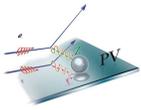


Optics Simulations

Dustin McNulty
UMass
mcnulty@jlab.org

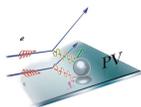
July 10, 2009



Optics Simulations

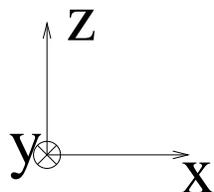
Outline

- New HRS Optics Tune for PRE_x
- A_T Hole Studies – Be X_o, FoM and Detector Size
- HAPPEX-III Focal Plane Distribution
- PRE_x Optics/detector beamtest during H-III Commissioning at 1pass
- Summary and Future Work



HAPPEX/PVDIS/PREX

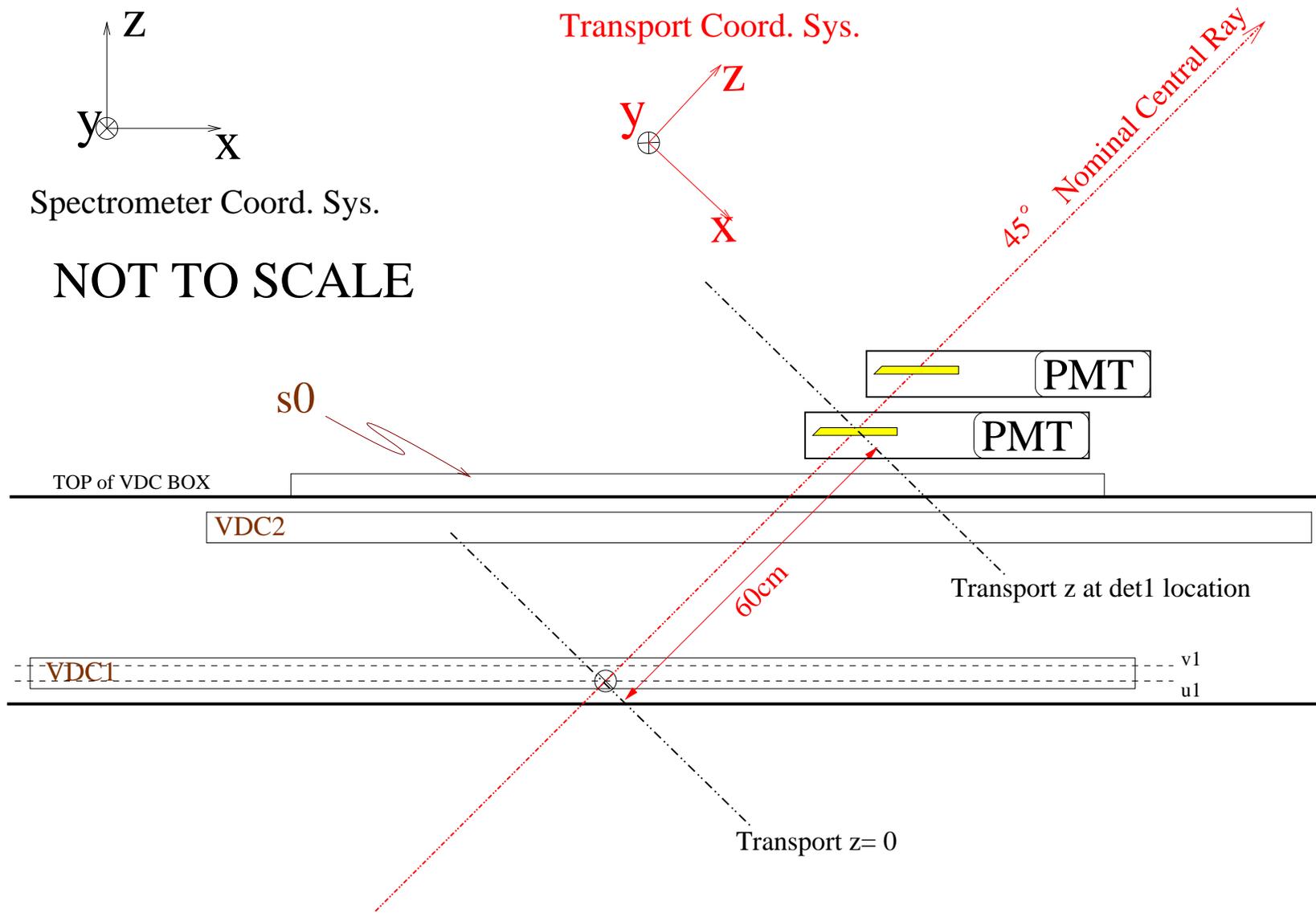
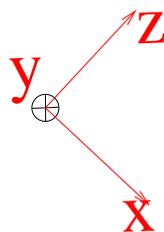
Jefferson Lab Hall A

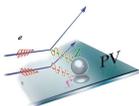


Spectrometer Coord. Sys.

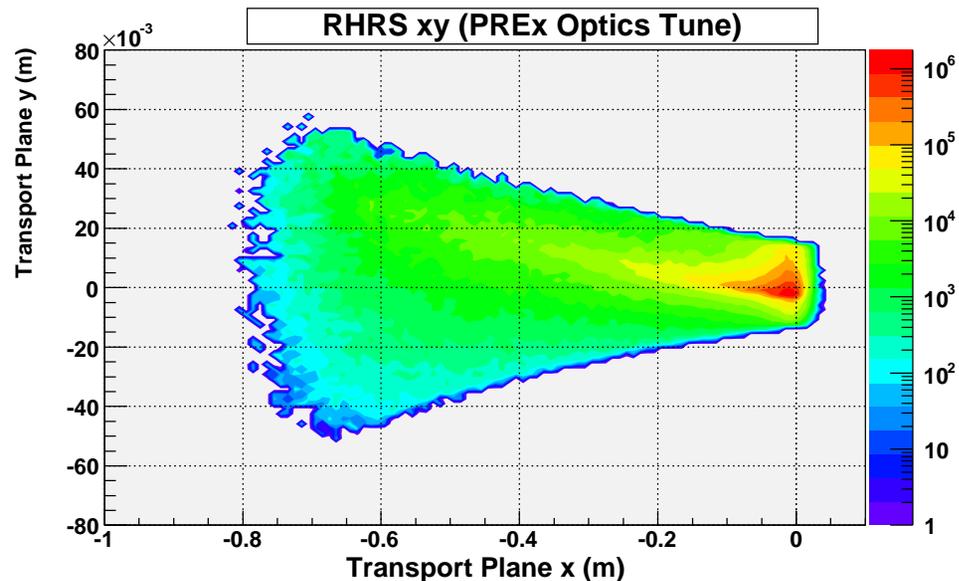
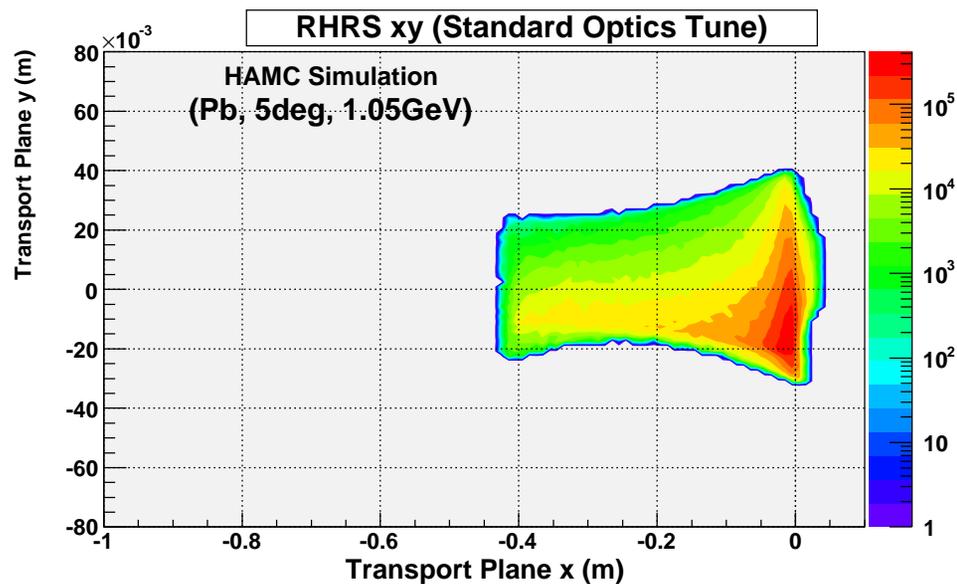
NOT TO SCALE

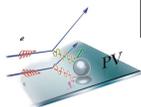
Transport Coord. Sys.



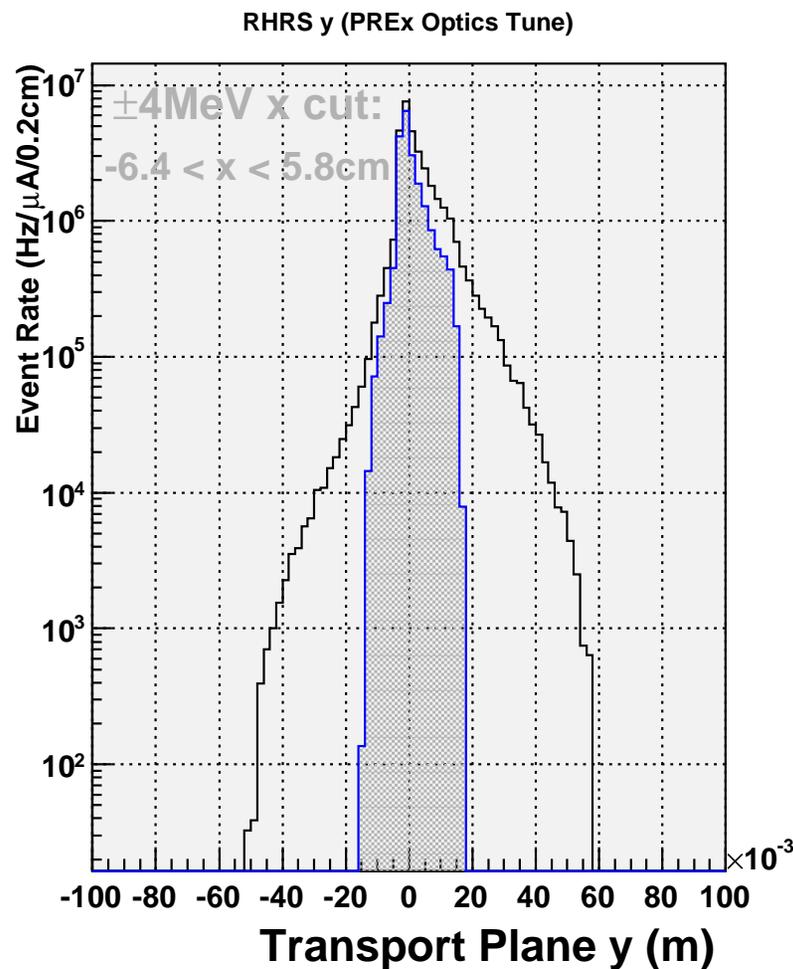
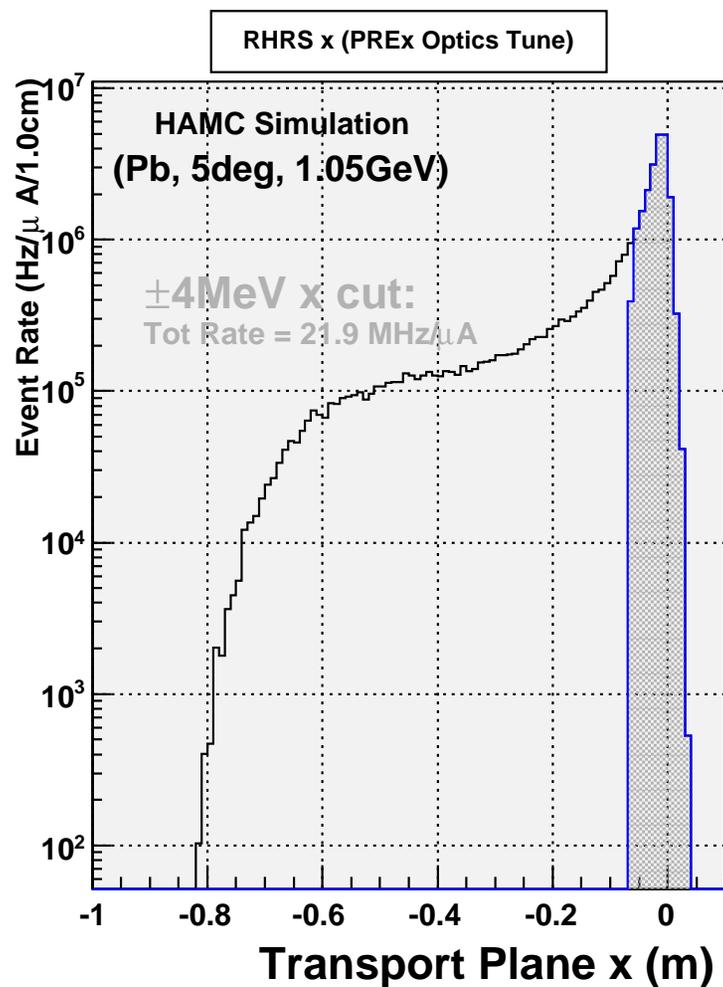


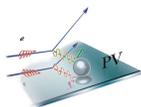
Standard (Old) and PREx (New) HRS Optics Tune



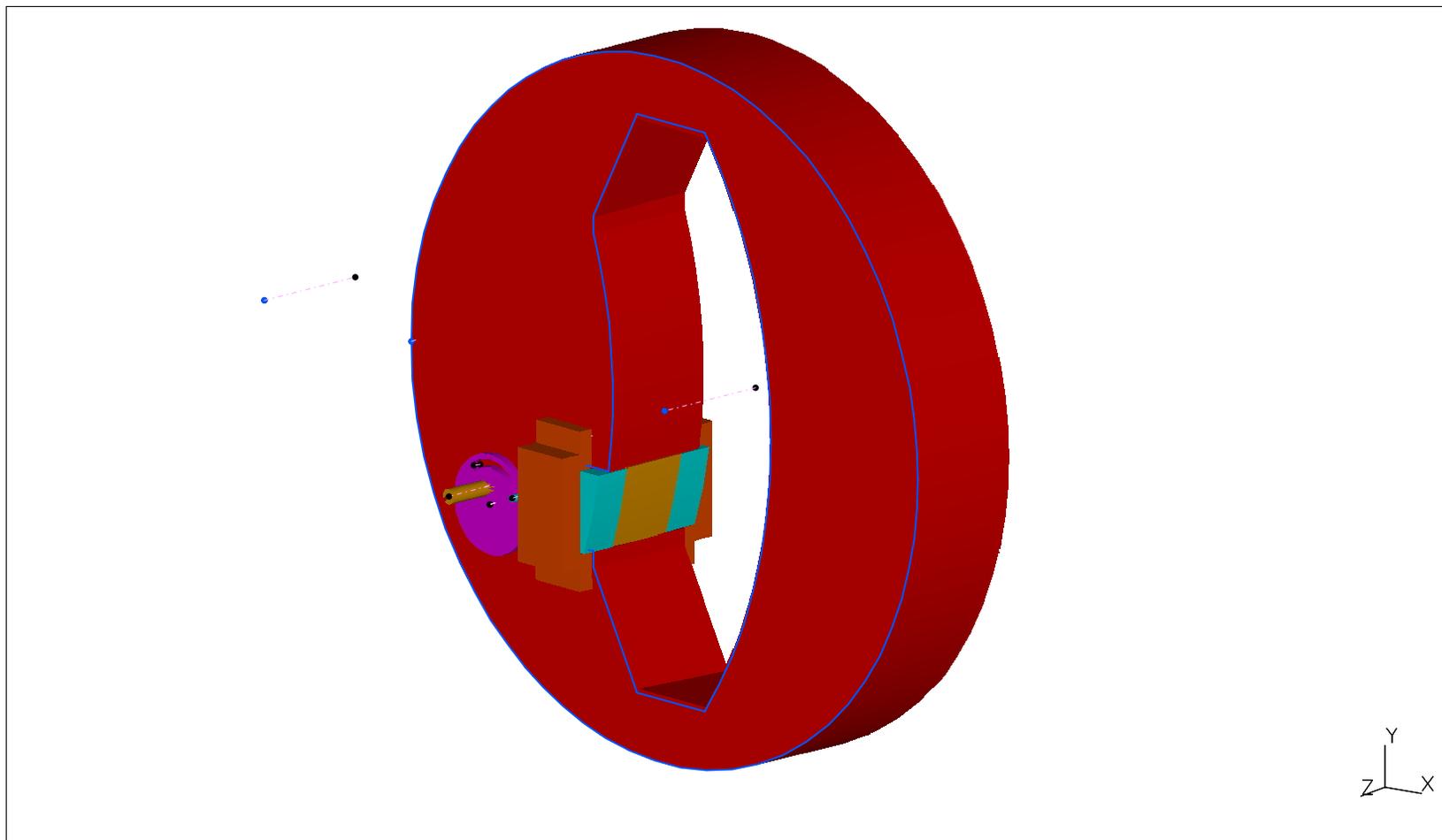


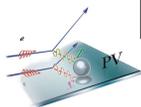
PREx: Transport Plane Rates



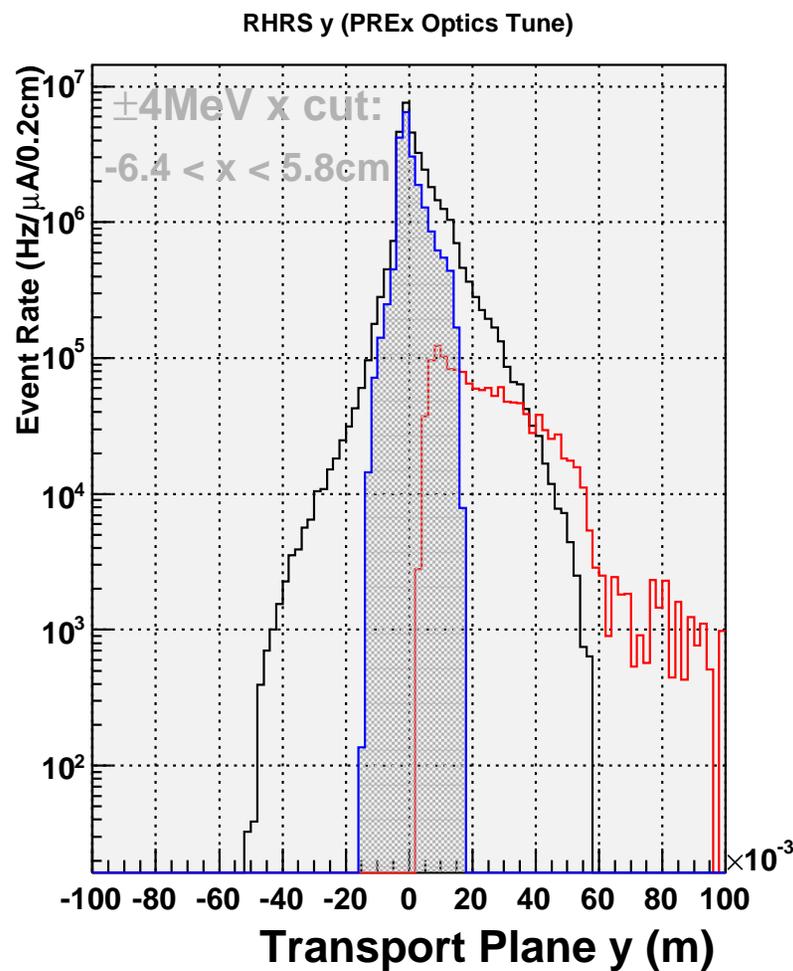
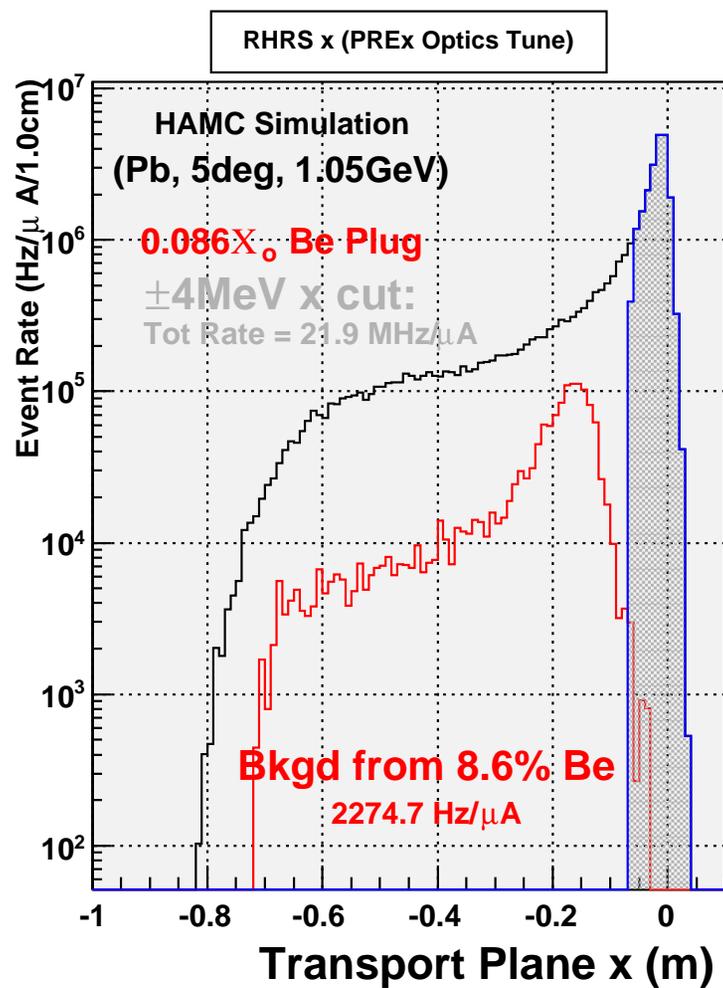


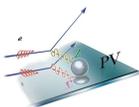
PREx Collimator and A_T Hole with ^9Be Plug





A_T Hole Event Distributions (PREx Tune)

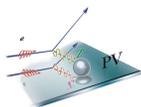




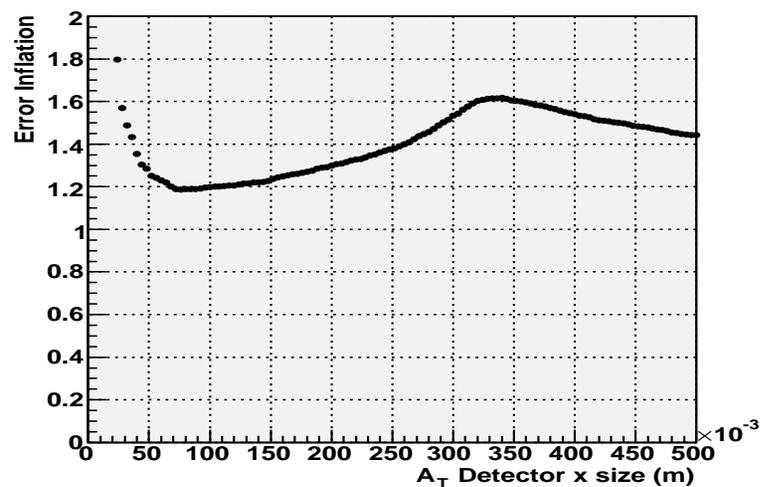
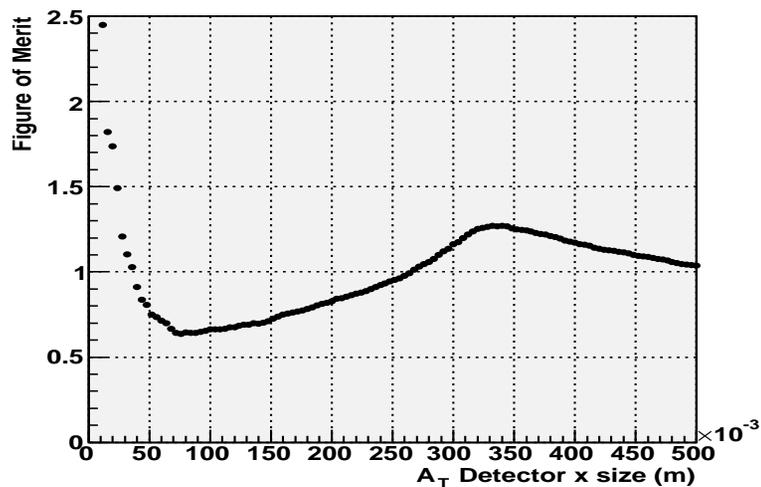
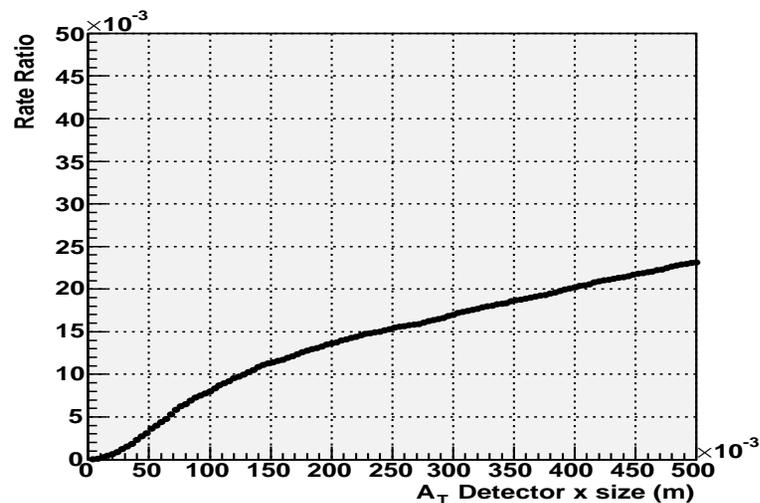
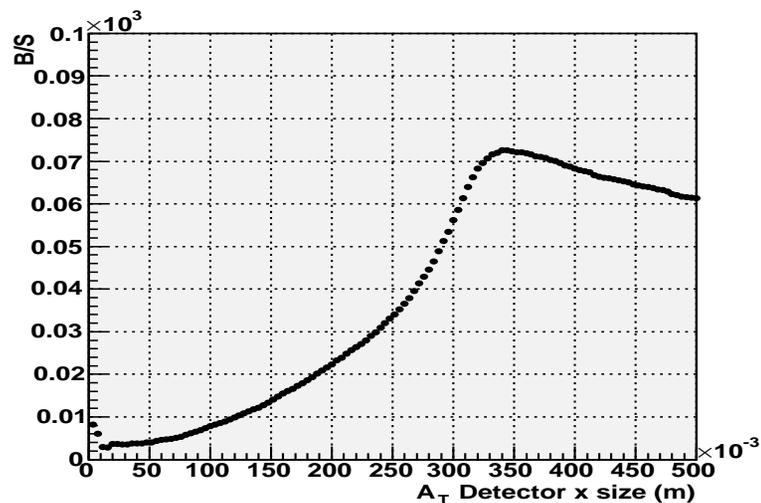
Paul's A_T Correction Figure of Merit

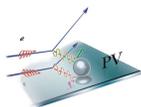
$$M = \frac{1}{E} \frac{1}{\sqrt{R}} \sqrt{1 + \frac{B}{S}} \quad (1)$$

- **E** is the enhancement in A_T sensitivity for the hole events (estimated to be ~ 50)
- **R** is the ratio of rates of the detected hole events relative to the main detector events (Additional factor of 2 reduction to account for radiative tail in hole events)
- **B/S** is the ratio of background to A_T hole signal
- The error inflation due to the correction is $\sqrt{1 + M^2}$
- For example, if $M = 0.64$, error inflation = 1.19

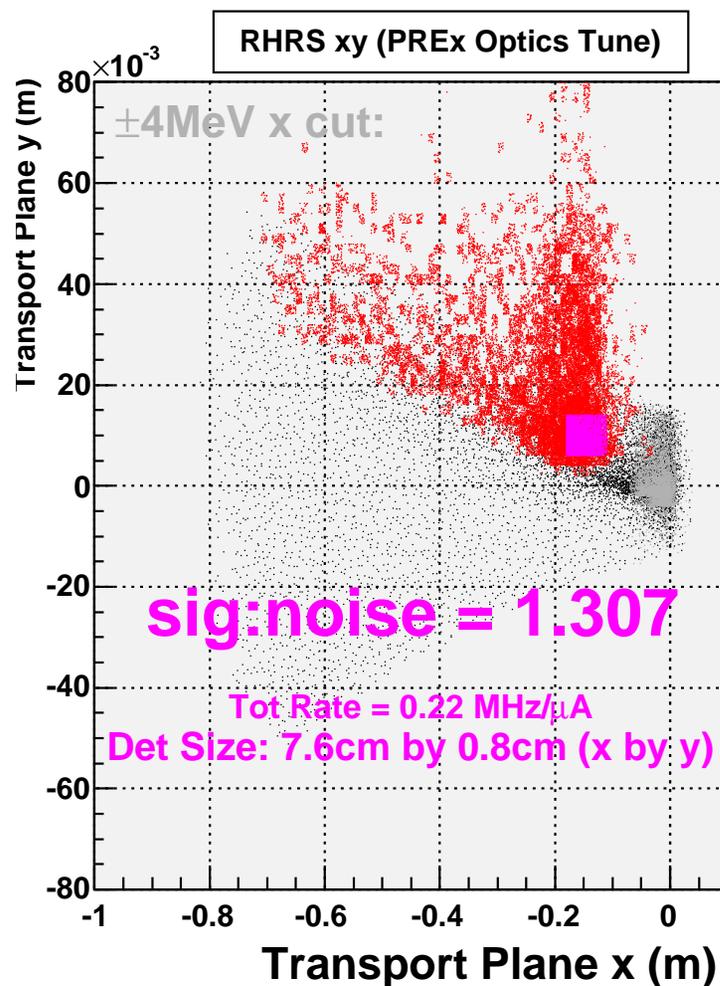
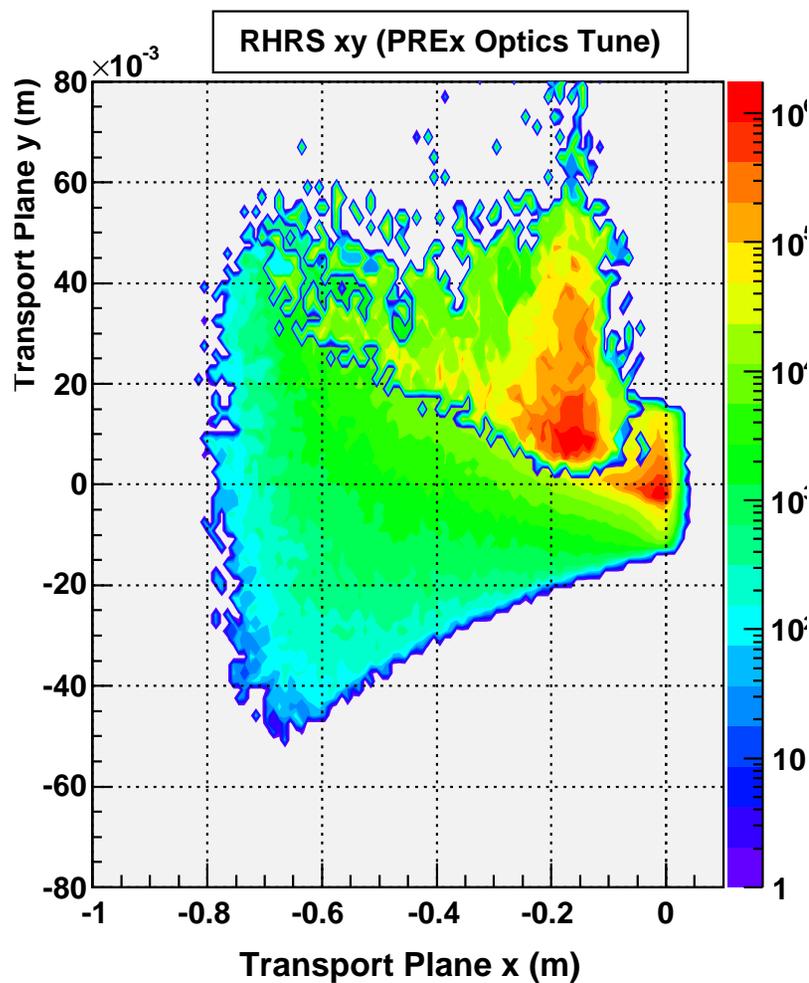


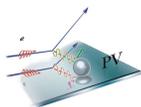
A_T Hole Figure of Merit Study



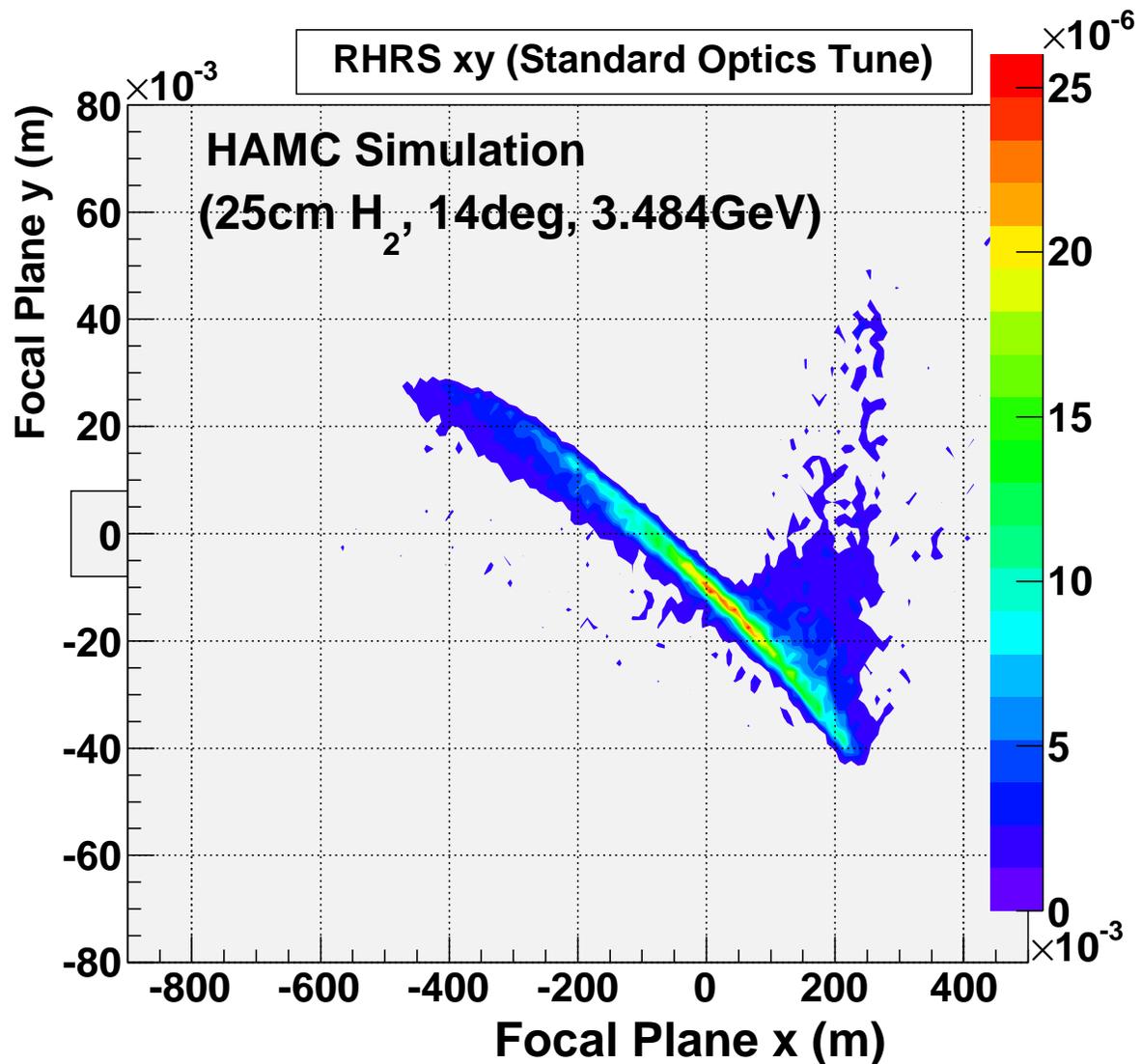


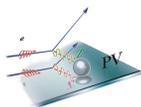
A_T Hole Detector (PREx Tune)



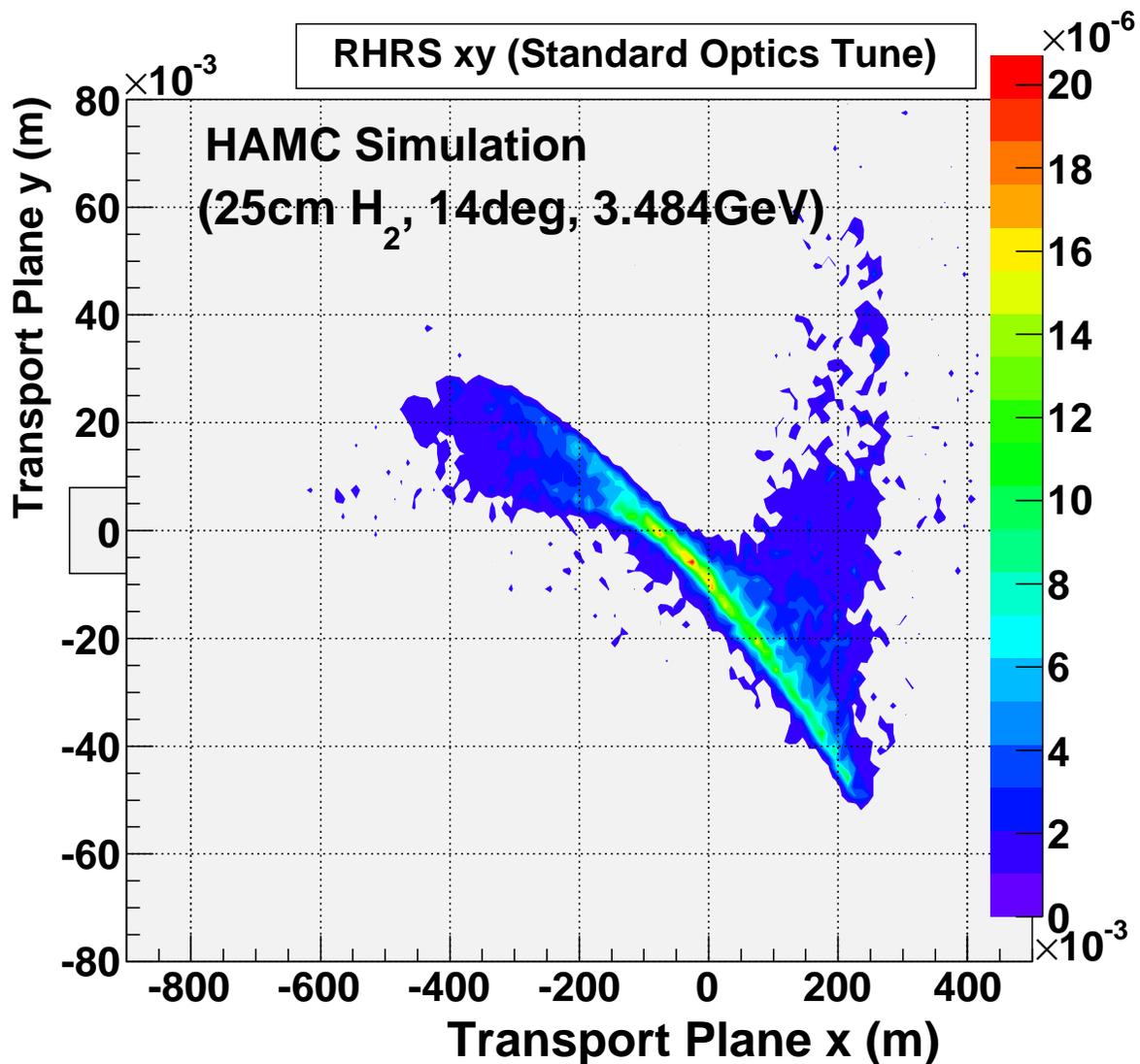


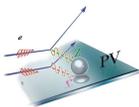
HAPPEX-III Focal Plane Distributions



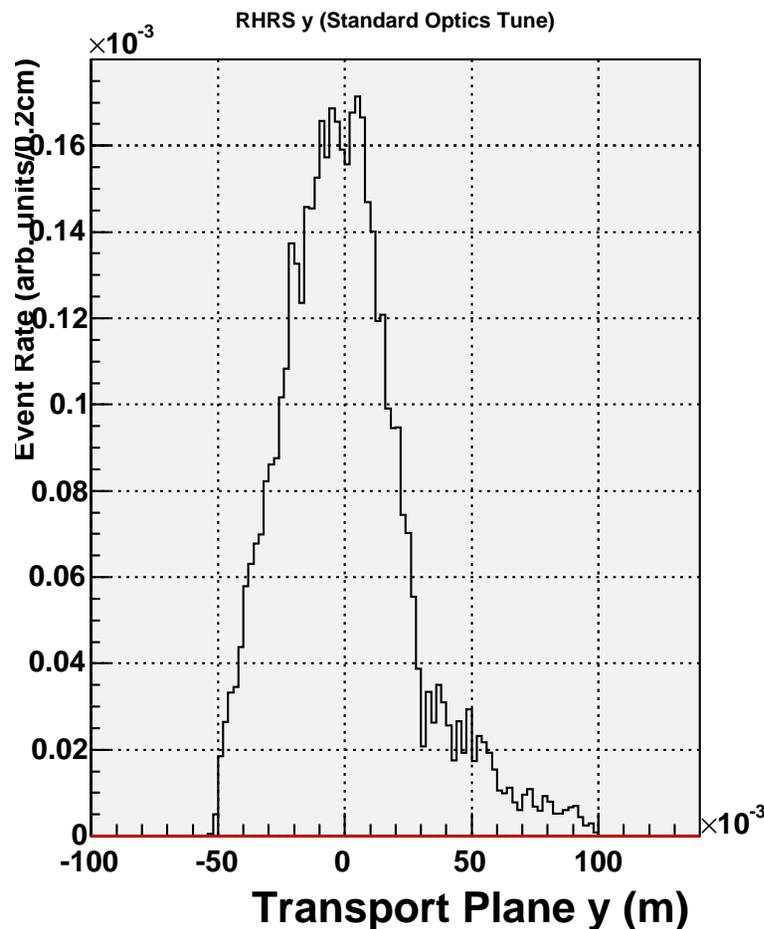
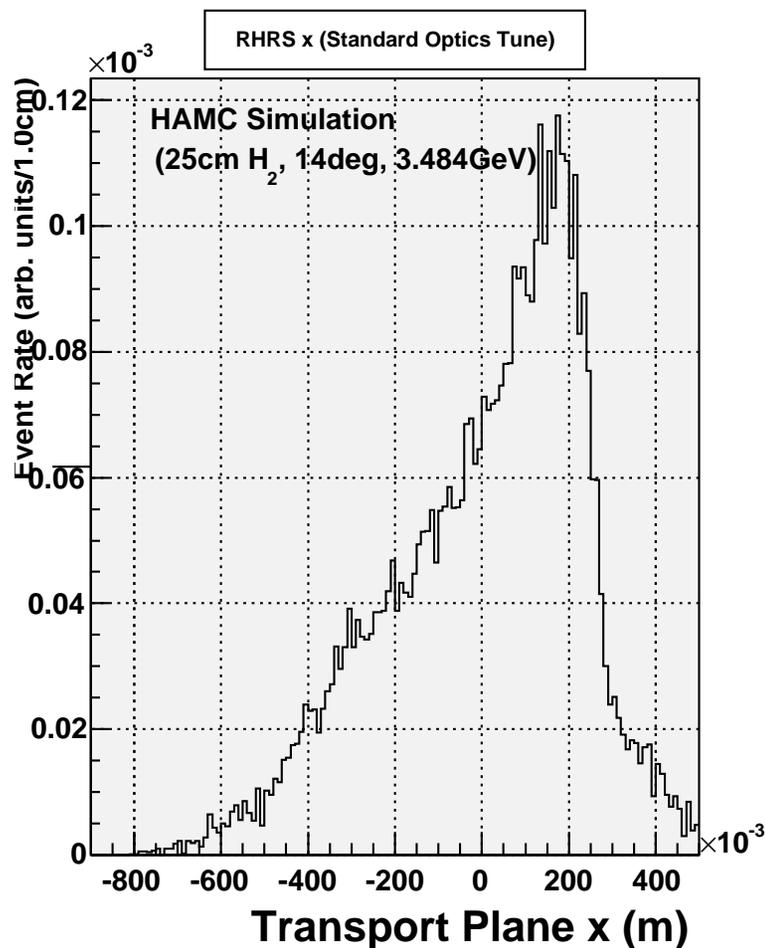


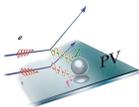
HAPPEX-III Transport Plane Distributions



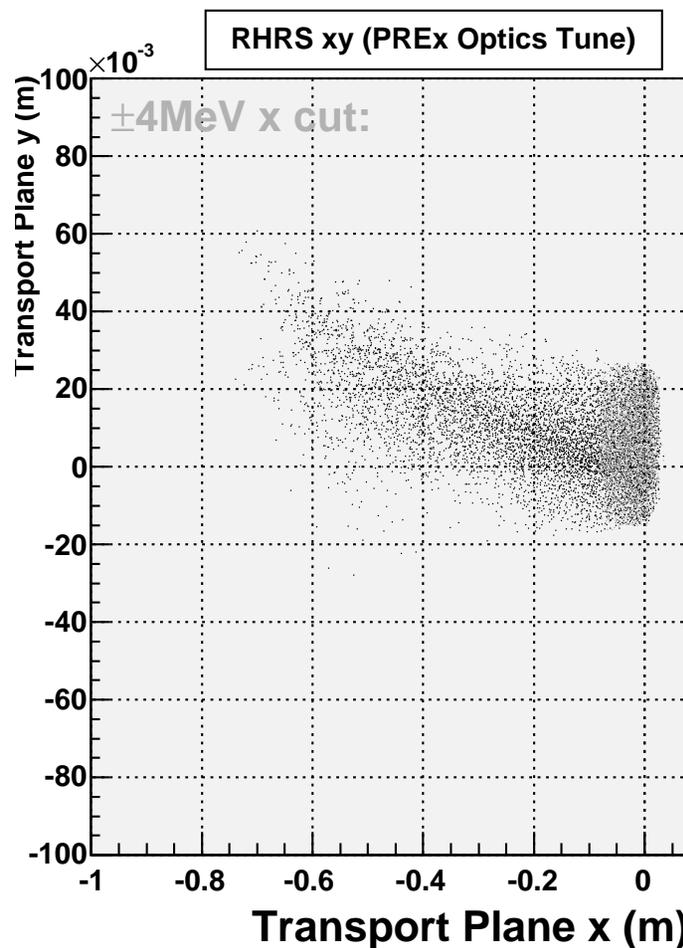
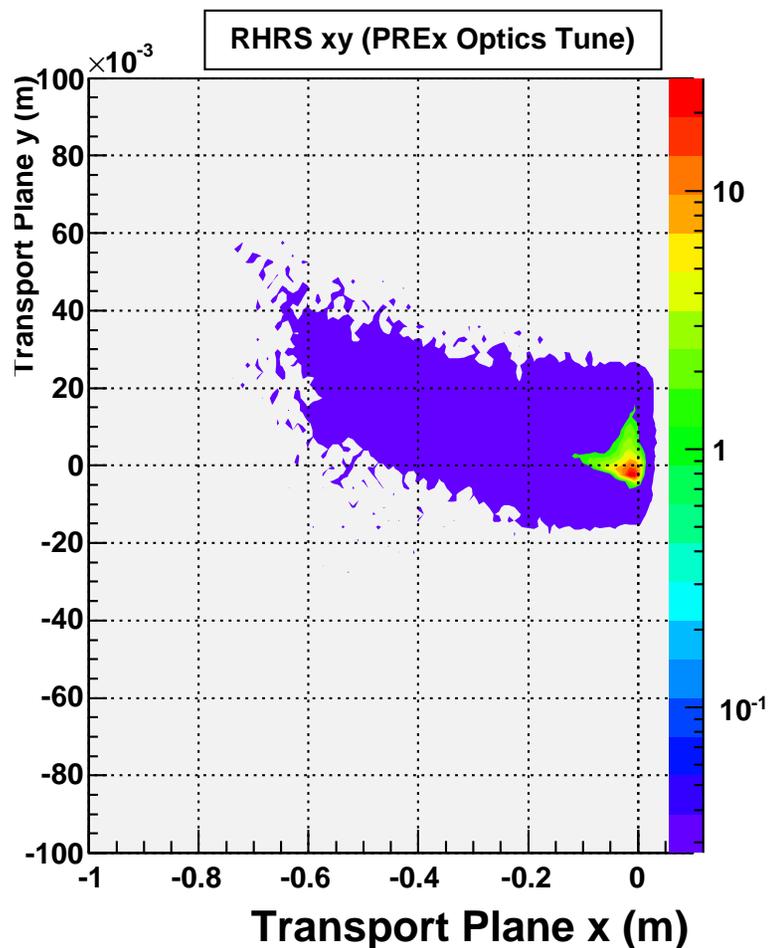


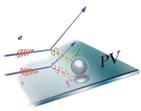
HAPPEX-III Transport Plane Distributions





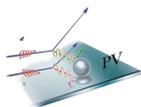
PREX beamtest during HAPPEX-III Commissioning



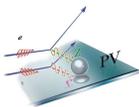


Summary and Future Work

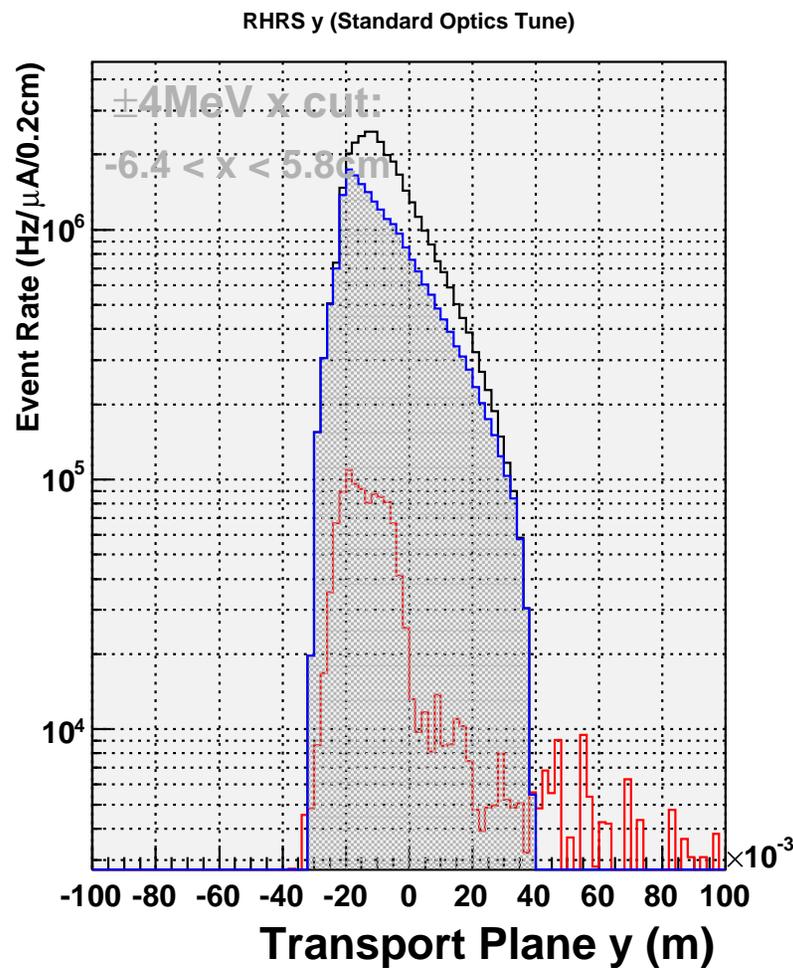
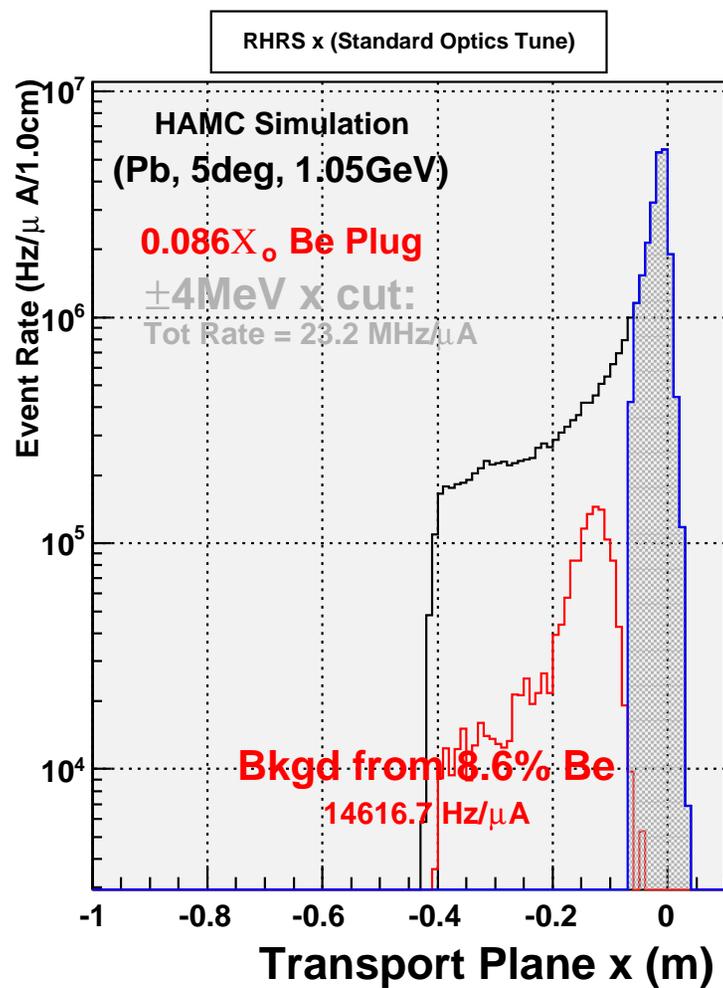
- Based on HAMC simulation of new PREX HRS optics, optimal prex detector size is roughly 10cm dispersive (x) by 4cm transverse (y)
- For A_T measurment/correction FoM studies, A_T detector size is optimized at 7.6cm in x and 0.8cm in y with estimated error inflation of $\sim 20\%$
- Examine sieve slit distributions using new PREX optics tune?
- Determine roll and pitch angles for HAPPEX-III detector positioning?
- Determine Rates for HAPPEX-III running
- Examine thick Ta spectra for H-III commissioning at 1pass?
-

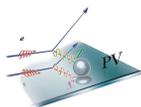


EXTRA SLIDES

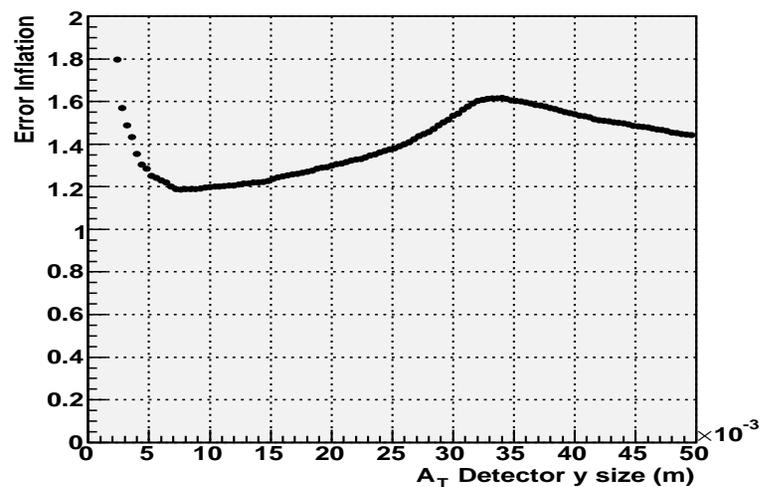
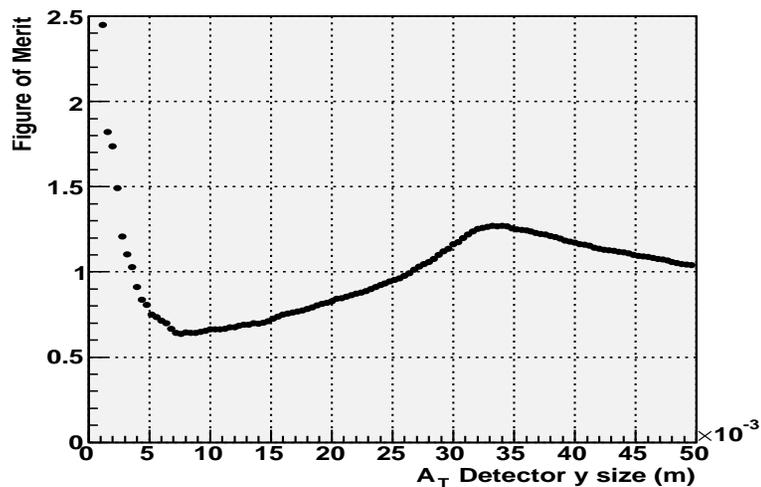
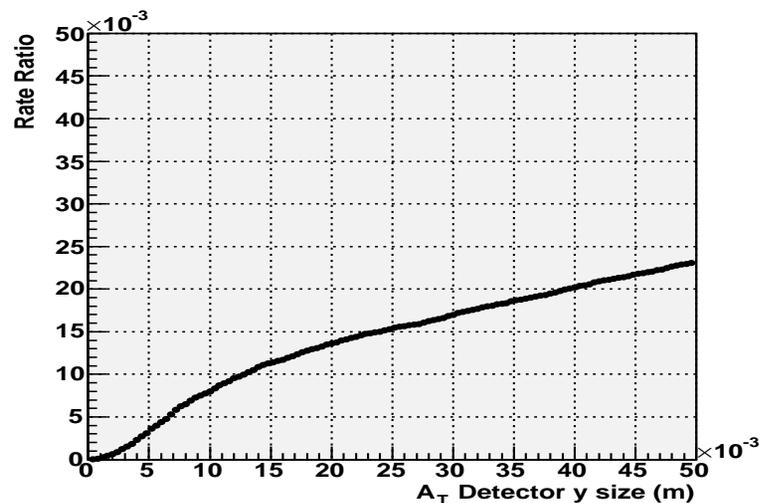
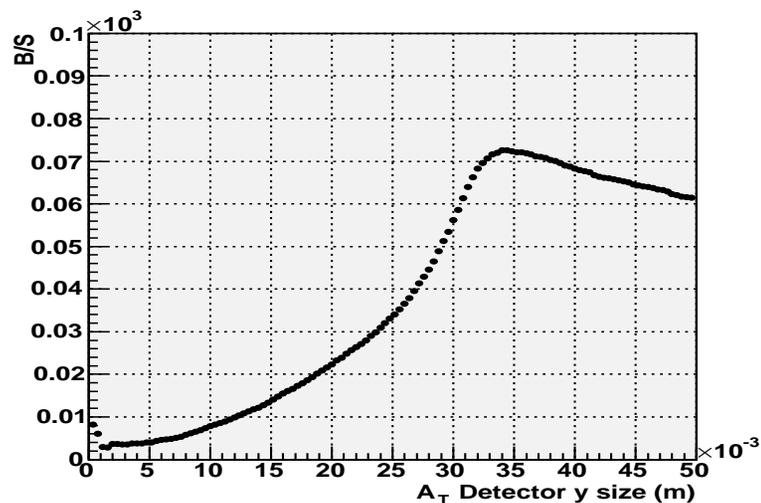


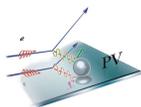
A_T Hole Event Distributions (Standard Tune)



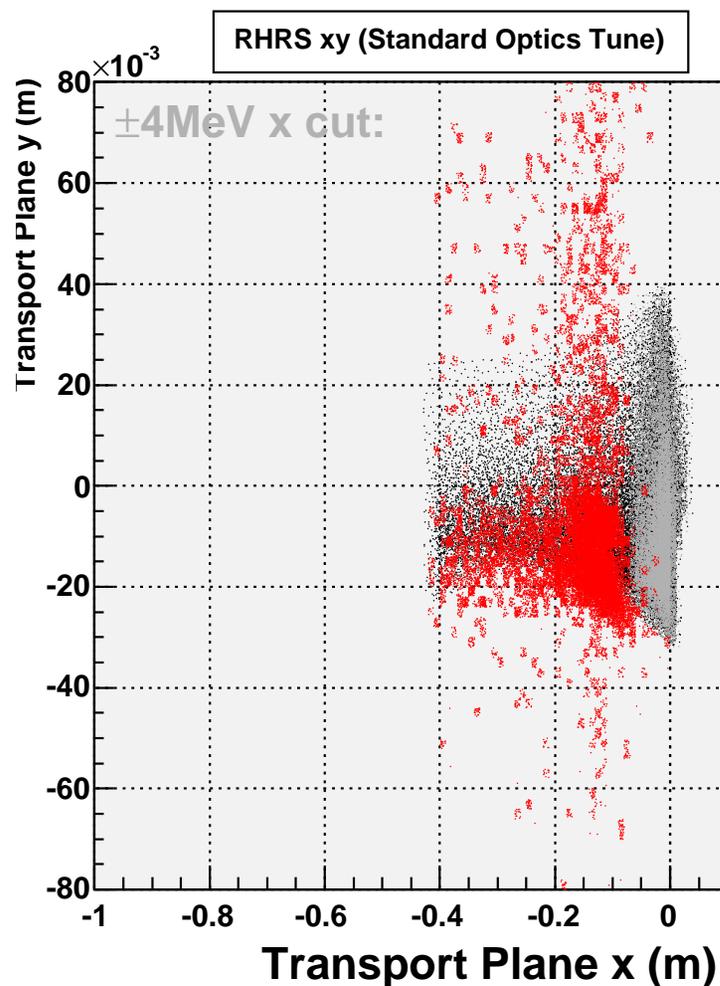
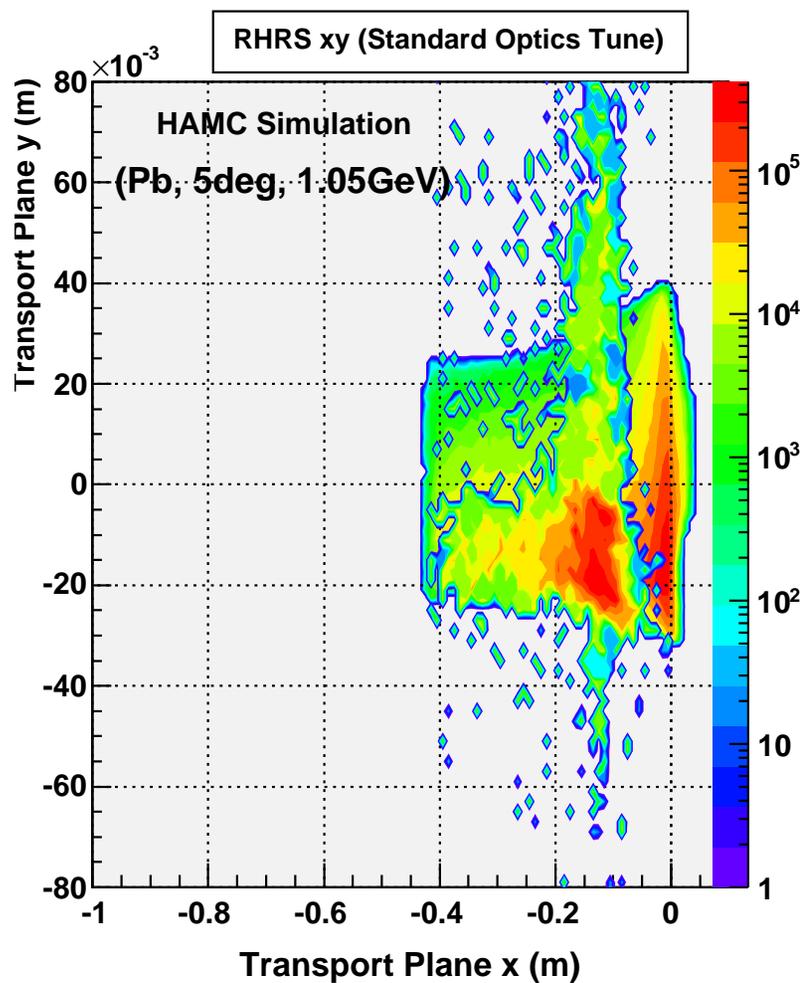


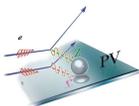
A_T Hole Figure of Merit Study



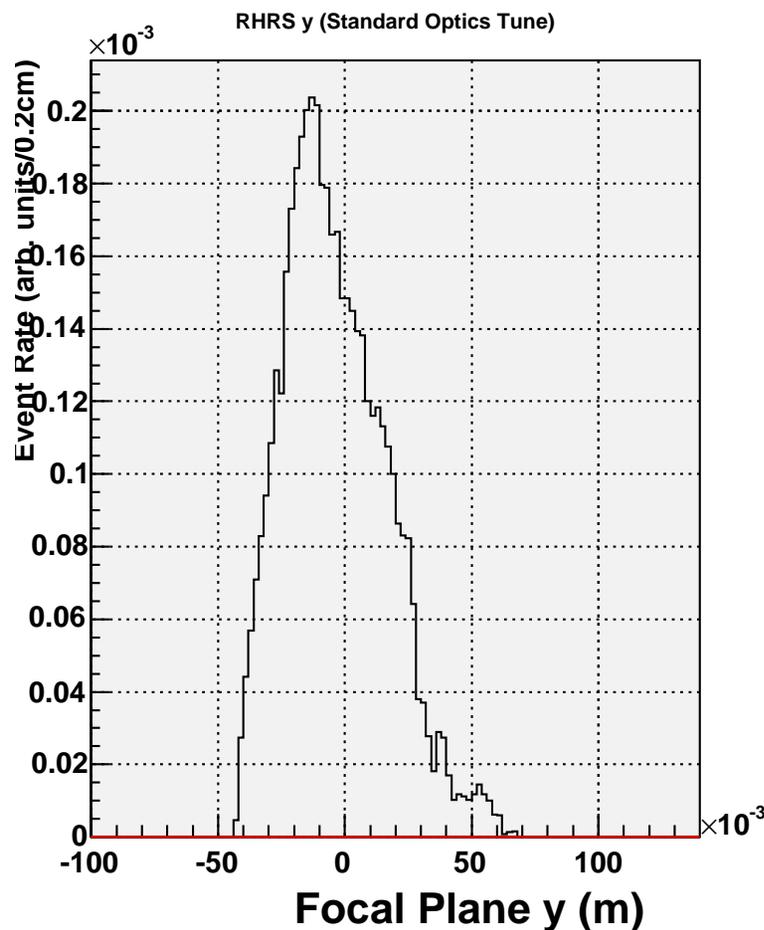
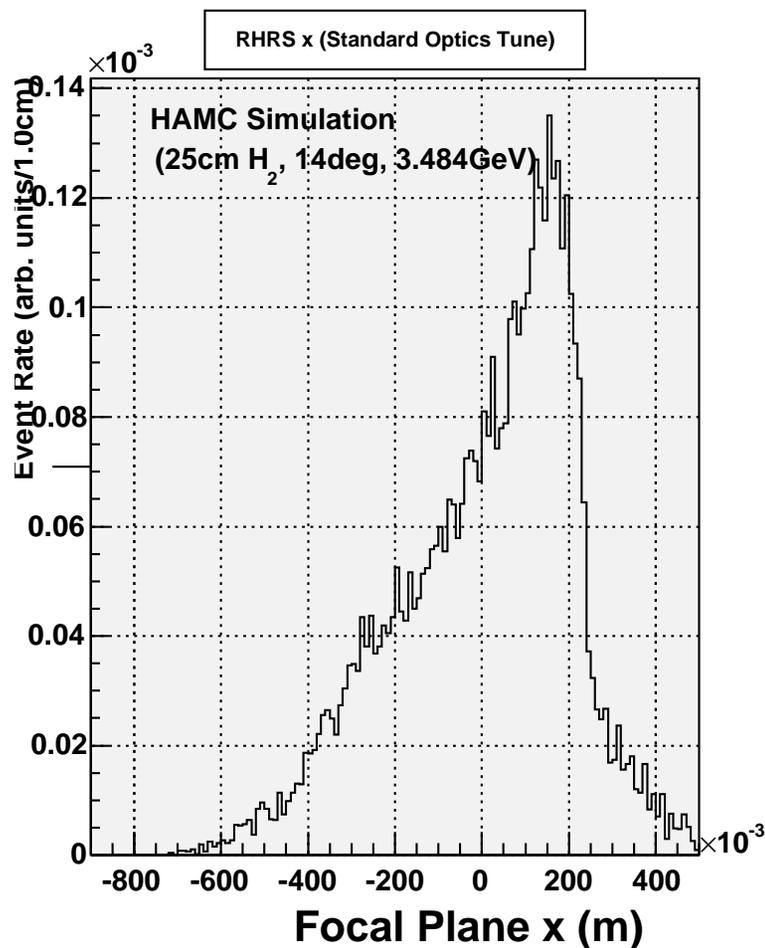


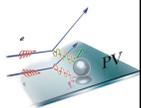
A_T Hole xy Distributions (Standard Tune)





HAPPEX-III Transport Plane Distributions





Test during HAPPEX-III Commissioning

