Main Detectors for PREX-II/CREX

Dustin McNulty Idaho State University mcnulty@jlab.org

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Main Detectors for PREX-II/CREX

Talk Outline:

- Main Detector Design
- Motion Control System
- GEM Stand with Tandem Mount
- HRS Detector Package
- PMT Linearity and Gain
- Summary and Future Work





Main Integrating Tandem Detector Design



- PREX-II/CREX main detector design based on UMass Design-3
- Rotatable tandem mount designed, prototyped, and final version constructed
- Final design has shorter quartz rails and incorporates mu-metal shields and 3D printed ABS-plastic enclosure with Kapton windows





Quartz Geometry Plans (Preliminary)



Beam's view. Note "stubby" quartz installed upstream, "full" quartz downstream – for illustrative purposes



Top view showing quartz-rail supports (at PMT end). No light guides or wrapping will be used.

- PREX-I quartz was 3.5 cm wide by 16 cm long by 6 mm/10 mm thick
- PREX-II and CREX quartz footprint will be similar to PREX-I
- Design can accommodate up to 4.8 cm wide quartz piece
- May need thinner quartz for PREX-II due to higher 1.8 GHz rate; possibly 4 mm thick...



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- Quartz spacing same as for rotary tandem mount (~ 16 cm)
- Used two Hamamatsu R7723Q pmts
- Quartz is wrapped with 1 mil Al. Mylai
- Took runs for each quartz thickness upstream and downstream
- Example raw data, pedestal fit, and ped-corrected ADC and PE dists



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PREX/CREX Tandem mount testbeam







Main Detector Components List: What we have & what is needed

- Left arm tandem mount complete
- Right arm tandem mount under construction (ready in a couple weeks)
- Have 4 R7723Q PMTs with characterized gain and linearity; would like at least 4 to 6 more for A_T detectors and spares (\$2k each -- \$8 \$12k total).
- Have 2 mu-metal shields in hand; purchased 2 more last month; may want 4 more for A_T detectors (\$500 each – total cost is \$2k)
- Have 10 mm and 6 mm thick PREX-I quartz; need to finalize geometry and purchase main detector quartz. Can likely use same pieces for PREX and CREX (Cost is expected to be roughly \$1k per piece for total of \$4k \$5k)
- We also need to purchase the "stubby" quartz pieces (if we want these for alignment validation studies during commissioning). Need at least 2 of these (total ~\$1.5k)
- Total cost of main detector and A_T components still to purchase: PMTs, bases and shields -- \$10k - \$14k; quartz \$5.5k - \$6.5k. Total: \$15k - \$20k



Tandem Mount motion system (Prototype shown)

RHRS





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- X, Y, and θ degrees of freedom for main detectors
- Velmex 5 and 15" Bi-slides for X and Y motion, respectively (from PREX-I, we've found 15" sliders but not 5" and no controllers or cables yet...)
- Velmex rotary stages (have one, need another)
- Transducers for position feedback (have 8 from PREX-I)
- A_T dets will use 2 4 Velmex X-slides per arm (have 4 from PREX-I)



Motion system Needs

- Stepper motor controllers: ~\$400 per channel. Need up to 7 channels per arm. Could purchase four 4-channel controllers (VXM-4) for total cost ~\$3200
- Need two 5" Bi-slides with motor: ~\$900 each for total cost of ~\$1800
- Need one rotary stage with motor and encoder: \sim \$2150
- Need 4 more X-slides, 2"(?) travel with motor for A_T's: ~\$500 each for total of ~\$2000
- May need cable extension kits (from controllers to computer): \$200 each. Total cost of \$800
- Total needed is \sim \$10k for motion system





 $\mathbf{PREX}/\mathbf{CREX} \text{ "small" } \mathbf{10x20} \text{ } \mathbf{cm}^2 \text{ } \mathbf{GEM} \text{ } \mathbf{trackers}$





- Custom CERN 10 cm by 20 cm active area triple GEM chambers
 - > 400 μ m pitch x/y, 4 + 2 Panasonic 130pin Readout connectors
 - Standard GEM spacing D-3mm-G1-2mm-G2-2mm-G3-2mm-RO
 - Standard HV filter circuit: uses CERN ceramic resistor
- Readout scheme based on INFN/UVA SBS rear-tracker: APV25FE → backplane PCB → VME MPD



GEM Readout Plans

- GEM readout scheme based on INFN/UVA SBS rear-tracker system:
 - Uses APV25FE rev4.1 cards (have 55 in-hand); each chamber requires 6 APVs
 - Requires new 4-slot and 2-slot "backplane" PCBs (have 36 in-hand)
 - Backplanes buss analog-out signals to MPD and pass digital ctrl signals to APVs
 - Have 6 VME MPDs (Multi-Purpose Digitizers); require 2 for each arm
 - Uses fast intel Linux ROCs (have 3 in-hand: GE XVB601); require 1 for each arm



Getting much advice and help from Paolo Musico and INFN group, Kondo Gnanvo, Chris Cuevas, Nilanga Liyanage, and Alexandre Camsonne





GEM DAQ Plans

- Each arm requires one VME crate with fast linux ROC and at least three other available slots; I am currently using two Dawn 4-slot VME crates
- Each arm will need a trigger interrupt module and 2 MPDs; currently using a CEAN V965 QDC and v1720 digitizer for triggering interrupts (we will likely want JLab TIs for this—but needs to be old Tis if use CODA 2)
- Using CODA 2.6.2 with MPD drivers and support from JLab DAQ group
- Planning to use 6 10 meter long high speed HDMI cables for analog and digital signals (have all cables in hand)
- Each arm will need a CAEN N1470 HV NIM module (or equivalent) with 3 available channels (-4000 V); I currently have one HV module. *We will need an HV supply for the other arm*
- Each arm will need LV power supply with 5.0, 2.5 and 1.25 V; I have one already built and plan to build another (note that 5.0 V is for PMT preAmps)







GEM DAQ Progress



- Two GEMs complete with readout electronics chain: APV ↔ backplanes (HDMI) ↔ MPD
- Discovered I²C addressing problem with backplanes; fixed with small jumper-wires
- Working with B. Moffit on CODA VME drivers specific for our setup – using v965 QDC or v1720 for triggering backplane interrupts
- CODA MPD (v3 firmware) system is up and running with two GEMs (so far); can take pedestal data; have working decoder; developing rudimentary analyzer, but have signal timing issues...
- Found out at SBS meeting last winter that there is a VME readout firmware problem with the factory MPDs. Paolo has fixed and given instructions on how to update MPD firmware
- Working wit Bryan Moffit and INFN group to get VME MPD4 DAQ acquiring signals



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Sample GEM Pedestals (very preliminary)



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GEM Chamber Mounting Concept





Photo showing rail support brackets



G10 platforms (1/16 in. thick) for GEMs: supports readout electronics



Two Chambers installed; gas flowing

- 1" Extruded aluminum framing system for GEM mount
- Each arm will use three GEM chambers: one upstream and two downstream of quartz
- GEM ladder-frame mounts to Velmex slider post using cleats

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Cosmic Stand (with Rarm Tandem Det & GEMs)









Main Detector Stands



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RHRS Tandem Quartz Mount with GEMs







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HRS Detector Package Torque Analysis



Using HRS hut coordinate system

• Origin defined at the center of the 5-inch travel (top) slider platform.

Center of Mass Analysis





To be determined by the second support frame





HRS Detector Package Torque Analysis

↑ X	Dot Color	Assembly	Weight (lbs)	Torque around y-axis	Torque around x-axis
Z Z	Green	GEM	30.0	(111-lbs) 81.90	(111-1bs) -120.0
4.0 in 4.96 in		Support Frame	2010	01150	12010
2.73 in 1.60 in	Blue	Rotary Stage	6.43	6.11	-15.69
1.30 in 0.95 i	Orange	Tandem Quartz Mount	10.32	-13.42	-96.08
	Red	Total Detector	46.75	74.59	-231.77
		Package			





HRS Detector Package Torque Analysis









Counter weight (+ supports): 26.4 lbs

- Using the new center of mass location and new load (old load + counterweight) of **73.17 lbs**, net torques were calculated.
- Net Torque about X-axis: (0.015 in)*(73.17 lbs) = 1.10 in-lbs
- Net Torque about Y-axis: (0.026 in)*(73.17 lbs) = 1.90 in-lbs
- Net Total Torque: $((1.10)^2 + (1.90)^2)^{1/2} = 2.19$ in-lbs





HRS Detector Package for PREX-II/CREX

- All HRS standard detector packages removed except for VDCs: No S1, S2, Cerenkov, or Calorimeter
- For event-mode operation: Use S3 (or S0) for triggering
- Additional array of large GEMs from UVA group to be installed above PREX detector package
- A_T detector not shown: will mount just above small GEMs
- Plan to use same hardware and mounting/installation concept developed for PREX-I







HRS Detector Package for PREX-II/CREX







PMT Linearity Studies: Apparatus)







PMT#2 (ZK5365 R7723Q) at 0.3 nA

nonLinearity vs HV for PMT#2, 0.3nA LL



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PMT#2 (ZK5365 R7723Q) at 0.5 nA

nonLinearity vs HV for PMT#2, 0.5nA LL



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PMT#2 (ZK5365 R7723Q) at 0.3 nA

120 Hz Flipping Frequency

240 Hz Flipping Frequency

Run	ΗV	PreAmp	non-Lin	Error	$\chi^2/6{ m df}$	Run	ΗV	PreAmp	non-Lin	Error	$\chi^2/6{ m df}$
2505	-1160	0.3	0.605	0.119	7.365	2533	-1160	0.3	0.301	0.166	16.49
2504	-1080	0.3	0.382	0.151	16.47	2530	-1080	0.3	0.071	0.176	15.60
2502	-1045	0.5	0.466	0.183	26.16	2528	-1045	0.5	0.367	0.229	22.85
2501	-980.0	0.5	0.408	0.175	28.09	2527	-980.0	0.5	0.388	0.271	30.19
2503	-915.0	1.0	0.294	0.147	19.07	2526	-915.0	1.0	0.236	0.211	24.16
2499	-855.0	1.0	0.126	0.114	2.584	2525	-850.0	1.0	0.158	0.184	22.17
2497	-805.0	2.0	0.021	0.147	17.43	2524	-805.0	2.0	-0.109	0.166	14.20
2496	-755.0	2.0	0.066	0.177	20.95	2523	-755.0	2.0	0.140	0.201	22.47
2508	-715.0	4.0	0.122	0.132	12.55	2516	-715.0	4.0	0.026	0.218	30.14
2507	-665.0	4.0	0.059	0.201	27.86	2515	-665.0	4.0	0.067	0.219	35.07
2491	-600.0	10.0	-0.145	0.145	12.96	2514	-600.0	10.0	-0.235	0.117	6.974
2490	-560.0	10.0	-0.040	0.110	4.177	2512	-560.0	10.0	-0.104	0.173	18.00

Note: The entries between double lines are acceptable.





Sign of non-linearity

- Low work function of dynode coatings
- Probability of thermal electrons from dynode coatings and photocathode
- At high currents, space charge can influence the electron trajectories causing collection losses
- At still higher currents, space charge can cause some electrons to return to the surface from which they originate
- Below certain HV, the space charge effect causes collection loss at higher LLs resulting negative non-linearity
- Above certain HV, the HV between the dynodes and between the last dynode and the anode is sufficient enough to overcome the space charge effect
- At higher HV, even the space charge gets dragged towards the anode increasing the collection at higher LLs resulting positive non-linearity

Mathematical Section PREX/CREX Collaboration



PMT Characteristics from Hamatsu Handbook



- At high currents, space charge can influence the electron trajectory, causing collection losses.
- Light level > certain level PMT output becomes saturated and is no longer proportional to the light level.
- Light hysteresis and voltage hysteresis of the PMT.
- Signal to noise ratio of the PMT.
- Drift (time stability) warm up helps minimize.
- PMT operation stability depends on the total stability of the power supply characteristics (drift, ripple, temperature, input regulation, load regulation).





PMT Gains (Work by Brady Lowe)

- Constructed second dark box (identical to linearity box) for dedicated PMT gain measurements
- Will measure gains of all PMTs from 2000 V down to as low as we can go; anticipated precision at sub 5% level
- This project is fairly advanced and is now in full production mode







PMT Linearity and Gain Summary

- So far we're able to find zero non-linearity crossings for specific HV and preAmp settings for all light levels tested so far (10 nA, 6 nA, 0.5 nA, and 0.3 nA). All these results can be found here: http://daq3.physics.isu.edu/linearity/PMT.html
- The above tested light levels were based on PE yields and 1 GHz/ 50 MHz focal plane rates for PREX/CREX
- Will reassess light levels given new focal plane rates (1.8 GHz for PREX and 30 MHz for CREX; planning to do optical simulations to explore different quartz thicknesses and ND filters for PREX-II main detectors
- $\bullet\,$ There is no significant difference between 120 and 240 Hz FF
- PMT gain calibrations are in full-swing and will be completed by this winter





GEM Summary

- "Small" GEM tracking system development well underway
 - 5 assembled and tested GEM chambers at ISU + 1 at SBU
 - All readout electronics in hand: 55 APVs, 6 MPDs, 20 two-slot and 16 four-slot backplanes; SBU has 18 APVs, 2 MPDs, 5 two-slot and 2 four-slot backplanes
 - GEM mounting concept developed and constructed
 - HV circuits assembled and burn-in procedure completed;
 but will likely change the divider for lower current
 - CODA DAQ with MPD drivers established; communicating with APVs and acquiring ped data; we have working decoder and rudementary analyzer; exploring basic functionality with source tests; hope to start cosmic tests soon



Main Detector Summary

- PREX-II/CREX main detector design essentially complete
 - Waiting for *final* PREX-II and CREX focal-plane footprints before finalizing quartz geometry
 - Will use bare, unwrapped quartz and no light guide
 - Rotary tandem mount concept vetted: Left arm tandem detector constructed and in cosmic test-stand
- Main detector PE yields and relative widths measured at MAMI for 6mm and 10mm thick tandem configuration
 - For unwrapped quartz, 6mm gives 37 peak PEs with 20% RMS/Mean; 10mm (downstream) gives 65 peak PEs with 28% RMS/Mean
 - Expected focal plane rates times these peak PE calibrations give PMT photocathode light levels—so we can prepare each PMT for optimal linearity—Devi working on this





Extra Slides

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Detector Configuration in HRS (Top View)



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