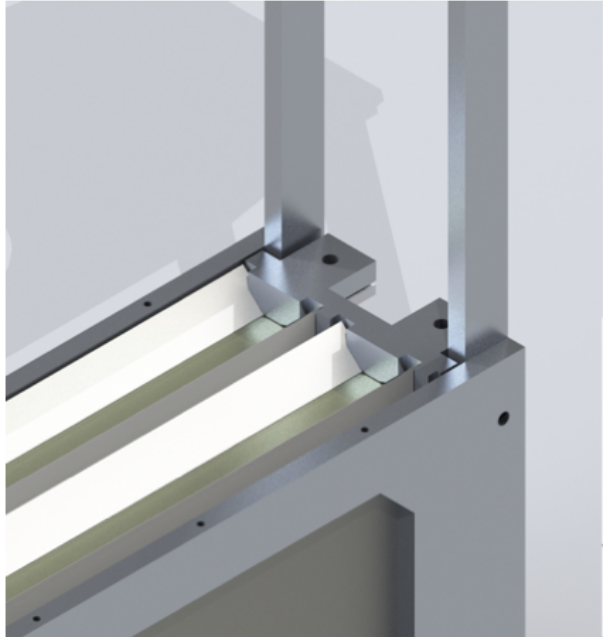
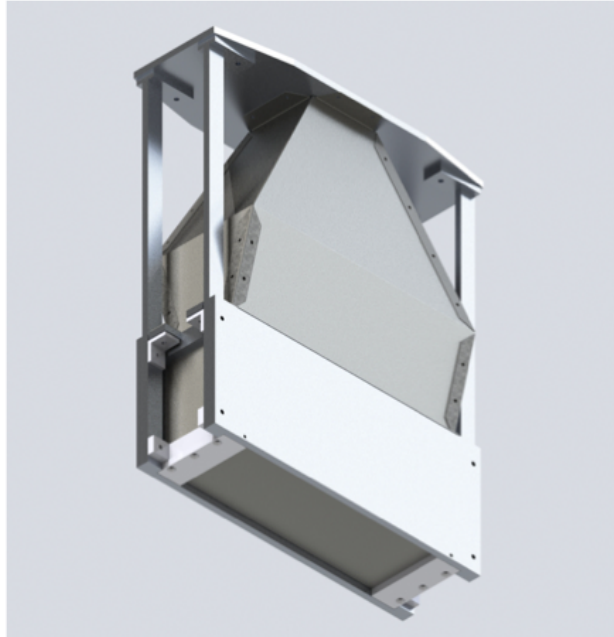


# Updated Full-Scale Prototype (1A) for Beamtest

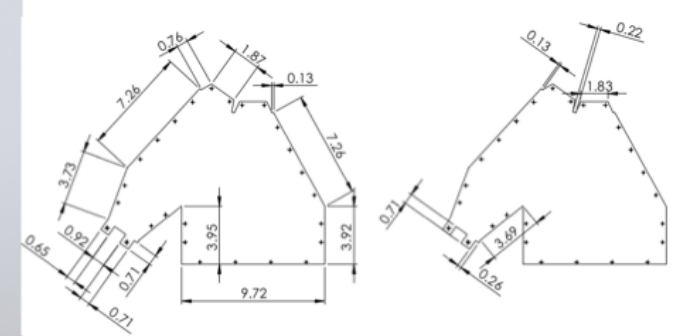


UNLESS OTHERWISE SPECIFIED:	NAME	DATE	Moller Collaboration	
DIMENSIONS ARE IN INCHES	DRAWN	DKS	1/14/18	TITLE
TOLERANCES:	CHECKED			Light Guide
FRACTIONAL ±	END APPR.			SIZE DWG. NO. REV
ANGULAR: MACH ± BEND ±	MFG APPR.			A I 0
TWO PLACE DECIMAL ±				SCALE: 1:10WEIGHT: SHEET 1 OF 9
THREE PLACE DECIMAL ±				
INTERPRET GEOMETRIC TOLERANCING PER:	Q.A.			
MATERIAL:	COMMENTS:			

ITEM NO.	PART	MATERIAL	QTY.
1	Light Guide - Back	0.002 ANODIZED MIRROR SILVER REFLECTIVE ALUMINUM SHEET	1
2	Light Guide - Front	0.002 ANODIZED MIRROR SILVER REFLECTIVE ALUMINUM SHEET	1
3	Long Flap	0.002 ANODIZED MIRROR SILVER REFLECTIVE ALUMINUM SHEET	2
4	Short Flap	0.002 ANODIZED MIRROR SILVER REFLECTIVE ALUMINUM SHEET	4
5	Suitcase	0.002 ANODIZED MIRROR SILVER REFLECTIVE ALUMINUM SHEET	2



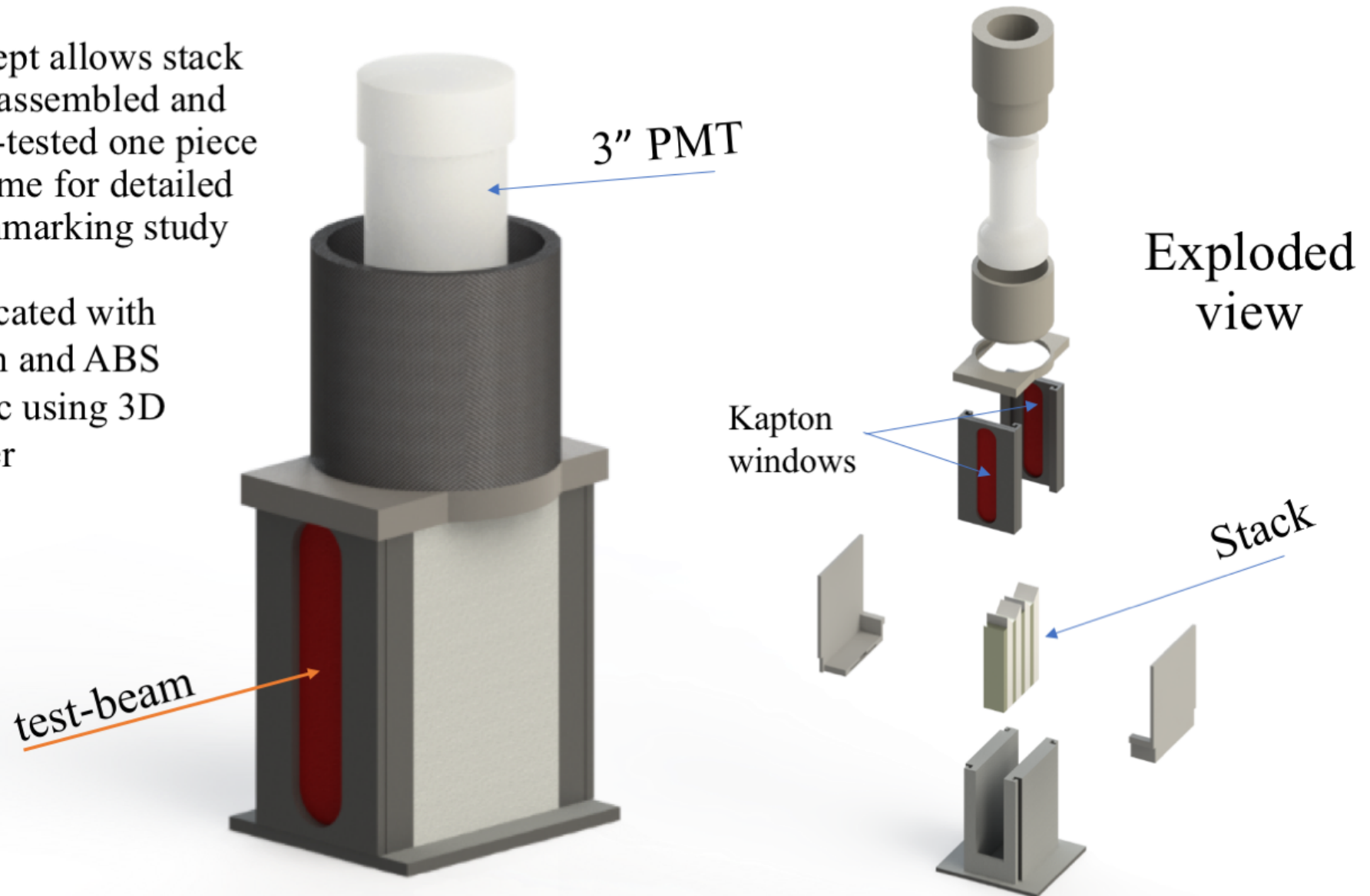
UNLESS OTHERWISE SPECIFIED:	NAME	DATE	Moller Collaboration	
DIMENSIONS ARE IN INCHES	DRAWN	DKS	1/14/18	TITLE
TOLERANCES:	CHECKED			Exploded View
FRACTIONAL ±	END APPR.			SIZE DWG. NO. REV
ANGULAR: MACH ± BEND ±	MFG APPR.			A II 0
TWO PLACE DECIMAL ±				SCALE: 1:10WEIGHT: SHEET 2 OF 9
THREE PLACE DECIMAL ±				
INTERPRET GEOMETRIC TOLERANCING PER:	Q.A.			
MATERIAL:	COMMENTS:			





# Shower-max Benchmarking Prototype concept

- Concept allows stack to be assembled and beam-tested one piece at a time for detailed benchmarking study
- Fabricated with Nylon and ABS plastic using 3D printer



**Config #1** (original baseline) benchmarking Prototype



## Quartz and Tungsten Ordered in Nov 2017

- For “benchmarking” prototype stack:
  - Quartz: 6 mm (thick) by 86 mm (tall) by 40 mm (wide) --4 pieces (\$975/piece = \$3.9k)
  - Quartz: 10 mm (thick) by 90 mm (tall) by 40 mm (wide) --4 pieces (\$1005/piece = \$4.0k)
  - Tungsten: 6 mm (thick) by 80 mm (tall) by 40 mm (wide) – 4 pieces (\$85/piece = \$340)
  - Tungsten: 8 mm (thick) by 80 mm (tall) by 40 mm (wide) – 4 pieces (\$110/piece = \$440)
  - Tungsten: 2 mm (thick) by 80 mm (tall) by 40 mm (wide) – 4 pieces (\$25/piece = \$100)
- For “full-scale” prototype stack:
  - Quartz: 6 mm (thick) by 111 mm (tall) by 246 mm (wide) -- 4 pieces (~\$1750/piece = \$7.0k)
  - Quartz: 10 mm (thick) by 115 mm (tall) by 246 mm (wide) -- 4 pieces (~\$1940/piece = \$7.8k)
  - Tungsten: 6 mm (thick) by 105 mm (tall) by 246 mm (wide) – 4 pieces (\$600/piece = \$2.5k)
  - Tungsten: 8 mm (thick) by 105 mm (tall) by 246 mm (wide) – 4 pieces (\$820/piece = \$3.2k)
  - Tungsten: 2 mm (thick) by 105 mm (tall) by 246 mm (wide) – 4 pieces (\$200/piece = \$0.8k)

Purchasing these pieces allows for Configs 1, 3, and 4 (A and B) to be tested

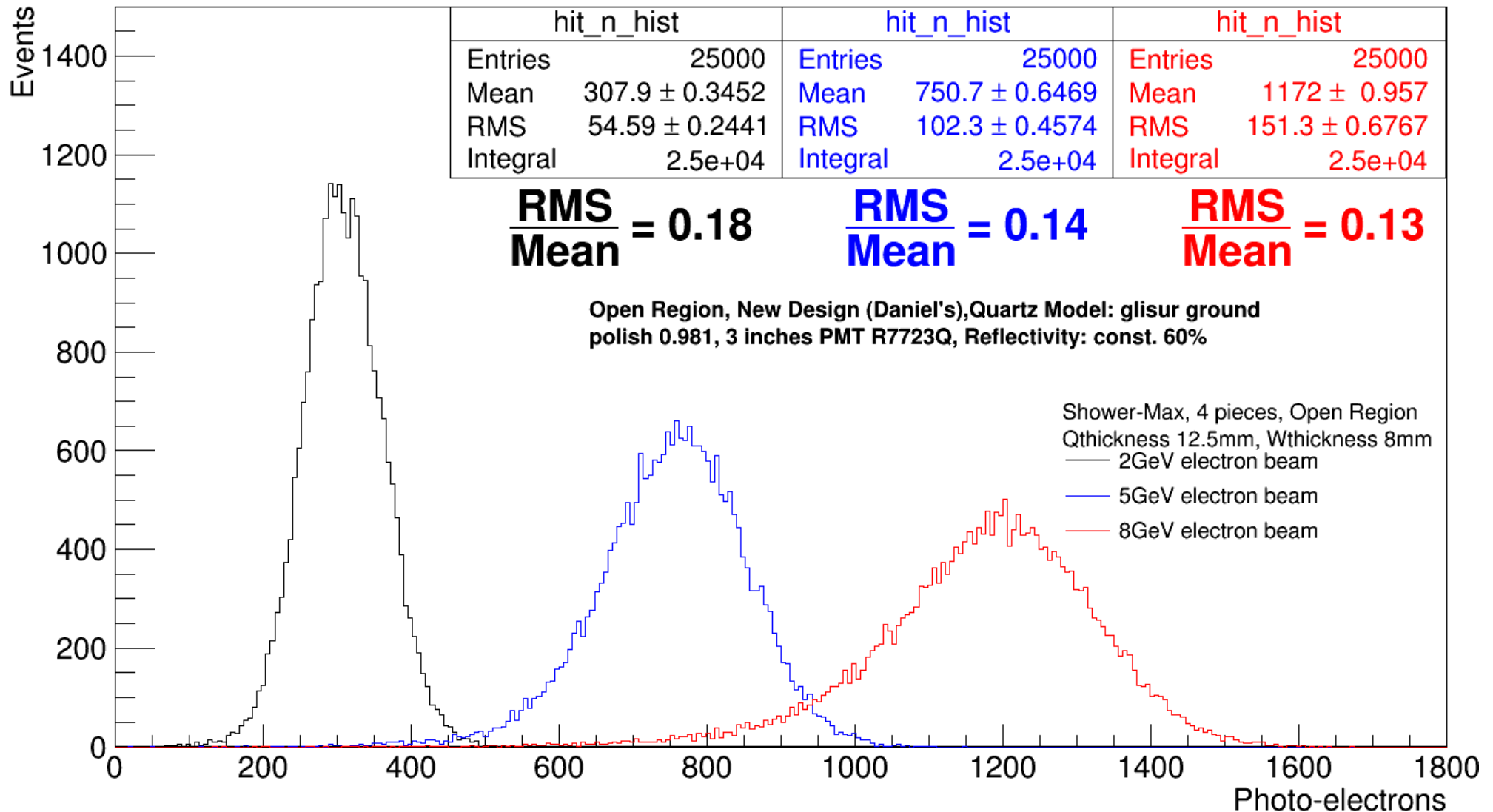
Total quartz: \$25k, total tungsten: \$7.5k: Total = \$32.5k

- Going with 6mm tiles allows construction of two benchmarking and two full-scale prototype sets
- Building two sets of prototypes will allow for more efficient testing during both SLAC testbeam and cosmic tests at SBU and ISU. We can each build a different configuration to test



## 4-layer baseline PE Dists for 2, 5, and 8 GeV

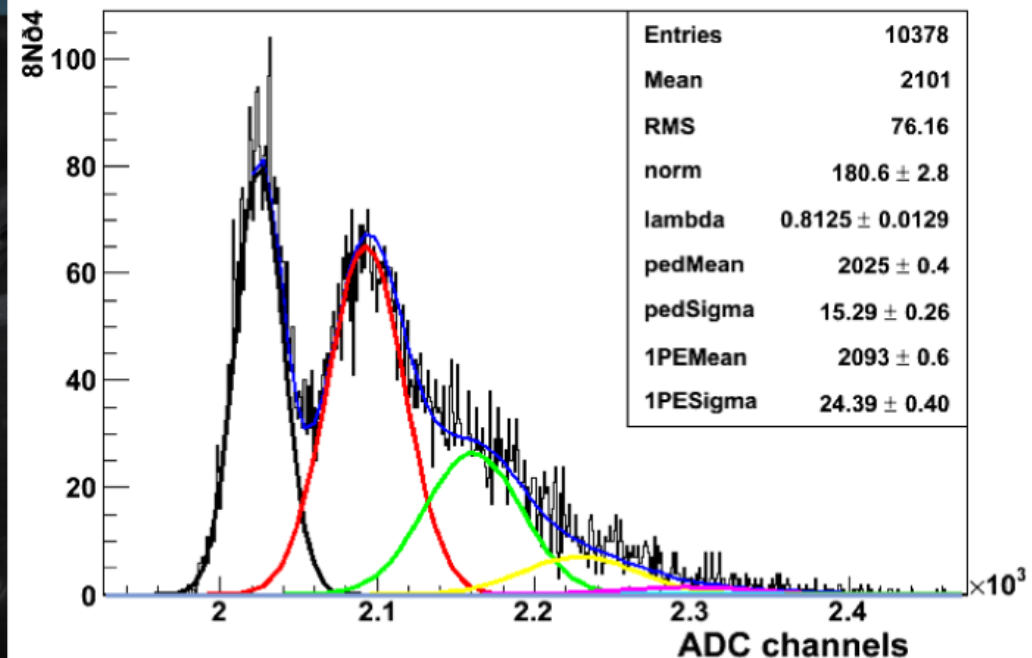
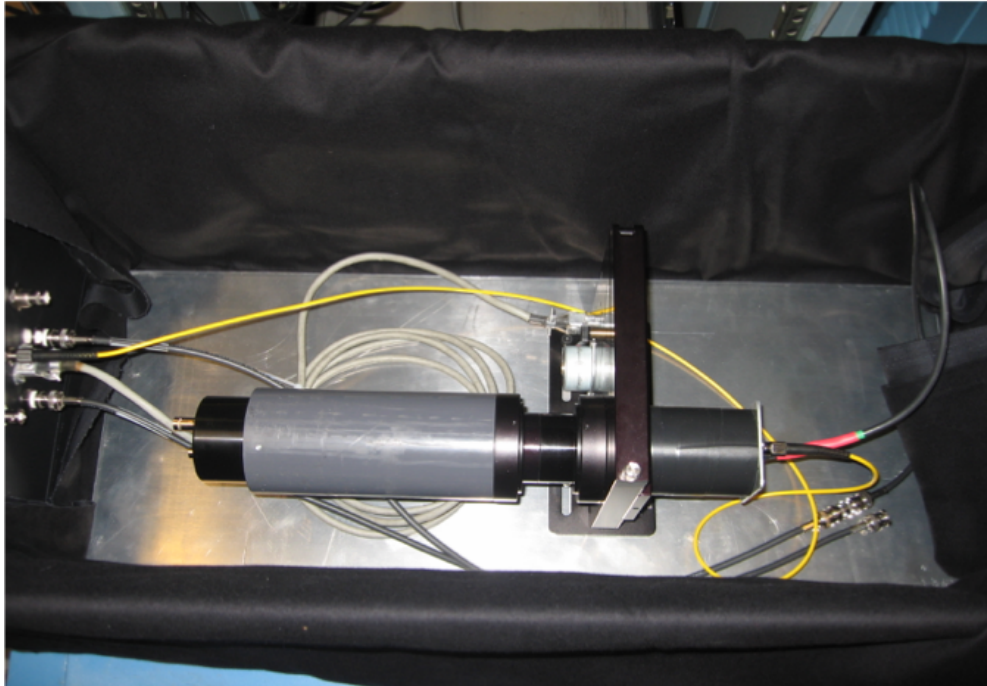
### PE Distribution: Showermax Open - 8mm W

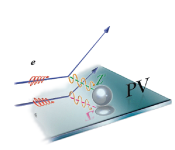




## Precision PMT Gain Measurements Planned

- ADC charge sensitivity calibrated
- PE peaks extracted using multi-Poisson fit algorithm
- Gains measured using linearity apparatus with CAEN LED driver, ND filter wheel, and CAEN fast amplifier
- Purchased 4 new R7723Q pmts (with Mod. base); purchasing two 3" ET9305QKB PMTs this spring





Thank you